

Gravitational Instantons,
Weyl Curvature, &
Conformally Kähler Geometry

Claude LeBrun
Stony Brook University

Differential Geometry and
Geometric Analysis Seminar
Princeton University
November 15, 2023

Joint work with

Joint work with

Olivier Biquard

Joint work with

Olivier Biquard
Sorbonne Université

Joint work with

Olivier Biquard
Sorbonne Université

and

Joint work with

Olivier Biquard
Sorbonne Université

and

Paul Gauduchon

Joint work with

Olivier Biquard
Sorbonne Université

and

Paul Gauduchon
École Polytechnique

Joint work with

Olivier Biquard
Sorbonne Université

and

Paul Gauduchon
École Polytechnique

e-print:

[arXiv:2310.14387](https://arxiv.org/abs/2310.14387) [math.DG]

Definition. *A gravitational instanton is a*

Definition. *A gravitational instanton is a complete,*

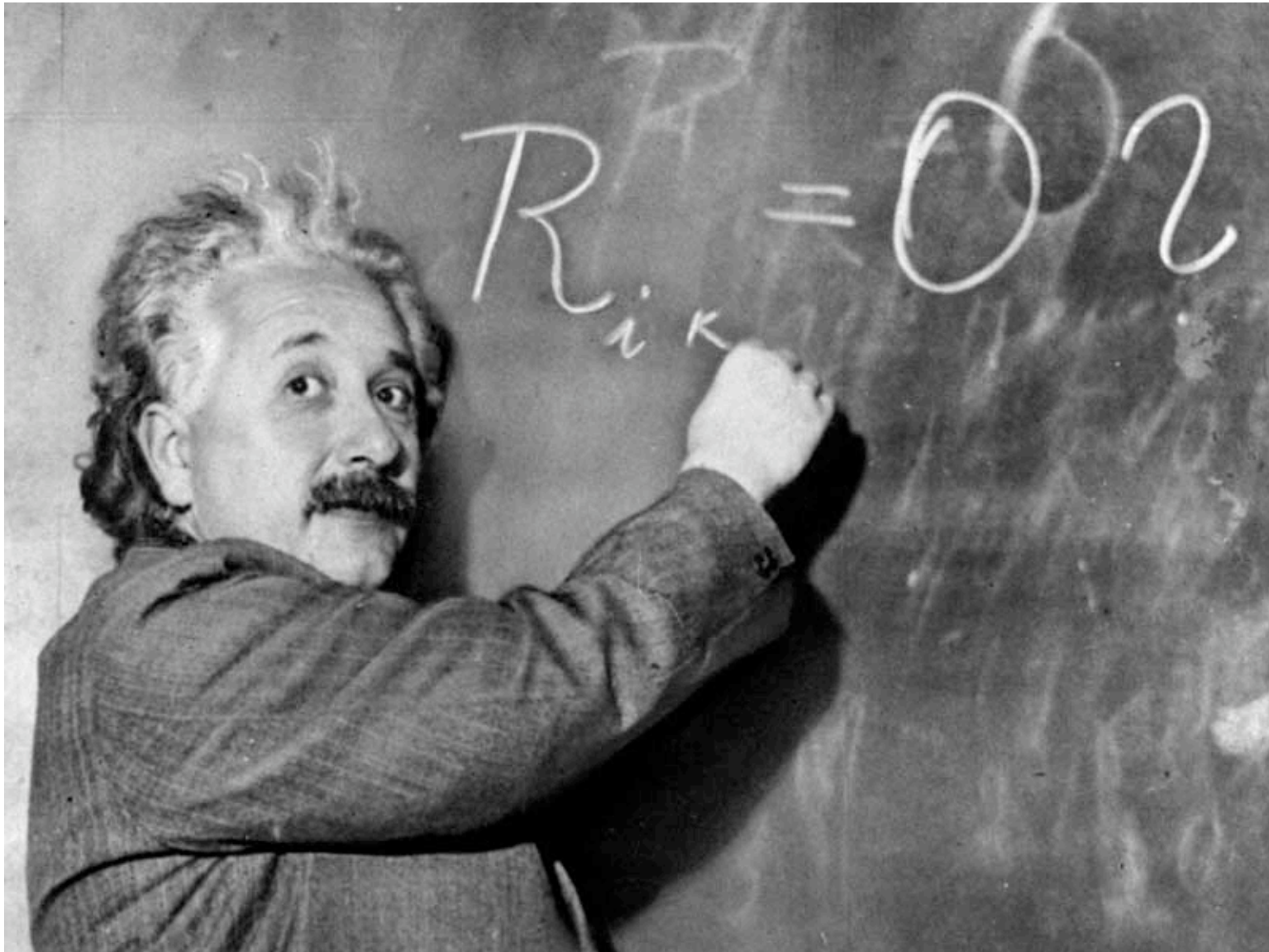
Definition. *A gravitational instanton is a complete, non-compact,*

Definition. *A gravitational instanton is a complete, non-compact, non-flat,*

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat*

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*



Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Terminology due to Gibbons & Hawking, late '70s

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Terminology due to Gibbons & Hawking, late '70s



Key examples:

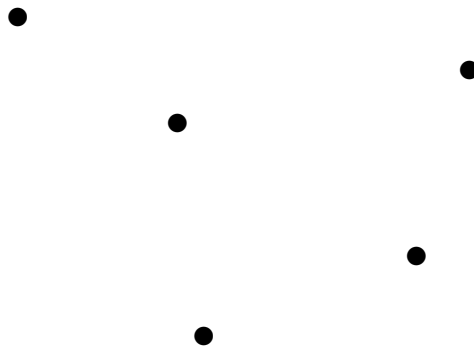
Key examples:

Discovered by Gibbons & Hawking, 1979.

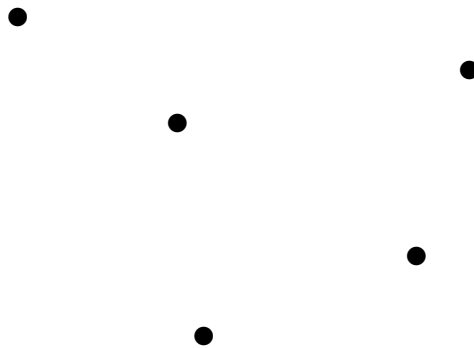
Key examples:

Discovered by Gibbons & Hawking, 1979.

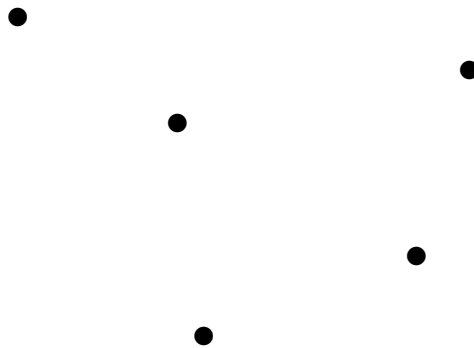
Data: ℓ points in \mathbb{R}^3 and a constant $\kappa^2 \geq 0$.



Data: ℓ points in \mathbb{R}^3 and a constant $\kappa^2 \geq 0$.

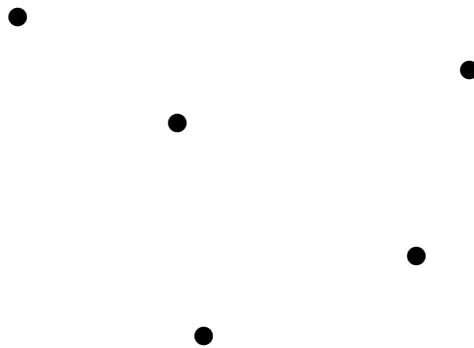


Data: ℓ points in \mathbb{R}^3 and κ^2



Data: ℓ points in \mathbb{R}^3 and $\kappa^2 \implies V$ with $\Delta V = 0$

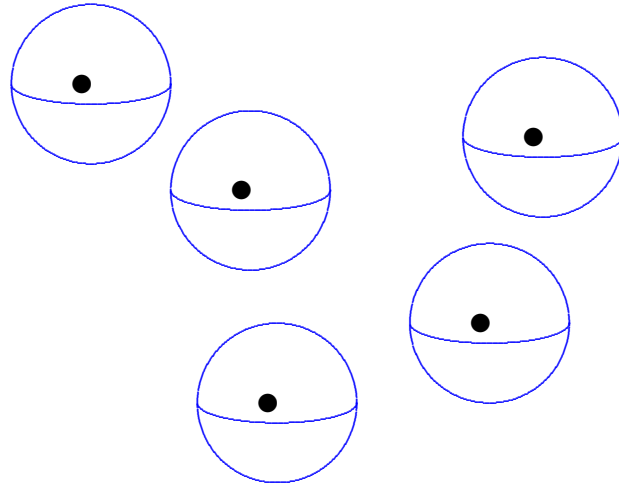
$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2^{\rho_j}}$$



