Curvature Functionals,

Kähler Metrics, &

the Geometry of 4-Manifolds V

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IHP, December 7, 2012

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More precisely, \exists such g with Einstein constant $\lambda \iff$ there is a Kähler form ω such that

$$c_1(M^4, J) = \lambda[\omega].$$

Moreover, this metric is unique, up to isometry, if $\lambda \neq 0$.

Del Pezzo surfaces:

 (M^4, J) for which c_1 is a Kähler class $[\omega]$.

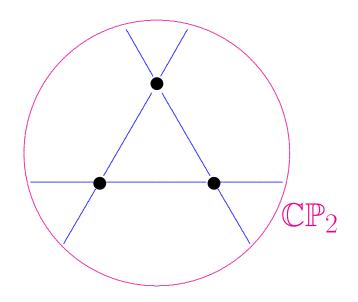
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Blow-up of \mathbb{CP}_2 at k distinct points, $0 \le k \le 8$, in general position, or $\mathbb{CP}_1 \times \mathbb{CP}_1$.



Theorem. Any Del Pezzo surface (M^4, J) admits an Einstein metric h which is conformal to a J-compatible Kähler metric g. In particular, this Einstein metric h is Hermitian with respect to J.

Rough strategy of proof:

Find Kähler metric which minimizes

$$g \mapsto \int_{M} s^2 d\mu_g$$

among all Kähler metrics g.

Here s = scalar curvature.

Note that Kähler class $[\omega]$ of g allowed to vary!

Corresponding problem with $[\omega]$ fixed:

Calabi's extremal Kähler metrics.

So minimize among extremal Kähler metrics.

Minimizer g has s > 0.

Einstein metric is $h = s^{-2}g$.

Theorem A. There is a conformally Kähler, Einstein metric h on $M = \mathbb{CP}_2 \# 2\overline{\mathbb{CP}_2}$ for which the conformally related Kähler g minimizes the functional

$$g \longmapsto \int_{M} s^2 d\mu_g$$

among all Kähler metrics on M. Consequently, h is an absolute minimizer of the functional

$$h \longmapsto \int_{M} |W|_{h}^{2} d\mu_{h}.$$

among all conformally Kähler metrics on M.

Theorem B. This minimizing Kähler metric g on $\mathbb{CP}_2\#2\overline{\mathbb{CP}_2}$ is conformal to an Einstein metric. Moreover, there is a 1-parameter family

$$[0,1)\ni t\longmapsto g_t$$

of extremal Kähler metrics on $\mathbb{CP}_2\#3\overline{\mathbb{CP}_2}$ s.t.

- g_0 is Kähler-Einstein, and such that
- $g_{t_j} \rightarrow g$ in the Gromov-Hausdorff sense for some $t_j \nearrow 1$.

$$0 = 12B = s\mathring{r} + 2\text{Hess}_0(s)$$

 \implies the conformal rescaling $h = s^{-2}g$ is Einstein courtesy of transformation rule

$$\mathring{r}(u^2g) = \mathring{r}(g) + (n-2)u \text{Hess}_0(u^{-1})$$
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Lemma. For any extremal Kähler g on any Del Pezzo M, scalar curvature s > 0 everywhere.

Any Kähler (M^4, g, J) satisfies

$$\frac{1}{32\pi^2} \int s^2 d\mu_g \ge \mathcal{A}([\omega])$$

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$$\mathcal{A}([\omega]) := \frac{(c_1 \cdot [\omega])^2}{[\omega]^2} + \frac{1}{32\pi^2} \|\mathcal{F}_{[\omega]}\|^2$$

where \mathcal{F} is Futaki invariant.

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Lemma. For all $[\omega]$ on any Del Pezzo M,

$$\mathcal{B}([\boldsymbol{\omega}]) < \frac{1}{4}$$

Theorem 1. Let $M = \mathbb{CP}_2 \# 2\overline{\mathbb{CP}}_2$ be the blow-up of \mathbb{CP}_2 at two distinct points,

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Then there is an extremal Kähler metric g on M with Kähler form $\omega \in [\omega]$.

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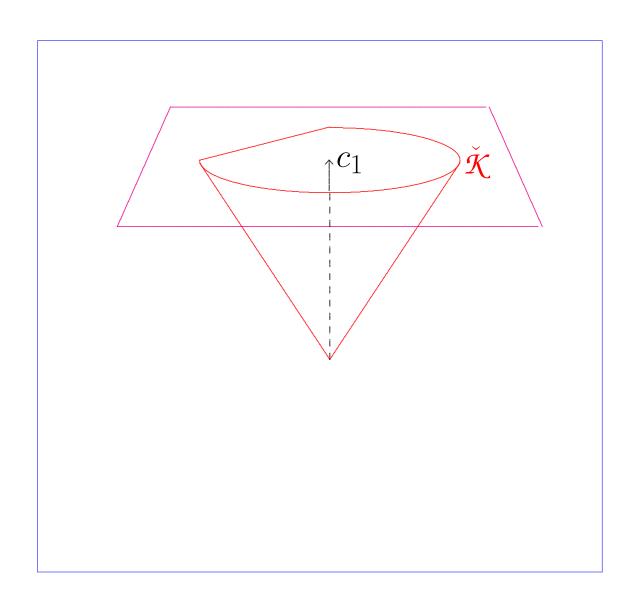
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$$\uparrow c_1$$