Data: ℓ points in \mathbb{R}^3 and $\kappa^2 \implies V$ with $\Delta V = 0$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\rho_j}$$

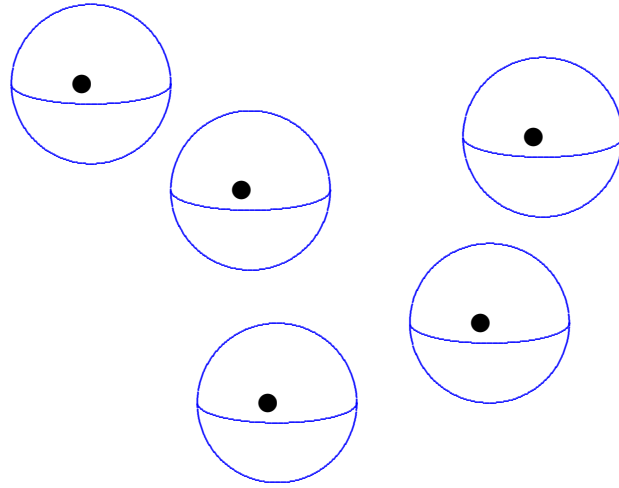
$F = \star dV$ closed 2-form,



Data: ℓ points in \mathbb{R}^3 and $\kappa^2 \implies V$ with $\Delta V = 0$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\rho_j}$$

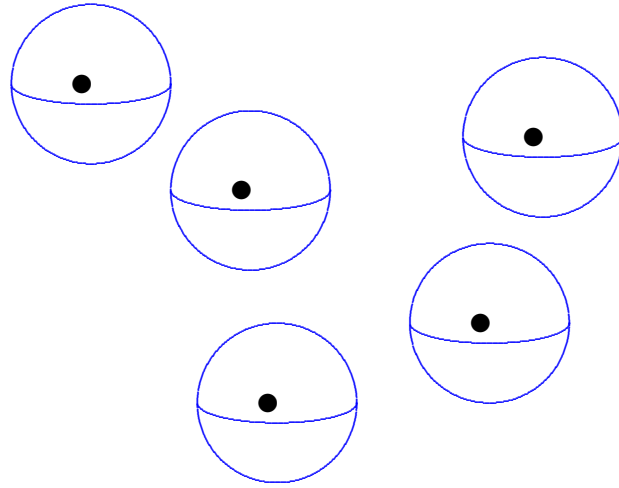
$F = \star dV$ closed 2-form, $[\frac{1}{2\pi}F] \in H^2(\mathbb{R}^3 - \{p_j\}, \mathbb{Z})$.



Data: ℓ points in \mathbb{R}^3 and $\kappa^2 \implies V$ with $\Delta V = 0$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\rho_j}$$

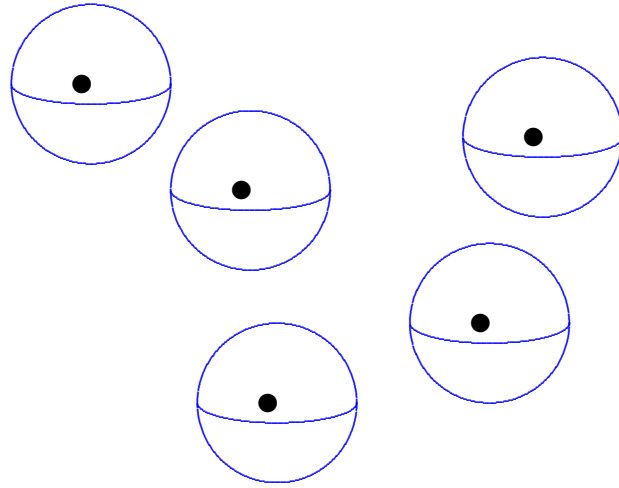
$F = \star dV$ curvature θ on $P \rightarrow \mathbb{R}^3 - \{\text{pts}\}$.



Data: ℓ points in \mathbb{R}^3 . $\implies V$ with $\Delta V = 0$

$$g = Vh + V^{-1}\theta^2$$

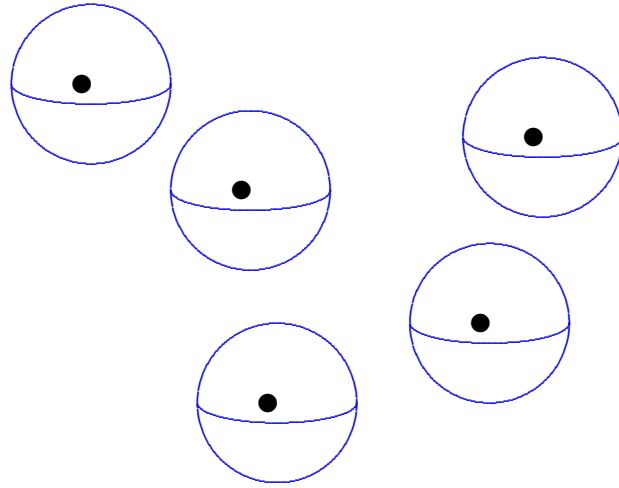
$F = \star dV$ curvature θ on $P \rightarrow \mathbb{R}^3 - \{\text{pts}\}$.



Data: ℓ points in \mathbb{R}^3 . $\implies V$ with $\Delta V = 0$

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

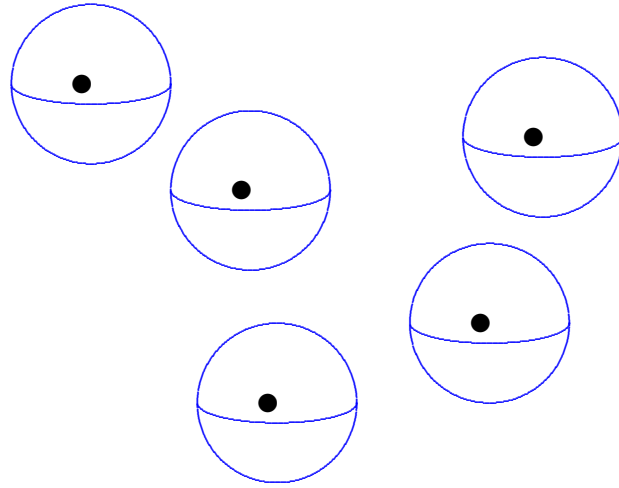
$F = \star dV$ curvature θ on $P \rightarrow \mathbb{R}^3 - \{\text{pts}\}$.



Data: ℓ points in \mathbb{R}^3 . $\implies V$ with $\Delta V = 0$

$$g = Vh + V^{-1}\theta^2$$

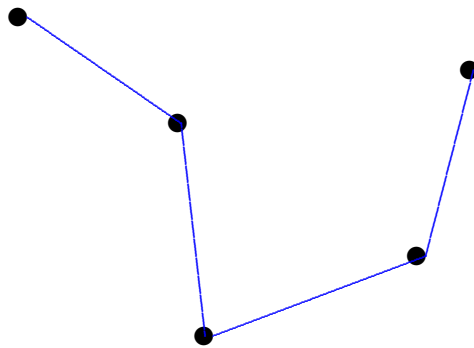
$F = \star dV$ curvature θ on $P \rightarrow \mathbb{R}^3 - \{\text{pts}\}$.



Data: ℓ points in \mathbb{R}^3 . $\implies V$ with $\Delta V = 0$

$$g = Vh + V^{-1}\theta^2$$

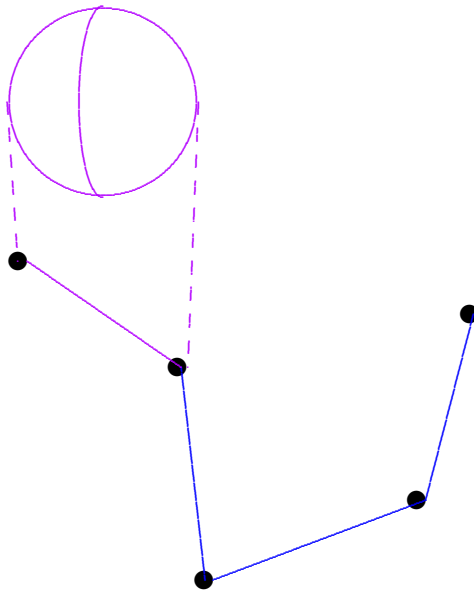
on P . Then take $M^4 =$ Riemannian completion.



Data: ℓ points in \mathbb{R}^3 . $\implies V$ with $\Delta V = 0$

$$g = Vh + V^{-1}\theta^2$$

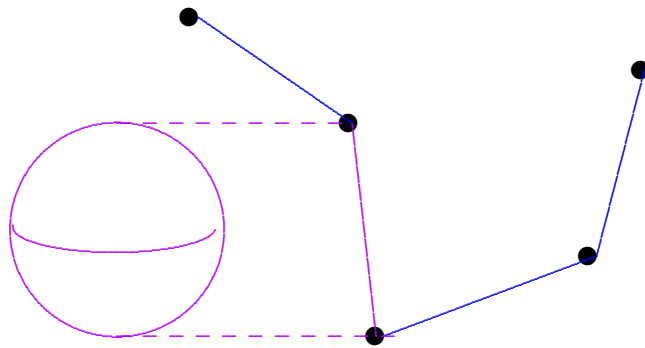
on P . Then take $M^4 =$ Riemannian completion.



Data: ℓ points in \mathbb{R}^3 . $\implies V$ with $\Delta V = 0$

$$g = Vh + V^{-1}\theta^2$$

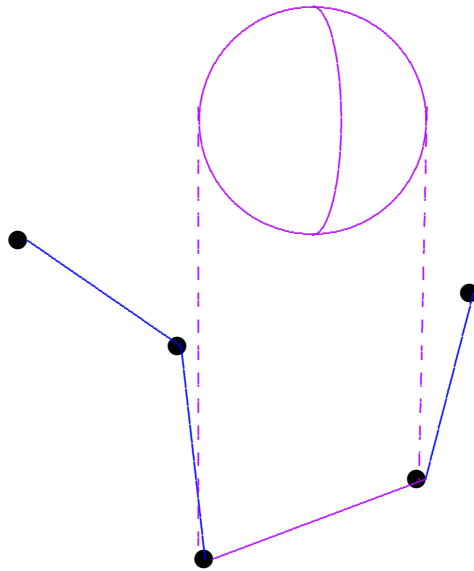
on P . Then take $M^4 =$ Riemannian completion.



Data: ℓ points in \mathbb{R}^3 . $\implies V$ with $\Delta V = 0$

$$g = Vh + V^{-1}\theta^2$$

on P . Then take $M^4 =$ Riemannian completion.



Data: ℓ points in \mathbb{R}^3 . $\implies V$ with $\Delta V = 0$

$$g = Vh + V^{-1}\theta^2$$

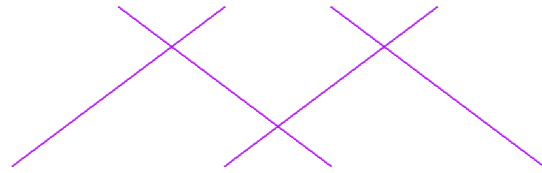
on P . Then take $M^4 =$ Riemannian completion.

Deform retracts to $k = \ell - 1$ copies of S^2 ,

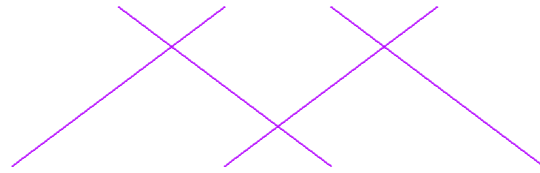
Deform retracts to $k = \ell - 1$ copies of S^2 ,
each with self-intersection -2 ,

Deform retracts to $k = \ell - 1$ copies of S^2 ,
each with self-intersection -2 ,
meeting transversely, & forming connected set:

Deform retracts to $k = \ell - 1$ copies of S^2 ,
each with self-intersection -2 ,
meeting transversely, & forming connected set:

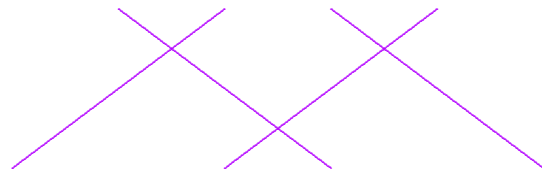


Deform retracts to $k = \ell - 1$ copies of S^2 ,
each with self-intersection -2 ,
meeting transversely, & forming connected set:



Configuration dual to Dynkin diagram A_k :

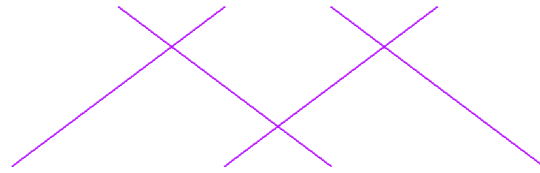
Deform retracts to $k = \ell - 1$ copies of S^2 ,
each with self-intersection -2 ,
meeting transversely, & forming connected set:



Configuration dual to Dynkin diagram A_k :



Deform retracts to $k = \ell - 1$ copies of S^2 ,
each with self-intersection -2 ,
meeting transversely, & forming connected set:

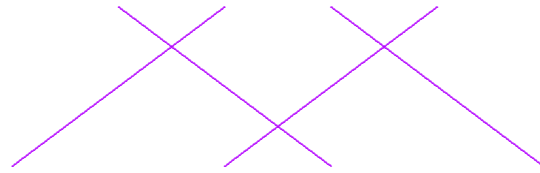


Configuration dual to Dynkin diagram A_k :



Diffeotype:

Deform retracts to $k = \ell - 1$ copies of S^2 ,
each with self-intersection -2 ,
meeting transversely, & forming connected set:



Configuration dual to Dynkin diagram A_k :



Diffeotype:

Plumb together k copies of T^*S^2
according to diagram.

Gibbons-Hawking gravitational instantons:

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2^{\rho_j}}$$

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2^{\rho_j}}$$

$$d\theta = \star dV$$

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\rho_j}$$

$$d\theta = \star dV$$

Kähler with respect to three complex structures

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2^{\rho_j}}$$

$$d\theta = \star dV$$

Kähler with respect to three complex structures

$$dx \mapsto V^{-1}\theta, \quad dy \mapsto dz$$

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2^{\rho_j}}$$

$$d\theta = \star dV$$

Kähler with respect to three complex structures

$$dx \mapsto V^{-1}\theta, \quad dy \mapsto dz$$

$$dy \mapsto V^{-1}\theta, \quad dz \mapsto dx$$

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\rho_j}$$

$$d\theta = \star dV$$

Kähler with respect to three complex structures

$$dx \mapsto V^{-1}\theta, \quad dy \mapsto dz$$

$$dy \mapsto V^{-1}\theta, \quad dz \mapsto dx$$

$$dz \mapsto V^{-1}\theta, \quad dx \mapsto dy$$

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\rho_j}$$

$$d\theta = \star dV$$

Kähler with respect to three complex structures

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\rho_j}$$

$$d\theta = \star dV$$

Kähler with respect to three complex structures

Hence holonomy $\subset \mathbf{Sp}(1) = \mathbf{SU}(2)$.

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\rho_j}$$

$$d\theta = \star dV$$

Kähler with respect to three complex structures

Hence holonomy $\subset \mathbf{Sp}(1) = \mathbf{SU}(2)$.

Hence Ricci-flat!

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\rho_j}$$

$$d\theta = \star dV$$

Kähler with respect to three complex structures

Hence holonomy $\subset \mathbf{Sp}(1) = \mathbf{SU}(2)$.

Hence Ricci-flat!

Calabi later called such metrics “hyper-Kähler.”

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2^{\rho_j}}$$

$$d\theta = \star dV$$

Kähler with respect to three complex structures

Hence holonomy $\subset \mathbf{Sp}(1) = \mathbf{SU}(2)$.

Hence Ricci-flat!

Calabi later called such metrics “hyper-Kähler.”

$M \rightarrow \mathbb{R}^3$ hyper-Kähler moment map of S^1 action.