$$H^{1,1}(M,\mathbb{R}) = H^2(M,\mathbb{R})$$

$$\uparrow c_1$$

$$\mathcal{T}([\omega]) = \frac{(c_1 \cdot [\omega])^2}{[\omega]^2} \le const$$



$$\check{\mathcal{K}}=\mathcal{K}/\mathbb{R}^+$$

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Generalizes work of Anderson/Tian-Viaclovsky

$$\mathcal{A}([\omega]) < \frac{3}{2}c_1^2(M).$$

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Recall:

$$(2\chi + 3\tau)(\mathbf{M}) = \frac{1}{4\pi^2} \int_{\mathbf{M}} \left(\frac{s^2}{24} + 2|W_+|^2 - \frac{|\mathring{\mathbf{r}}|^2}{2} \right) d\mu_g$$

$$\mathcal{A}([\omega]) < \frac{3}{2}c_1^2(M).$$

Gauss-Bonnet \Longrightarrow

$$Y_{[g]}^2 \ge 64\pi^2 \left(\frac{3}{2}c_1^2 - \mathcal{A}([\omega])\right)$$

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have Sobolev bound on convex cone

$$\mathcal{T}([\omega]) \le \frac{3}{2}c_1^2 - \frac{1}{4}$$

ullet $\exists subsequence which <math>C^{\infty}$ converges modulo diffeomorphims; or

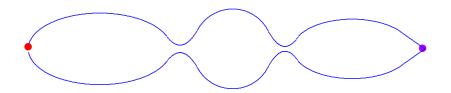
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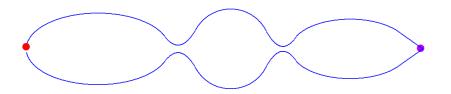
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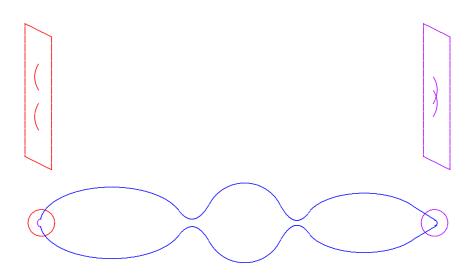
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Then \exists subsequence which Gromov-Hausdorff converges to an extremal Kähler metric on a compact complex 2-orbifold.

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Suggests continuity method...

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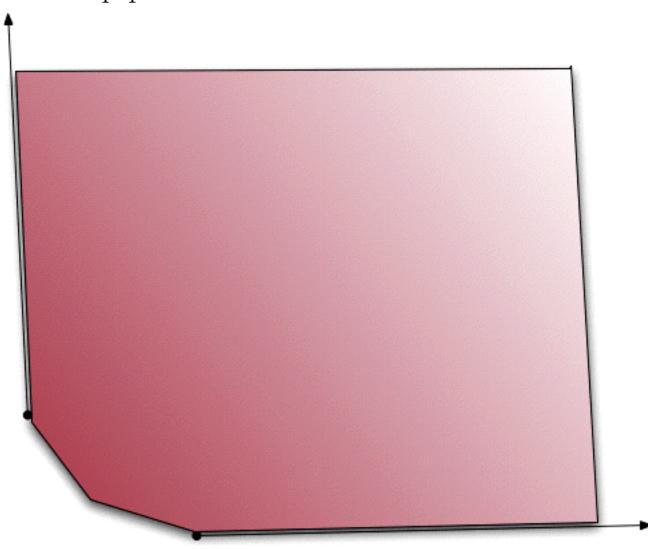
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Difficulty: rule out deepest bubbles.

Lemma. If M is toric, any deepest bubble (X, g_{∞}) must be toric, too, with $H_2(X, \mathbb{R}) \neq 0$ generated by holomorphic \mathbb{CP}_1 's.

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Moment map profile:



• Bubble arises by rescaling region of manifold by scales $\nearrow \infty$.

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Consider line segment

$$[\omega_t] = (1 - t)c_1 + t[\omega]$$

of Kähler classes,

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$$[\omega_{t_i}] \cdot [S] > 0$$

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$$(u_{j}c_{1} + v_{j}\Omega) \cdot [S] > 0, \quad \exists u_{j}, v_{j} > 0$$
$$[S] \cdot [S] = -k < 0$$
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of Kähler classes, and suppose extremal metric exists for $t \in [0, \mathfrak{t})$. If bubbling occurred for $t_j \nearrow \mathfrak{t}$, then, setting $\Omega = [\omega_{\mathfrak{t}}]$, would have $[S] \in H_2(M, \mathbb{Z})$ with

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It follows that bubbling off cannot occur!