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2\rho_j}$$

$$d\theta = \star dV$$

Kähler with respect to three complex structures

Hence holonomy $\subset \mathbf{Sp}(1) = \mathbf{SU}(2)$.

Hence Ricci-flat!

Calabi later called such metrics “hyper-Kähler.”

Gibbons and Hawking were unaware of all this!

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2^{\rho_j}}$$

$$d\theta = \star dV$$

Gibbons-Hawking gravitational instantons:

Gibbons-Hawking gravitational instantons:

These spaces have just one end,

Gibbons-Hawking gravitational instantons:

These spaces have just one end,

cf. Cheeger-Gromoll splitting theorem!

Gibbons-Hawking gravitational instantons:

These spaces have just one end,

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

When $\kappa = 0$, they are ALE:

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2^{\rho_j}}$$

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

When $\kappa = 0$, they are ALE:

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

When $\kappa = 0$, they are ALE:

“Asymptotically locally Euclidean”

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

When $\kappa = 0$, they are ALE:

“Asymptotically locally Euclidean”

$$g_{jk} = \delta_{jk} + O(|x|^{-4})$$

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

When $\kappa = 0$, they are ALE:

“Asymptotically locally Euclidean”

$$g_{jk} = \delta_{jk} + O(|x|^{-4})$$

In particular, volume of large ball is

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

When $\kappa = 0$, they are ALE:

“Asymptotically locally Euclidean”

$$g_{jk} = \delta_{jk} + O(|x|^{-4})$$

In particular, volume of large ball is

$$\text{Vol}(B_\rho) \sim \text{const} \cdot \rho^4$$

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

When $\kappa = 0$, they are ALE:

“Asymptotically locally Euclidean”

$$g_{jk} = \delta_{jk} + O(|x|^{-4})$$

In particular, volume of large ball is

$$\text{Vol}(B_\rho) \sim \frac{\pi^2/2}{\ell} \cdot \rho^4$$

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

When $\kappa = 0$, they are ALE:

“Asymptotically locally Euclidean”

$$g_{jk} = \delta_{jk} + O(|x|^{-4})$$

In particular, volume of large ball is

$$\text{Vol}(B_\rho) \sim \frac{\pi^2/2}{\ell} \cdot \rho^4$$

Notice that $\ell = 1$ case is just flat \mathbb{R}^4 !

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

When $\kappa = 0$, they are ALE:

“Asymptotically locally Euclidean”

$$g_{jk} = \delta_{jk} + O(|x|^{-4})$$

In particular, volume of large ball is

$$\text{Vol}(B_\rho) \sim \frac{\pi^2/2}{\ell} \cdot \rho^4$$

Notice that $\ell = 1$ case is just flat \mathbb{R}^4 !

cf. Bishop-Gromov inequality!

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

But when $\kappa \neq 0$, they are instead ALF:

Gibbons-Hawking gravitational instantons:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = \kappa^2 + \sum_{j=1}^{\ell} \frac{1}{2^{\rho_j}}$$

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

But when $\kappa \neq 0$, they are instead ALF:

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

But when $\kappa \neq 0$, they are instead ALF:

“Asymptotically locally flat”

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

But when $\kappa \neq 0$, they are instead ALF:

“Asymptotically locally flat”

Curvature still falls off at infinity,

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

But when $\kappa \neq 0$, they are instead ALF:

“Asymptotically locally flat”

Curvature still falls off at infinity,

$$|\mathcal{R}| \sim \text{const} \cdot \rho^{-3}$$

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

But when $\kappa \neq 0$, they are instead ALF:

“Asymptotically locally flat”

Curvature still falls off at infinity,

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

But when $\kappa \neq 0$, they are instead ALF:

“Asymptotically locally flat”

Curvature still falls off at infinity,

but volume growth is only cubic:

$$\text{Vol}(B_\rho) \sim \text{const} \cdot \rho^3$$

Gibbons-Hawking gravitational instantons:

These spaces have just one end, $\approx (\mathbb{R}^4 - \{0\})/\mathbb{Z}_\ell$

But when $\kappa \neq 0$, they are instead ALF:

“Asymptotically locally flat”

Curvature still falls off at infinity,

but volume growth is only cubic:

$$\text{Vol}(B_\rho) \sim \text{const} \cdot \rho^3$$

This last property distinguishes the ALF spaces from other classes of gravitational instantons:

ALG, ALH, ALG*, ALH*, ...

Example.

Example. Taub-NUT metric:

Example. Taub-NUT metric:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = 1 + \frac{1}{2\rho}$$

Example. Taub-NUT metric:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = 1 + \frac{1}{2\rho}$$

Can also write as

$$g = \frac{r+1}{4r}dr^2 + r(1+r)[\sigma_1^2 + \sigma_2^2] + \frac{r}{r+1}\sigma_3^2$$

Example. Taub-NUT metric:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = 1 + \frac{1}{2\rho}$$

Can also write as

$$g = \frac{r+1}{4r}dr^2 + r(1+r)[\sigma_1^2 + \sigma_2^2] + \frac{r}{r+1}\sigma_3^2$$

for left-invariant coframe $\{\sigma_j\}$ on $S^3 = \mathbf{SU}(2)$.

Example. Taub-NUT metric:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = 1 + \frac{1}{2\rho}$$

Can also write as

$$g = \frac{r+1}{4r}dr^2 + r(1+r)[\sigma_1^2 + \sigma_2^2] + \frac{r}{r+1}\sigma_3^2$$

for left-invariant coframe $\{\sigma_j\}$ on $S^3 = \mathbf{SU}(2)$.

Taub-NUT becomes Hermitian metric on \mathbb{C}^2 .

Example. Taub-NUT metric:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = 1 + \frac{1}{2\varrho}$$

Can also write as

$$g = \frac{r+1}{4r}dr^2 + r(1+r)[\sigma_1^2 + \sigma_2^2] + \frac{r}{r+1}\sigma_3^2$$

for left-invariant coframe $\{\sigma_j\}$ on $S^3 = \mathbf{SU}(2)$.

Taub-NUT becomes Hermitian metric on \mathbb{C}^2 .

$$dr \mapsto \frac{2r}{1+r}\sigma_3, \quad \sigma_1 \mapsto \sigma_2$$

Example. Taub-NUT metric:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = 1 + \frac{1}{2\rho}$$

Can also write as

$$g = \frac{r+1}{4r}dr^2 + r(1+r)[\sigma_1^2 + \sigma_2^2] + \frac{r}{r+1}\sigma_3^2$$

for left-invariant coframe $\{\sigma_j\}$ on $S^3 = \mathbf{SU}(2)$.

Taub-NUT becomes Hermitian metric on \mathbb{C}^2 .

Example. Taub-NUT metric:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = 1 + \frac{1}{2\rho}$$

Can also write as

$$g = \frac{r+1}{4r}dr^2 + r(1+r)[\sigma_1^2 + \sigma_2^2] + \frac{r}{r+1}\sigma_3^2$$

for left-invariant coframe $\{\sigma_j\}$ on $S^3 = \mathbf{SU}(2)$.

Taub-NUT becomes Hermitian metric on \mathbb{C}^2 .

This J determines opposite orientation from the hyper-Kähler complex structures.

Example. Taub-NUT metric:

$$g = V(dx^2 + dy^2 + dz^2) + V^{-1}\theta^2$$

$$V = 1 + \frac{1}{2\rho}$$

Can also write as

$$g = \frac{r+1}{4r}dr^2 + r(1+r)[\sigma_1^2 + \sigma_2^2] + \frac{r}{r+1}\sigma_3^2$$

for left-invariant coframe $\{\sigma_j\}$ on $S^3 = \mathbf{SU}(2)$.

Taub-NUT becomes Hermitian metric on \mathbb{C}^2 .

Non-Kähler, but **conformally** Kähler!

Hawking also explored non-hyper-Kähler examples. . .

Example.

Example. Riemannian Schwarzschild metric:

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Kähler!

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Kähler! In fact, extremal Kähler!

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Kähler! In fact, extremal Kähler!

$$\bar{\partial} \nabla^{1,0} s = 0$$

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Kähler! In fact, extremal Kähler!

$$\bar{\partial} \nabla^{1,0} s = 0$$

Andrzej Derdziński '83:

Bach-flat Kähler metrics are extremal!

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Kähler! In fact, extremal Kähler!

$$\bar{\partial} \nabla^{1,0} s = 0$$

Andrzej Derdziński '83:

$$B_{ab} := (\nabla^c \nabla^d + \frac{1}{2} \dot{r}^{cd}) W_{acbd}$$

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Kähler! In fact, extremal Kähler!

$$\bar{\partial} \nabla^{1,0} s = 0$$

Andrzej Derdziński '83:

$$B_{ab} := (\nabla^c \nabla^d + \frac{1}{2} \dot{r}^{cd}) W_{acbd} = 0$$

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Hawking: set $t = 4m\theta$ and $\varrho = 2m + \frac{r^2}{8m}$.

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Hawking: set $t = 4m\theta$ and $\varrho = 2m + \frac{r^2}{8m}$.

This makes g into a Ricci-flat metric on $\mathbb{R}^2 \times S^2$.

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Hawking: set $t = 4m\theta$ and $\varrho = 2m + \frac{r^2}{8m}$.

This makes g into a Ricci-flat metric on $\mathbb{R}^2 \times S^2$.

$$g = dr^2 + r^2 d\theta^2 + 4m^2 g_{S^2} + O(r^2)$$

Example. Riemannian Schwarzschild metric:

$$g = \left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 + \varrho^2 g_{S^2}$$

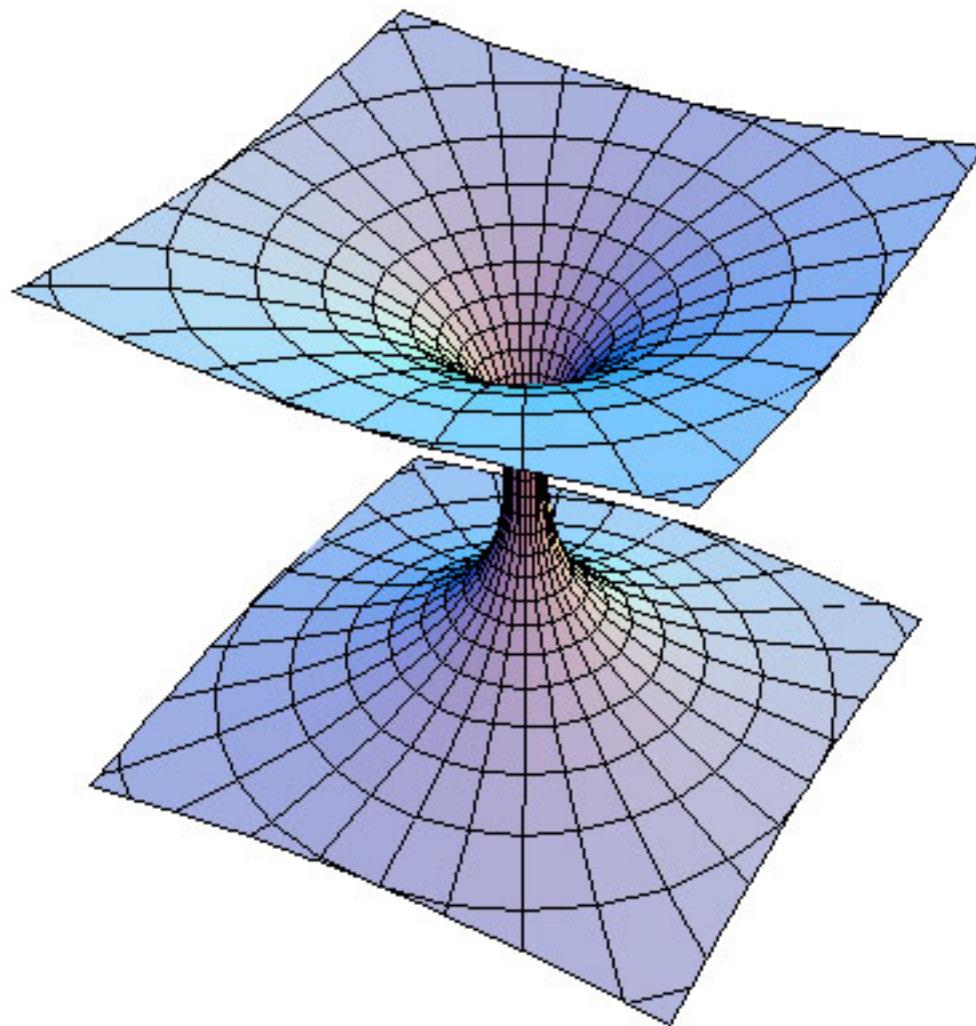
Conformal to

$$h = \frac{1}{\varrho^2} \left[\left(1 - \frac{2m}{\varrho}\right)^{-1} d\varrho^2 + \left(1 - \frac{2m}{\varrho}\right) dt^2 \right] + g_{S^2}$$

Hawking: set $t = 4m\theta$ and $\varrho = 2m + \frac{r^2}{8m}$.

This makes g into a Ricci-flat metric on $\mathbb{R}^2 \times S^2$.

Makes h into extremal Kähler metric on $\mathbb{C} \times \mathbb{C}P_1$.



$$\mathbb{R} \times S^2 \subset \mathbb{R}^2 \times S^2$$

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

Hitchin, Kronheimer, Cherkis-Hitchin, Minerbe, Hein, Chen-Chen, Hein-Sun-Viaclovsky-Zhang. . .

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

This might lend some credence to the aphorism...

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

“Mathematicians are like Frenchmen:

— J.W. von Goethe

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

“Mathematicians are like Frenchmen: you tell them something,

— J.W. von Goethe

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

“Mathematicians are like Frenchmen: you tell them something, they translate it into their own language,

— J.W. von Goethe

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

“Mathematicians are like Frenchmen: you tell them something, they translate it into their own language, and before you know it,

— J.W. von Goethe

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

“Mathematicians are like Frenchmen: you tell them something, they translate it into their own language, and before you know it, it’s something else entirely.”

— J.W. von Goethe

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

But now my French collaborators

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

But now my French collaborators Biquard and Gauduchon

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

But now my French collaborators Biquard and Gauduchon have fortunately done us all the favor of reminding us

Definition. *A gravitational instanton is a complete, non-compact, non-flat, Ricci-flat Riemannian 4-manifold.*

Many excellent mathematical papers cleverly narrow the definition for technical convenience, by assuming at the outset that the metric is hyper-Kähler.

But now my French collaborators Biquard and Gauduchon have fortunately done us all the favor of reminding us that the hyper-Kähler gravitons are only one small part of the story!

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat,*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric,*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric,*

\mathbb{T}^2 acts effectively and isometrically

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF ,*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, *ALF*,*

$$M - (\text{compact set}) \approx \Sigma \times \mathbb{R}^+$$

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, *ALF*,*

$$M - (\text{compact set}) \approx \Sigma \times \mathbb{R}^+$$

Σ finitely covered by S^3 or $S^2 \times S^1$

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF,*

$$M - (\text{compact set}) \approx \Sigma \times \mathbb{R}^+$$

Σ finitely covered by S^3 or $S^2 \times S^1$

Σ equipped with vector field T and 1-form η

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF,*

$$M - (\text{compact set}) \approx \Sigma \times \mathbb{R}^+$$

Σ finitely covered by S^3 or $S^2 \times S^1$

Σ equipped with vector field T and 1-form η

$$\eta(T) = 1, \quad \mathcal{L}_T \eta = 0,$$

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF,*

$$M - (\text{compact set}) \approx \Sigma \times \mathbb{R}^+$$

Σ finitely covered by S^3 or $S^2 \times S^1$

Σ equipped with vector field T and 1-form η

$$\eta(T) = 1, \quad \mathcal{L}_T \eta = 0, \quad \text{and}$$

$T\Sigma/T$ equipped with curvature +1 metric γ .