Theorem 2. Let $M = \mathbb{CP}_2 \# 3\mathbb{CP}_2$ be the blow-up of \mathbb{CP}_2 at three non-collinear points, and let $[\omega]$ be a Kähler class on M for which

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Also works when approaching boundary of Kähler cone, but can bubble off (-1)-curves.

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of extremal Kähler metrics on $\mathbb{CP}_2\#3\overline{\mathbb{CP}}_2$

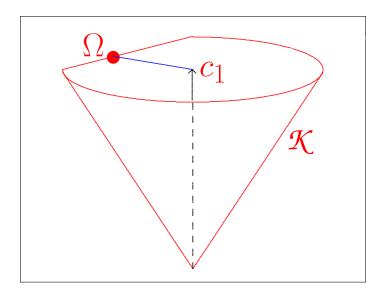
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- $g_{t_j} \rightarrow g$ in the Gromov-Hausdorff sense for some $t_j \nearrow 1$.

Uniform bound $\mathcal{B}([\omega]) < 1/4$ now implies that

$$A = T + B$$

has minimizer $[\omega]$ represented by conformally Einstein Kähler metric.

$$h(J\cdot, J\cdot) = h.$$

Then either

- \bullet (M, J, h) is Kähler-Einstein; or
- $M \approx \mathbb{CP}_2 \# \overline{\mathbb{CP}}_2$, and h is a constant times the Page metric; or
- $M \approx \mathbb{CP}_2 \# 2\mathbb{CP}_2$ and h is a constant times the Chen-LeBrun-Weber metric.

Uniqueness:

Theorem C. Let (M^4, J) be a compact complex surface, and suppose that h is an Einstein metric on M which is Hermitian with respect to J:

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- \bullet (M, J, h) is Kähler-Einstein; or
- $M \approx \mathbb{CP}_2 \# \mathbb{CP}_2$, and h is a constant times the Page metric; or
- $M \approx \mathbb{CP}_2 \# 2\overline{\mathbb{CP}}_2$ and h is a constant times the Chen-LeBrun-Weber metric.

Exceptional cases: \mathbb{CP}_2 blown up at 1 or 2 points.

More precisely, there is a Hermitian, Einstein metric h with Einstein constant $\lambda \iff (M, J)$ carries a Kähler class $[\omega]$ such that

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Warning: when h is non-Kähler, its relation to ω is surprisingly complicated!

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In other words,

$$h = fg$$

 \exists Kähler metric g, smooth function $f: M \to \mathbb{R}^+$.

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Similarly for $S^{2n+1} \times S^{2m+1}$.

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May normalize so that either $f = s^{-2}$ or f = 1.

Calabi:

Extremal Kähler metrics = critical points of

$$g \mapsto \int_{M} s^2 d\mu_g$$

where $g = g_{\omega}$ for J and $[\omega] \in H^2(M, \mathbb{R})$ fixed.

Euler-Lagrange equations \iff

 $\nabla^{1,0}s$ is a holomorphic vector field.

Donaldson/Mabuchi/Chen-Tian: unique in Kähler class, modulo bihomorphisms. Riemann curvature of g

$$\mathcal{R}: \Lambda^2 \to \Lambda^2$$

splits into 4 irreducible pieces:

$$\Lambda^{+*} \qquad \Lambda^{-*}$$

$$\Lambda^{+} \qquad W_{+} + \frac{s}{12} \qquad \mathring{r}$$

$$\Lambda^{-} \qquad \mathring{r} \qquad W_{-} + \frac{s}{12}$$

where

s = scalar curvature

 \mathring{r} = trace-free Ricci curvature

 $W_{+} = \text{self-dual Weyl curvature } (conformally invariant)$

 W_{-} = anti-self-dual Weyl curvature

Kähler case:

$$\Lambda^{1,1} = \mathbb{R}\omega \oplus \Lambda^{-}$$

$$\Lambda^{+} = \mathbb{R}\omega \oplus \Re e(\Lambda^{2,0})$$

$$\nabla J = 0 \Longrightarrow \mathcal{R} \in \operatorname{End}(\Lambda^{1,1}) \Longrightarrow$$

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Notice that W_{+} has a repeated eigenvalue.

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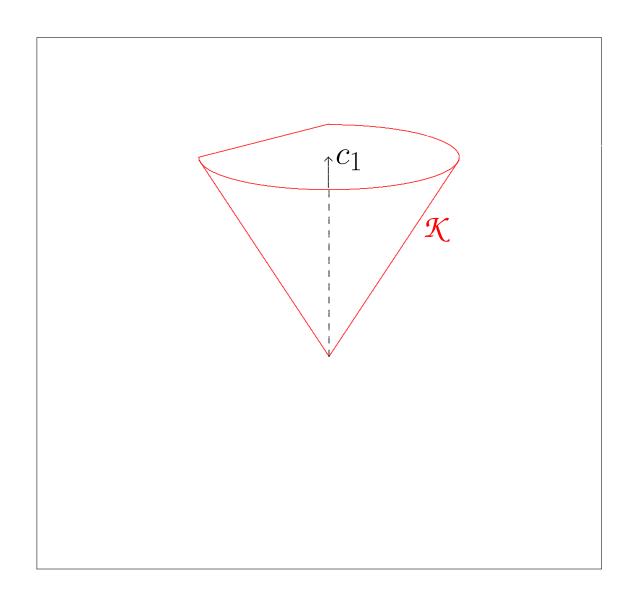
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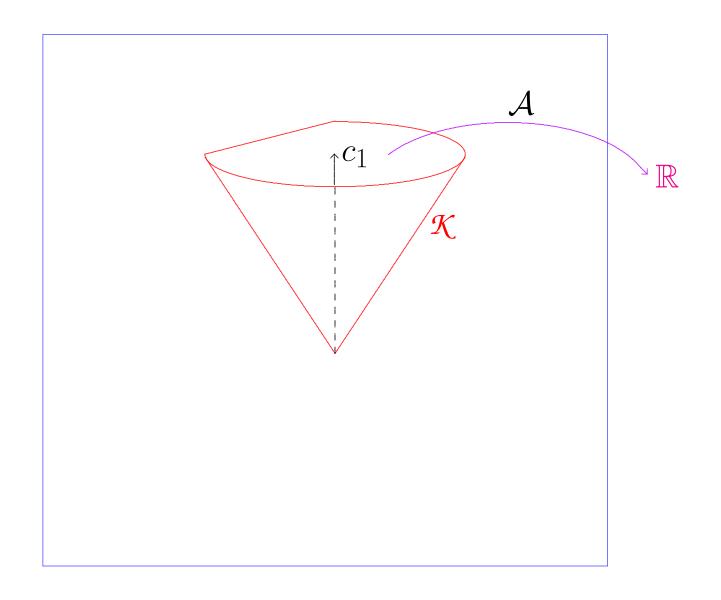
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- $[\omega]$ is a critical point of $\mathcal{A}: \mathcal{K} \to \mathbb{R}$.

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$$0 = 12s^{-1}B = \mathring{r} + 2s^{-1}\text{Hess}_0(s)$$

$$\rho + 2i\partial\bar{\partial}\log s > 0.$$

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Theorem. Let (M^4, J) be a Del Pezzo surface. Then, up to automorphisms and rescaling, there is a unique Bach-flat Kähler metric g on M. This metric is characterized by the fact that it minimizes the Calabi functional

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Hermitian, Einstein metric then given by

$$h = s^{-2}g$$

and uniqueness Theorem A follows.

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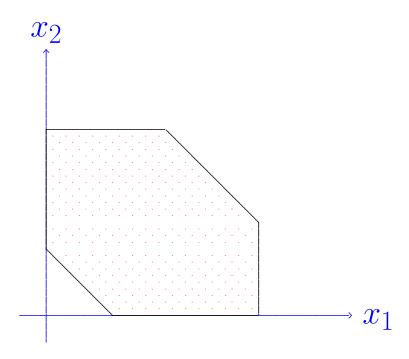
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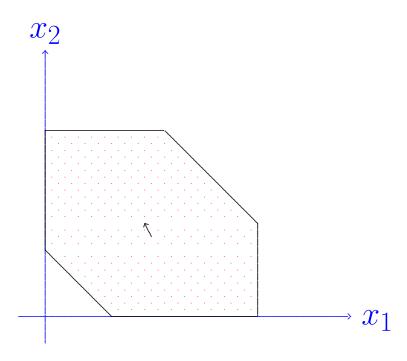
Only three cases are non-trivial:

$$\mathbb{CP}_2 \# k \overline{\mathbb{CP}}_2, \quad k = 1, 2, 3.$$

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$$\mathcal{A}([\boldsymbol{\omega}]) = \frac{|\partial P|^2}{2} \left(\frac{1}{|P|} + \vec{\mathfrak{D}} \cdot \Pi^{-1} \vec{\mathfrak{D}} \right)$$

To prove Theorem, show that

$$\mathcal{A}: \check{\mathcal{K}}
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has unique critical point for relevant M.

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 ${\cal A}$ is explicit rational function —