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, *ALF*,*

$$g = d\varrho^2 + \varrho^2\gamma + \eta^2 + \mathcal{U}$$

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF,*

$$g = d\rho^2 + \rho^2\gamma + \eta^2 + \mathcal{U}$$

$$\mathcal{U} = O(\rho^{-1}), \quad \nabla\mathcal{U} = O(\rho^{-2}), \quad \dots \quad \nabla^3\mathcal{U} = O(\rho^{-4})$$

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, *ALF*,*

$$\implies \text{Vol}(B_\rho) \sim \text{const} \cdot \rho^3$$

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, *ALF*, and Hermitian*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, *ALF*, and Hermitian with respect to some integrable complex structure J .*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, *ALF*, and Hermitian with respect to some integrable complex structure J .*

$$g(J\cdot, J\cdot) = g$$

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, *ALF*, and Hermitian with respect to some integrable complex structure J .*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, *ALF*, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler.*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, *ALF*, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*

Diffeomorphic to \mathbb{R}^4

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*

Diffeomorphic to $\mathbb{C}P_2 - \{pt\}$

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*
- *a metric of the Kerr family;*

Diffeomorphic to $S^2 \times \mathbb{R}^2$

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*
- *a metric of the Kerr family; or*
- *a metric in the Chen-Teo family.*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*
- *a metric of the Kerr family; or*
- *a metric in the Chen-Teo family.*

Diffeomorphic to $\mathbb{C}P_2 - S^1$

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*
- *a metric of the Kerr family; or*
- *a metric in the Chen-Teo family.*

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*
- *a metric of the Kerr family; or*
- *a metric in the Chen-Teo family.*

By the Riemannian Goldberg-Sachs Theorem,

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*
- *a metric of the Kerr family; or*
- *a metric in the Chen-Teo family.*

By the Riemannian Goldberg-Sachs Theorem, the Hermitian assumption is

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*
- *a metric of the Kerr family; or*
- *a metric in the Chen-Teo family.*

By the Riemannian Goldberg-Sachs Theorem, the Hermitian assumption is equivalent to assuming that the Ricci-flat g is conformally Kähler.

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*
- *a metric of the Kerr family; or*
- *a metric in the Chen-Teo family.*

This assertion is peculiar to dimension 4.

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*
- *a metric of the Kerr family; or*
- *a metric in the Chen-Teo family.*

This assertion is peculiar to dimension 4.
It is false in all higher dimensions!

Special character of dimension 4:

Special character of dimension 4:

On oriented (M^4, g) ,

$$\Lambda^2 = \Lambda^+ \oplus \Lambda^-$$

Special character of dimension 4:

On oriented (M^4, g) ,

$$\Lambda^2 = \Lambda^+ \oplus \Lambda^-$$

where Λ^\pm are (± 1) -eigenspaces of

$$\star : \Lambda^2 \rightarrow \Lambda^2,$$

$$\star^2 = 1.$$

Special character of dimension 4:

On oriented (M^4, g) ,

$$\Lambda^2 = \Lambda^+ \oplus \Lambda^-$$

where Λ^\pm are (± 1) -eigenspaces of

$$\star : \Lambda^2 \rightarrow \Lambda^2,$$

$$\star^2 = 1.$$

Λ^+ self-dual 2-forms.

Λ^- anti-self-dual 2-forms.

Special character of dimension 4:

On oriented (M^4, g) ,

$$\Lambda^2 = \Lambda^+ \oplus \Lambda^-$$

where Λ^\pm are (± 1) -eigenspaces of

$$\star : \Lambda^2 \rightarrow \Lambda^2,$$

$$\star^2 = 1.$$

Λ^+ self-dual 2-forms.

Λ^- anti-self-dual 2-forms.

Only depends on the conformal class

$$[g] := \{u^2 g \mid u : M \rightarrow \mathbb{R}^+\}.$$

Special character of dimension 4:

On oriented (M^4, g) ,

$$\Lambda^2 = \Lambda^+ \oplus \Lambda^-$$

where Λ^\pm are (± 1) -eigenspaces of

$$\star : \Lambda^2 \rightarrow \Lambda^2,$$

$$\star^2 = 1.$$

Λ^+ self-dual 2-forms.

Λ^- anti-self-dual 2-forms.

Reversing orientation interchanges $\Lambda^+ \leftrightarrow \Lambda^-$.

Riemann curvature of g

$$\mathcal{R} : \Lambda^2 \rightarrow \Lambda^2$$

Riemann curvature of g

$$\mathcal{R} : \Lambda^2 \rightarrow \Lambda^2$$

splits into 4 irreducible pieces:

$$\mathcal{R} = \left(\begin{array}{c|c} W_+ + \frac{s}{12} & \overset{\circ}{r} \\ \hline \overset{\circ}{r} & W_- + \frac{s}{12} \end{array} \right) .$$

Riemann curvature of g

$$\mathcal{R} : \Lambda^2 \rightarrow \Lambda^2$$

splits into 4 irreducible pieces:

	Λ^{+*}	Λ^{-*}
Λ^+	$W_+ + \frac{s}{12}$	$\overset{\circ}{r}$
Λ^-	$\overset{\circ}{r}$	$W_- + \frac{s}{12}$

Riemann curvature of g

$$\mathcal{R} : \Lambda^2 \rightarrow \Lambda^2$$

splits into 4 irreducible pieces:

	Λ^{+*}	Λ^{-*}
Λ^+	$W_+ + \frac{s}{12}$	$\overset{\circ}{r}$
Λ^-	$\overset{\circ}{r}$	$W_- + \frac{s}{12}$

where

s = scalar curvature

$\overset{\circ}{r}$ = trace-free Ricci curvature

W_+ = self-dual Weyl curvature

W_- = anti-self-dual Weyl curvature

Riemann curvature of g

$$\mathcal{R} : \Lambda^2 \rightarrow \Lambda^2$$

splits into 4 irreducible pieces:

	Λ^{+*}	Λ^{-*}
Λ^+	$W_+ + \frac{s}{12}$	$\overset{\circ}{r}$
Λ^-	$\overset{\circ}{r}$	$W_- + \frac{s}{12}$

where

s = scalar curvature

$\overset{\circ}{r}$ = trace-free Ricci curvature

W_+ = self-dual Weyl curvature (*conformally invariant*)

W_- = anti-self-dual Weyl curvature //

Theorem (Biquard-Gauduchon '23). *Let (M^4, g) be a smooth, complete, non-flat, Ricci-flat 4-manifold that is toric, ALF, and Hermitian with respect to some integrable complex structure J . Also assume that (M, g, J) is not Kähler. Then (M, g) is one of the following explicit examples:*

- *the (reverse-oriented) Taub-NUT metric;*
- *the Taub-bolt metric;*
- *a metric of the Kerr family; or*
- *a metric in the Chen-Teo family.*

Each of the metrics g in question is conformal to an extremal Kähler metric h with $s > 0$.

Each of the metrics g in question is conformal to an extremal Kähler metric h with $s > 0$.

Indeed, one has $g = \text{const} \cdot s^{-2}h$.

Each of the metrics g in question is conformal to an extremal Kähler metric h with $s > 0$.

Indeed, one has $g = \text{const} \cdot s^{-2}h$.

Notational warning:

Here, g and h interchanged relative to our e-print!

Each of the metrics g in question is conformal to an extremal Kähler metric h with $s > 0$.

Indeed, one has $g = \text{const} \cdot s^{-2}h$.

Each of the metrics g in question is conformal to an extremal Kähler metric h with $s > 0$.

Indeed, one has $g = \text{const} \cdot s^{-2}h$.

The end geometry of these Kähler metrics locally modeled on $\mathbb{C}\mathbb{P}_1 \times$ hyperbolic cusp.

Each of the metrics g in question is conformal to an extremal Kähler metric h with $s > 0$.

Indeed, one has $g = \text{const} \cdot s^{-2}h$.

The end geometry of these Kähler metrics locally modeled on $\mathbb{C}\mathbb{P}_1 \times$ hyperbolic cusp.

Finite 4-volume, w/ cross-section of 3-volume $\rightarrow 0$.

Each of the metrics g in question is conformal to an extremal Kähler metric h with $s > 0$.

Indeed, one has $g = \text{const} \cdot s^{-2}h$.

The end geometry of these Kähler metrics locally modeled on $\mathbb{C}\mathbb{P}_1 \times$ hyperbolic cusp.

Finite 4-volume, w/ cross-section of 3-volume $\rightarrow 0$.

Fact that h has $s > 0$ means that both g and h satisfy Peng Wu's criterion:

Each of the metrics g in question is conformal to an extremal Kähler metric h with $s > 0$.

Indeed, one has $g = \text{const} \cdot s^{-2}h$.

The end geometry of these Kähler metrics locally modeled on $\mathbb{C}\mathbb{P}_1 \times$ hyperbolic cusp.

Finite 4-volume, w/ cross-section of 3-volume $\rightarrow 0$.

Fact that h has $s > 0$ means that both g and h satisfy Peng Wu's criterion:

$$\det(W^+) > 0$$

Each of the metrics g in question is conformal to an extremal Kähler metric h with $s > 0$.

Indeed, one has $g = \text{const} \cdot s^{-2}h$.

The end geometry of these Kähler metrics locally modeled on $\mathbb{C}\mathbb{P}_1 \times$ hyperbolic cusp.

Finite 4-volume, w/ cross-section of 3-volume $\rightarrow 0$.

Fact that h has $s > 0$ means that both g and h satisfy Peng Wu's criterion:

$$\det(W^+) > 0$$

$$W^+ = \begin{bmatrix} \alpha & & \\ & \beta & \\ & & \gamma \end{bmatrix}, \quad \alpha \geq \beta \geq \gamma$$

Each of the metrics g in question is conformal to an extremal Kähler metric h with $s > 0$.

Indeed, one has $g = \text{const} \cdot s^{-2}h$.

The end geometry of these Kähler metrics locally modeled on $\mathbb{C}\mathbb{P}_1 \times$ hyperbolic cusp.

Finite 4-volume, w/ cross-section of 3-volume $\rightarrow 0$.

Fact that h has $s > 0$ means that both g and h satisfy Peng Wu's criterion:

$$\det(W^+) > 0$$

$$W^+ \sim \begin{bmatrix} + & & \\ & - & \\ & & - \end{bmatrix}$$

Equation $\delta W^+ = 0$ conformally invariant w/ weight.

Equation $\delta W^+ = 0$ conformally invariant w/ weight.

If $g = f^2 h$ satisfies

$$\delta W^+ = 0$$

Equation $\delta W^+ = 0$ conformally invariant w/ weight.

If $g = f^2 h$ satisfies

$$\delta W^+ = 0$$

then h instead satisfies

Equation $\delta W^+ = 0$ conformally invariant w/ weight.

If $g = f^2 h$ satisfies

$$\delta W^+ = 0$$

then h instead satisfies

$$\delta(fW^+) = 0$$

Equation $\delta W^+ = 0$ conformally invariant w/ weight.

If $g = f^2 h$ satisfies

$$\delta W^+ = 0$$

then h instead satisfies

$$\delta(fW^+) = 0$$

which in turn implies the Weitzenböck formula

$$0 = \nabla^* \nabla (fW^+) + \frac{s}{2} fW^+ - 6fW^+ \circ W^+ + 2f|W^+|^2 I$$

Equation $\delta W^+ = 0$ conformally invariant w/ weight.

If $g = f^2 h$ satisfies

$$\delta W^+ = 0$$

then h instead satisfies

$$\delta(fW^+) = 0$$

which in turn implies the Weitzenböck formula

$$0 = \nabla^* \nabla (fW^+) + \frac{s}{2} fW^+ - 6fW^+ \circ W^+ + 2f|W^+|^2 I$$

for $fW^+ \in \text{End}(\Lambda^+)$.

Application to Wu's criterion:

Let $\alpha \geq \beta \geq \gamma$ be eigenvalues of W^+ :

$$W^+ = \begin{bmatrix} \alpha & & \\ & \beta & \\ & & \gamma \end{bmatrix}$$

$$\alpha + \beta + \gamma = 0$$

$$\alpha > 0, \quad \gamma < 0, \quad \text{if } W^+ \neq 0$$

$$\det(W^+) = \alpha\beta\gamma$$

$$\det(W^+) > 0 \implies \alpha \text{ has multiplicity 1.}$$

Application to Wu's criterion:

Let $\alpha \geq \beta \geq \gamma$ be eigenvalues of W^+ :

$$W^+ = \begin{bmatrix} \alpha & & \\ & \beta & \\ & & \gamma \end{bmatrix}$$

$$\alpha + \beta + \gamma = 0$$

$$\alpha > 0, \quad \gamma < 0, \quad \text{if } W^+ \neq 0$$

$$\det(W^+) = \alpha\beta\gamma$$

$$\det(W^+) > 0 \implies \alpha \text{ has multiplicity } 1.$$

So $\alpha = \alpha_g : M \rightarrow \mathbb{R}^+$ a smooth function.

Application to Wu's criterion:

Let $\alpha \geq \beta \geq \gamma$ be eigenvalues of W^+ :

$$W^+ = \begin{bmatrix} \alpha & & \\ & \beta & \\ & & \gamma \end{bmatrix}$$

$$\alpha + \beta + \gamma = 0$$

$$\alpha > 0, \quad \gamma < 0, \quad \text{if } W^+ \neq 0$$

$$\det(W^+) = \alpha\beta\gamma$$

$$\det(W^+) > 0 \implies \alpha \text{ has multiplicity } 1.$$

So $\alpha = \alpha_g : M \rightarrow \mathbb{R}^+$ a smooth function. Set

$$f = \alpha_g^{-1/3}, \quad h = f^{-2}g = \alpha_g^{2/3}g.$$

Lemma. *Let (M, g) be an oriented, simply-connected Einstein 4-manifold*

Lemma. *Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies*

Lemma. *Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies*

$$\det(W^+) > 0$$

everywhere.

Lemma. *Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies*

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric

Lemma. Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g,$$

Lemma. *Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies*

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g,$$

and let $\omega \in \Lambda^+$ be defined

Lemma. *Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies*

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g,$$

and let $\omega \in \Lambda^+$ be defined by

$$W_h^+(\omega) = \alpha_h \omega, \quad |\omega|_h = \sqrt{2}.$$

Lemma. *Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies*

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g,$$

and let $\omega \in \Lambda^+$ be defined (up to sign) by

$$W_h^+(\omega) = \alpha_h \omega, \quad |\omega|_h = \sqrt{2}.$$

Lemma. Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g,$$

and let $\omega \in \Lambda^+$ be defined (up to sign) by

$$W_h^+(\omega) = \alpha_h \omega, \quad |\omega|_h = \sqrt{2}.$$

Then

$$0 \geq |\nabla \omega|^2 + 3 \langle \omega, (d + d^*)^2 \omega \rangle$$

Lemma. Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g,$$

and let $\omega \in \Lambda^+$ be defined (up to sign) by

$$W_h^+(\omega) = \alpha_h \omega, \quad |\omega|_h = \sqrt{2}.$$

Then

$$0 \geq |\nabla \omega|^2 + 3 \langle \omega, (d + d^*)^2 \omega \rangle$$

at every point of M ,

Lemma. Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g,$$

and let $\omega \in \Lambda^+$ be defined (up to sign) by

$$W_h^+(\omega) = \alpha_h \omega, \quad |\omega|_h = \sqrt{2}.$$

Then

$$0 \geq |\nabla \omega|^2 + 3 \langle \omega, (d + d^*)^2 \omega \rangle$$

at every point of M , with respect to the conformally rescaled metric h .

Lemma. Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g,$$

and let $\omega \in \Lambda^+$ be defined (up to sign) by

$$W_h^+(\omega) = \alpha_h \omega, \quad |\omega|_h = \sqrt{2}.$$

Then

$$3 d[\omega \wedge \star d\omega] \geq \star \left(\frac{1}{2} |\nabla \omega|^2 + 3 |d\omega|^2 \right).$$

at every point of M , with respect to the conformally rescaled metric h .

Proposition.