 $3 \left[3 + 28\gamma + 96\gamma^2 + 168\gamma^3 + 164\gamma^4 + 80\gamma^5 + 16\gamma^6 + 16\beta^6 (1+\gamma)^4 + 16\alpha^6 (1+\beta+\gamma)^4 + 16\beta^5 (5 + 24\gamma + 43\gamma^2 + 37\gamma^3 + 15\gamma^4 + 2\gamma^5) + 4\beta^4 (41 + 228\gamma + 478\gamma^2 + 496\gamma^3 + 263\gamma^4 + 16\gamma^4 + 16\gamma$ $60\gamma^5 + 4\gamma^6) + 8\beta^3(21 + 135\gamma + 326\gamma^2 + 392\gamma^3 + 248\gamma^4 + 74\gamma^5 + 8\gamma^6) + 4\beta(7 + 58\gamma + 176\gamma^2 + 270\gamma^3 + 228\gamma^4 + 96\gamma^5 + 16\gamma^6) + 4\beta^2(24 + 176\gamma + 479\gamma^2 + 652\gamma^3 + 478\gamma^4 + 176\gamma^2 + 176\gamma$ $172\gamma^{5} + 24\gamma^{6}) + 16\alpha^{5}(5 + 2\beta^{5} + 24\gamma + 43\gamma^{2} + 37\gamma^{3} + 15\gamma^{4} + 2\gamma^{5} + \beta^{4}(15 + 14\gamma) + \beta^{3}(37 + 70\gamma + 30\gamma^{2}) + \beta^{2}(43 + 123\gamma + 108\gamma^{2} + 30\gamma^{3}) + \beta(24 + 92\gamma + 123\gamma^{2} + 70\gamma^{3} + 123\gamma^{2} +$ $14\gamma^{4})) + 4\alpha^{4}(41 + 4\beta^{6} + 228\gamma + 478\gamma^{2} + 496\gamma^{3} + 263\gamma^{4} + 60\gamma^{5} + 4\gamma^{6} + \beta^{5}(60 + 56\gamma) + \beta^{4}(263 + 476\gamma + 196\gamma^{2}) + 8\beta^{3}(62 + 169\gamma + 139\gamma^{2} + 35\gamma^{3}) + 2\beta^{2}(239 + 876\gamma + 1089\gamma^{2} + 108\gamma^{2}) + 3\beta^{2}(239 + 876\gamma + 108\gamma^{2} + 108\gamma^{2} + 108\gamma^{2}) + 3\beta^{2}(239 + 876\gamma + 108\gamma^{2} + 108\gamma^{2} + 108\gamma^{2}) + 3\beta^{2}(239 + 876\gamma + 108\gamma^{2} +$ $556\gamma^{3} + 98\gamma^{4}) + 4\beta(57 + 263\gamma + 438\gamma^{2} + 338\gamma^{3} + 119\gamma^{4} + 14\gamma^{5})) + 8\alpha^{3}(21 + 135\gamma + 326\gamma^{2} + 392\gamma^{3} + 248\gamma^{4} + 74\gamma^{5} + 8\gamma^{6} + 8\beta^{6}(1 + \gamma) + 2\beta^{5}(37 + 70\gamma + 30\gamma^{2}) + 4\beta^{4}(62 + 32\gamma^{2} + 338\gamma^{2} + 119\gamma^{4} + 14\gamma^{5})) + 8\alpha^{3}(21 + 135\gamma + 326\gamma^{2} + 392\gamma^{3} + 248\gamma^{4} + 74\gamma^{5} + 8\gamma^{6} + 8\beta^{6}(1 + \gamma) + 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526\gamma^{4} + 184\gamma^{5} + 123\gamma^{6}) + 3\beta^{2} (135 + 736\gamma^{2} + 1412\gamma^{3} + 676\gamma^{4} + 140\gamma^{5} + 8\gamma^{6}) + 4\beta^{2} (147 + 123\gamma^{6} + 1412\gamma^{6} + 123\gamma^{6} + 123\gamma^{6}$ $24\gamma^{6})) + 4\alpha^{2}(24 + 176\gamma + 479\gamma^{2} + 652\gamma^{3} + 478\gamma^{4} + 172\gamma^{5} + 24\gamma^{6} + 24\beta^{6}(1 + \gamma)^{2} + 4\beta^{5}(43 + 123\gamma + 108\gamma^{2} + 30\gamma^{3}) + 2\beta^{4}(239 + 876\gamma + 1089\gamma^{2} + 556\gamma^{3} + 98\gamma^{4}) + 4\beta^{3}(163 + 123\gamma + 108\gamma^{2} + 30\gamma^{3}) + 2\beta^{4}(239 + 876\gamma + 1089\gamma^{2} + 556\gamma^{3} + 98\gamma^{4}) + 4\beta^{3}(163 + 123\gamma + 108\gamma^{2} + 30\gamma^{3}) + 2\beta^{4}(239 + 876\gamma + 1089\gamma^{2} + 556\gamma^{3} + 98\gamma^{4}) + 4\beta^{3}(163 + 123\gamma + 108\gamma^{2} + 30\gamma^{3}) + 2\beta^{4}(239 + 876\gamma + 1089\gamma^{2} + 556\gamma^{3} + 98\gamma^{4}) + 4\beta^{3}(163 + 123\gamma + 108\gamma^{2} + 30\gamma^{3}) + 2\beta^{4}(239 + 876\gamma + 1089\gamma^{2} + 556\gamma^{3} + 98\gamma^{4}) + 4\beta^{3}(163 + 123\gamma + 108\gamma^{2} + 108\gamma^{2}$ $735\gamma + 1179\gamma^2 + 856\gamma^3 + 278\gamma^4 + 30\gamma^5) + 4\beta(44 + 278\gamma + 645\gamma^2 + 735\gamma^3 + 438\gamma^4 + 123\gamma^5 + 12\gamma^6) + \beta^2(479 + 2580\gamma + 5058\gamma^2 + 4716\gamma^3 + 2178\gamma^4 + 432\gamma^5 + 24\gamma^6)) \Big] / \\$ $\left[1+10\gamma+36\gamma^{2}+64\gamma^{3}+60\gamma^{4}+24\gamma^{5}+24\beta^{5}(1+\gamma)^{5}+24\alpha^{5}(1+\beta+\gamma)^{5}+12\beta^{4}(1+\gamma)^{2}(5+20\gamma+23\gamma^{2}+10\gamma^{3})+16\beta^{3}(4+28\gamma+72\gamma^{2}+90\gamma^{3}+57\gamma^{4}+15\gamma^{5})+12\beta^{4}(1+\gamma)^{2}(5+20\gamma+23\gamma^{2}+10\gamma^{3})+16\beta^{3}(4+28\gamma+72\gamma^{2}+90\gamma^{3}+57\gamma^{4}+15\gamma^{5})+12\beta^{4}(1+\gamma)^{2}(5+20\gamma+23\gamma^{2}+10\gamma^{3})+16\beta^{3}(4+28\gamma+72\gamma^{2}+90\gamma^{3}+57\gamma^{4}+15\gamma^{5})+12\beta^{4}(1+\gamma)^{2}(5+20\gamma+23\gamma^{2}+10\gamma^{3})+16\beta^{3}(4+28\gamma+72\gamma^{2}+90\gamma^{3}+57\gamma^{4}+15\gamma^{5})+12\beta^{4}(1+\gamma)^{2}(5+20\gamma+23\gamma^{2}+10\gamma^{3})+16\beta^{3}(4+28\gamma+72\gamma^{2}+90\gamma^{3}+57\gamma^{4}+15\gamma^{5})+12\beta^{4}(1+\gamma)^{2}(5+20\gamma+23\gamma^{2}+10\gamma^{3})+16\beta^{3}(4+28\gamma+72\gamma^{2}+90\gamma^{3}+57\gamma^{4}+15\gamma^{5})+16\beta^{3}(1+\gamma)^{2$ $12\beta^{2}(3 + 24\gamma + 69\gamma^{2} + 96\gamma^{3} + 68\gamma^{4} + 20\gamma^{5}) + 2\beta(5 + 45\gamma + 144\gamma^{2} + 224\gamma^{3} + 180\gamma^{4} + 60\gamma^{5}) + 12\alpha^{4}(1 + \beta + \gamma)^{2}(5 + 20\gamma + 23\gamma^{2} + 10\gamma^{3} + 10\beta^{3}(1 + \gamma) + \beta^{2}(23 + 46\gamma + 10\gamma^{2} + 1$ $16\gamma^{2}) + 2\beta(10 + 30\gamma + 23\gamma^{2} + 5\gamma^{3})) + 16\alpha^{3}(4 + 28\gamma + 72\gamma^{2} + 90\gamma^{3} + 57\gamma^{4} + 15\gamma^{5} + 15\beta^{5}(1 + \gamma)^{2} + 3\beta^{4}(19 + 57\gamma + 50\gamma^{2} + 13\gamma^{3}) + 3\beta^{3}(30 + 120\gamma + 155\gamma^{2} + 78\gamma^{3} + 15\gamma^{2} +$ $13\gamma^{4}) + 3\beta^{2}(24 + 120\gamma + 206\gamma^{2} + 155\gamma^{3} + 50\gamma^{4} + 5\gamma^{5}) + \beta(28 + 168\gamma + 360\gamma^{2} + 360\gamma^{3} + 171\gamma^{4} + 30\gamma^{5})) + 12\alpha^{2}(3 + 24\gamma + 69\gamma^{2} + 96\gamma^{3} + 68\gamma^{4} + 20\gamma^{5} + 20\beta^{5}(1 + \gamma)^{3} + 20\gamma^{5}) + 3\beta^{2}(24 + 120\gamma + 206\gamma^{2} + 155\gamma^{3} + 50\gamma^{4} + 5\gamma^{5}) + \beta(28 + 168\gamma + 360\gamma^{2} + 360\gamma^{3} + 171\gamma^{4} + 30\gamma^{5})) + 12\alpha^{2}(3 + 24\gamma + 69\gamma^{2} + 96\gamma^{3} + 68\gamma^{4} + 20\gamma^{5} + 20\beta^{5}(1 + \gamma)^{3} + 20\gamma^{5}(1 + \gamma)^{3}) + 3\beta^{2}(24 + 120\gamma + 206\gamma^{2} + 155\gamma^{3} + 50\gamma^{4} + 5\gamma^{5}) + \beta(28 + 168\gamma + 360\gamma^{2} + 360\gamma^{3} + 171\gamma^{4} + 30\gamma^{5})) + 12\alpha^{2}(3 + 24\gamma + 69\gamma^{2} + 96\gamma^{3} + 68\gamma^{4} + 20\gamma^{5} + 20\beta^{5}(1 + \gamma)^{3}) + 3\beta^{2}(24 + 120\gamma + 206\gamma^{2} + 155\gamma^{3} + 50\gamma^{4} + 5\gamma^{5}) + \beta(28 + 168\gamma + 360\gamma^{2} + 360\gamma^{2} + 360\gamma^{2} + 360\gamma^{2} + 36\gamma^{4} + 360\gamma^{2} + 36\gamma^{4} + 36$ $\beta^{4} (68 + 272\gamma + 366\gamma^{2} + 200\gamma^{3} + 36\gamma^{4}) + 4\beta^{3} (24 + 120\gamma + 206\gamma^{2} + 155\gamma^{3} + 50\gamma^{4} + 5\gamma^{5}) + 2\beta (12 + 84\gamma + 207\gamma^{2} + 240\gamma^{3} + 136\gamma^{4} + 30\gamma^{5}) + \beta^{2} (69 + 414\gamma + 864\gamma^{2} + 120\gamma^{2} + 120\gamma$ $824\gamma^{3} + 366\gamma^{4} + 60\gamma^{5})) + 2\alpha(5 + 45\gamma + 144\gamma^{2} + 224\gamma^{3} + 180\gamma^{4} + 60\gamma^{5} + 60\beta^{5}(1 + \gamma)^{4} + 12\beta^{4}(15 + 75\gamma + 136\gamma^{2} + 114\gamma^{3} + 43\gamma^{4} + 5\gamma^{5}) + 12\beta^{2}(12 + 84\gamma + 207\gamma^{2} + 12\beta^{2}) + 12\beta^{2}(12 + 84\gamma^{2} + 12\beta^{2}) + 12\beta^{2}(12$ $240\gamma^{3} + 136\gamma^{4} + 30\gamma^{5}) + 8\beta^{3}(28 + 168\gamma + 360\gamma^{2} + 360\gamma^{3} + 171\gamma^{4} + 30\gamma^{5}) + 3\beta(15 + 120\gamma + 336\gamma^{2} + 448\gamma^{3} + 300\gamma^{4} + 80\gamma^{5})) \Big]$

$$\mathcal{A}: \check{\mathcal{K}}
ightarrow \mathbb{R}$$

has unique critical point for relevant M.

Here
$$\check{\mathcal{K}} = \mathcal{K}/\mathbb{R}^+$$
.

A is explicit rational function — but quite complicated!

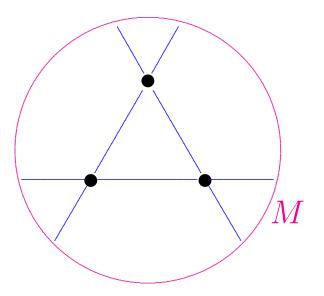
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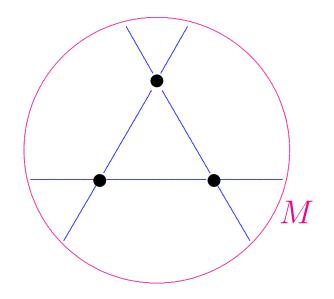
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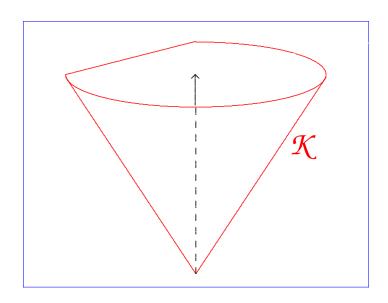
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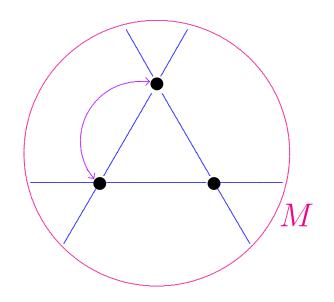
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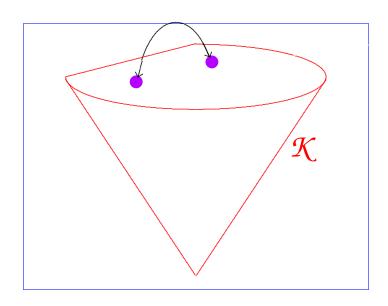
Proof proceeds by showing critical point invariant under certain discrete automorphisms of M.

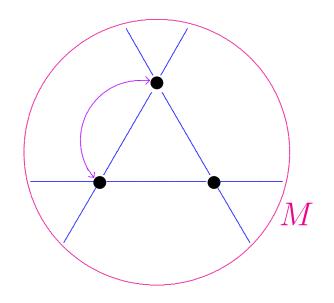


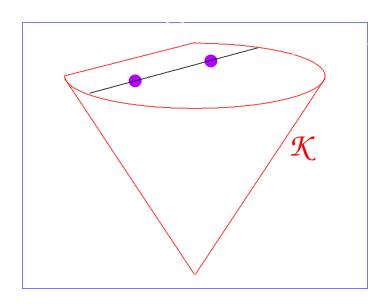


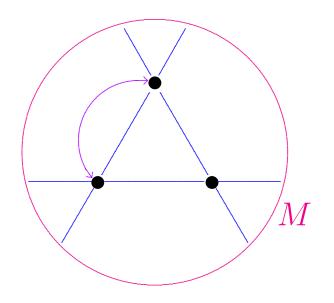


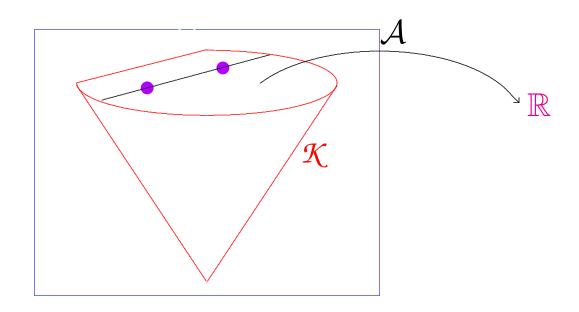












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Done by showing \mathcal{A} convex on appropriate lines.

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Final step then just calculus in one variable...

Proposition. Modulo rescalings and biholomorphisms, there is exactly one conformally Kähler, Einstein metric h on $M = \mathbb{CP}_2 \# \overline{\mathbb{CP}}_2$. This metric coincides with the Page metric.

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Proposition. Modulo rescalings and biholomorphisms, there is only one conformally Kähler, Einstein metric h on $M = \mathbb{CP}_2 \# 3\overline{\mathbb{CP}_2}$. This metric is actually Kähler-Einstein, and is exactly the metric discovered by Siu.

 $M pprox \left\{ egin{align*} \mathbb{CP}_2 \# k \overline{\mathbb{CP}}_2, \\ M pprox \left\{ egin{align*} \mathcal{CP}_2 \# k \overline{\mathbb{CP}}_2, \\ \mathcal{CP}_2 \# k \overline{\mathbb{CP}_2}, \\ \mathcal{CP}_2 \# k \overline{\mathbb{CP}}_2, \\ \mathcal{CP}_2 \# k \overline{\mathbb{$

 $M \approx \begin{cases} \mathbb{CP}_2 \# k \overline{\mathbb{CP}_2}, & 0 \le k \le 8, \end{cases}$

$$pact\ oriented\ 4\text{-}manifold\ which\ adm}$$

$$grable\ complex\ structure\ J.\ Then\ M$$

$$an\ Einstein\ metric\ g\ with\ \lambda\geq 0\ if\ a$$

$$\left\{\begin{array}{c} \mathbb{CP}_2\#k\overline{\mathbb{CP}}_2,\quad 0\leq k\leq 8,\\ S^2\times S^2,\end{array}\right.$$

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Theorem E. Suppose that
$$M$$
 is a samp pact oriented 4-manifold which adm grable complex structure J . Then M an Einstein metric g with $\lambda \geq 0$ if a $\mathbb{CP}_2 \# k \overline{\mathbb{CP}}_2$, $0 \leq k \leq 8$, $S^2 \times S^2$, $K3$, $K3/\mathbb{Z}_2$, T^4 ,

an Einstein metric
$$g$$
 with $\lambda \geq 0$ if and
$$\begin{cases} \mathbb{CP}_2 \# k \overline{\mathbb{CP}}_2, & 0 \leq k \leq 8, \\ S^2 \times S^2, \\ K3, \\ K3/\mathbb{Z}_2, \\ T^4, \\ T^4/\mathbb{Z}_2, T^4/\mathbb{Z}_3, T^4/\mathbb{Z}_4, T^4/\mathbb{Z}_6, \end{cases}$$

```
\begin{cases}
\mathbb{CP}_{2} \# k \overline{\mathbb{CP}}_{2}, & 0 \leq k \leq 8, \\
S^{2} \times S^{2}, \\
K3, \\
K3/\mathbb{Z}_{2}, \\
T^{4}, \\
T^{4}/\mathbb{Z}_{2}, T^{4}/\mathbb{Z}_{3}, T^{4}/\mathbb{Z}_{4}, T^{4}/\mathbb{Z}_{6}, \\
T^{4}/(\mathbb{Z}_{2} \oplus \mathbb{Z}_{2}), T^{4}/(\mathbb{Z}_{3} \oplus \mathbb{Z}_{3}), or T^{4}/(\mathbb{Z}_{2} \oplus \mathbb{Z}_{4}).
\end{cases}
```

Theorem. Let M be the 4-manifold underlying a compact complex surface. Suppose that M an Einstein metric g.

• M is a surface of general type;

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- M is not too non-minimal

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in the sense that it is obtained from its minimal model X by blowing up at $k < c_1^2(X)/3$ points.

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Same conclusion holds in symplectic case.

Question. Are there any non-minimal complex surfaces M of general type which actually admit Einstein metrics?

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If so, quite different from Kähler-Einstein metrics!

Question. Are the constructed Einstein metrics on rational surfaces the only ones? For example, are there non-standard Einstein metrics on $S^2 \times S^2$ or on \mathbb{CP}_2 ?

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Question. When a 4-manifold M admits a Kähler-Einstein metric g with s > 0, Gursky has shown that, among all metrics with s > 0, it is a minimizes $\int |W|^2 d\mu$, and that the only minimizers are other K-E metrics. Are the Page and C-L-W metrics similarly minimizers for this problem? What happens if we consider metrics which do not have s > 0?

End, Part V