Proposition. *Let (M, g) be an oriented, simply-connected Riemannian 4-manifold*

Proposition. *Let (M, g) be an oriented, simply-connected Riemannian 4-manifold that satisfies $\delta W^+ = 0$ and $\det(W^+) > 0$ everywhere.*

Proposition. *Let (M, g) be an oriented, simply-connected Riemannian 4-manifold that satisfies $\delta W^+ = 0$ and $\det(W^+) > 0$ everywhere. Also suppose that M is a nested union $M = \cup_j U_j$*

Proposition. *Let (M, g) be an oriented, simply-connected Riemannian 4-manifold that satisfies $\delta W^+ = 0$ and $\det(W^+) > 0$ everywhere. Also suppose that M is a nested union $M = \cup_j U_j$ of compact domains*

Proposition. Let (M, g) be an oriented, simply-connected Riemannian 4-manifold that satisfies $\delta W^+ = 0$ and $\det(W^+) > 0$ everywhere. Also suppose that M is a nested union $M = \cup_j U_j$ of compact domains

$$U_1 \subset U_2 \subset \cdots \subset U_j \subset \cdots$$

with smooth boundary

Proposition. Let (M, g) be an oriented, simply-connected Riemannian 4-manifold that satisfies $\delta W^+ = 0$ and $\det(W^+) > 0$ everywhere. Also suppose that M is a nested union $M = \cup_j U_j$ of compact domains

$$U_1 \subset U_2 \subset \cdots \subset U_j \subset \cdots$$

with smooth boundary such that

$$\lim_{j \rightarrow \infty} \int_{\partial U_j} \omega \wedge \star d\omega = 0.$$

Proposition. Let (M, g) be an oriented, simply-connected Riemannian 4-manifold that satisfies $\delta W^+ = 0$ and $\det(W^+) > 0$ everywhere. Also suppose that M is a nested union $M = \cup_j U_j$ of compact domains

$$U_1 \subset U_2 \subset \cdots \subset U_j \subset \cdots$$

with smooth boundary such that

$$\lim_{j \rightarrow \infty} \int_{\partial U_j} \omega \wedge \star d\omega = 0.$$

Then $h = \alpha_g^{2/3} g$ is a Kähler metric.

Proposition. Let (M, g) be an oriented, simply-connected Riemannian 4-manifold that satisfies $\delta W^+ = 0$ and $\det(W^+) > 0$ everywhere. Also suppose that M is a nested union $M = \cup_j U_j$ of compact domains

$$U_1 \subset U_2 \subset \cdots \subset U_j \subset \cdots$$

with smooth boundary such that

$$\lim_{j \rightarrow \infty} \int_{\partial U_j} \omega \wedge \star d\omega = 0.$$

Then $h = \alpha_g^{2/3} g$ is a Kähler metric.

Proof.

$$3 d[\omega \wedge \star d\omega] \geq \star \left(\frac{1}{2} |\nabla \omega|^2 + 3 |d\omega|^2 \right).$$

Lemma. Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g,$$

and let $\omega \in \Lambda^+$ be defined (up to sign) by

$$W_h^+(\omega) = \alpha_h \omega, \quad |\omega|_h = \sqrt{2}.$$

Then

$$3 d[\omega \wedge \star d\omega] \geq \star \left(\frac{1}{2} |\nabla \omega|^2 + 3 |d\omega|^2 \right).$$

at every point of M , with respect to the conformally rescaled metric h .

Lemma. Let (M, g) be an oriented, simply-connected Einstein 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g,$$

and let $\omega \in \Lambda^+$ be defined (up to sign) by

$$W_h^+(\omega) = \alpha_h \omega, \quad |\omega|_h = \sqrt{2}.$$

Then

$$3 d[\omega \wedge \star d\omega] \geq \star \left(\frac{1}{2} |\nabla \omega|^2 + 3 |d\omega|^2 \right).$$

at every point of M , with respect to the conformally rescaled metric h . Moreover,

$$2\sqrt{6} |W^+|_h + |s_h| \geq |\omega \wedge \star d\omega|^2$$

everywhere on (M, h) .

Theorem.

Theorem. *Let (M, g) be an oriented, simply-connected Ricci-flat*

Theorem. *Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies*

Theorem. *Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies*

$$\det(W^+) > 0$$

everywhere.

Theorem. *Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies*

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric

Theorem. *Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies*

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g.$$

Theorem. *Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies*

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g.$$

Suppose that M is expressed as a nested union
 $M = \cup_j U_j$

Theorem. *Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies*

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g.$$

Suppose that M is expressed as a nested union $M = \cup_j U_j$ of compact domains

$$U_1 \subset U_2 \subset \cdots \subset U_j \subset \cdots$$

Theorem. Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g.$$

Suppose that M is expressed as a nested union $M = \cup_j U_j$ of compact domains

$$U_1 \subset U_2 \subset \cdots \subset U_j \subset \cdots$$

with smooth ∂

Theorem. Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g.$$

Suppose that M is expressed as a nested union $M = \cup_j U_j$ of compact domains

$$U_1 \subset U_2 \subset \cdots \subset U_j \subset \cdots$$

with smooth ∂ s.t. $\text{Vol}^{(3)}(\partial U_j, h) < C$ and

Theorem. Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g.$$

Suppose that M is expressed as a nested union $M = \cup_j U_j$ of compact domains

$$U_1 \subset U_2 \subset \cdots \subset U_j \subset \cdots$$

with smooth ∂ s.t. $\text{Vol}^{(3)}(\partial U_j, h) < C$ and

$$\lim_{j \rightarrow \infty} \int_{\partial U_j} |W^+{}_h| d\check{\mu}_h = \lim_{j \rightarrow \infty} \int_{\partial U_j} |s_h| d\check{\mu}_h = 0.$$

Theorem. Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g.$$

Suppose that M is expressed as a nested union $M = \cup_j U_j$ of compact domains

$$U_1 \subset U_2 \subset \cdots \subset U_j \subset \cdots$$

with smooth ∂ s.t. $\text{Vol}^{(3)}(\partial U_j, h) < \mathbf{C}$ and

$$\lim_{j \rightarrow \infty} \int_{\partial U_j} |W^+ h| d\check{\mu}_h = \lim_{j \rightarrow \infty} \int_{\partial U_j} |s_h| d\check{\mu}_h = 0.$$

Then (M, h) is an extremal Kähler manifold

Theorem. Let (M, g) be an oriented, simply-connected Ricci-flat 4-manifold that satisfies

$$\det(W^+) > 0$$

everywhere. Let h be the conformally rescaled metric defined by

$$h = \alpha_g^{2/3} g.$$

Suppose that M is expressed as a nested union $M = \cup_j U_j$ of compact domains

$$U_1 \subset U_2 \subset \cdots \subset U_j \subset \cdots$$

with smooth ∂ s.t. $\text{Vol}^{(3)}(\partial U_j, h) < C$ and

$$\lim_{j \rightarrow \infty} \int_{\partial U_j} |W^+ h| d\check{\mu}_h = \lim_{j \rightarrow \infty} \int_{\partial U_j} |s_h| d\check{\mu}_h = 0.$$

Then (M, h) is an extremal Kähler manifold with non-constant, positive scalar curvature.

Theorem A.

Theorem A. *Let (M, g_0) be*

Theorem A. *Let (M, g_0) be one of the ALF toric Hermitian gravitational instantons*

Theorem A. *Let (M, g_0) be one of the **ALF** toric **Hermitian** gravitational instantons featured in **Biquard-Gauduchon** classification.*

Theorem A. *Let (M, g_0) be one of the ALF toric Hermitian gravitational instantons featured in Biquard-Gauduchon classification. Then any other Ricci-flat Riemannian metric g on M*

Theorem A. *Let (M, g_0) be one of the ALF toric Hermitian gravitational instantons featured in Biquard-Gauduchon classification. Then any other Ricci-flat Riemannian metric g on M which is sufficiently C_1^3 -close to g_0*

Theorem A. Let (M, g_0) be one of the *ALF toric Hermitian* gravitational instantons featured in *Biquard-Gauduchon* classification. Then any other Ricci-flat Riemannian metric g on M which is sufficiently C_1^3 -close to g_0

$$\|g - g_0\|_{C_1^3} < \varepsilon$$

Theorem A. Let (M, g_0) be one of the *ALF toric Hermitian* gravitational instantons featured in *Biquard-Gauduchon* classification. Then any other Ricci-flat Riemannian metric g on M which is sufficiently C_1^3 -close to g_0

$$\|g - g_0\|_{C_1^3} < \varepsilon$$

$$\|\mathcal{U}\|_{C_1^3} := \sup_M \sum_{j=0}^3 (1 + \text{dist})^{j+1} |\nabla^j \mathcal{U}|_{g_0}$$

Theorem A. Let (M, g_0) be one of the *ALF toric Hermitian* gravitational instantons featured in *Biquard-Gauduchon* classification. Then any other Ricci-flat Riemannian metric g on M which is sufficiently C_1^3 -close to g_0

$$\|g - g_0\|_{C_1^3} < \varepsilon$$

$$\|\mathcal{U}\|_{C_1^3} := \sup_M \sum_{j=0}^3 (1 + \text{dist})^{j+1} |\nabla^j \mathcal{U}|_{g_0}$$

$$|\mathcal{U}|_{g_0} = O(\varrho^{-1}), \quad |\nabla \mathcal{U}|_{g_0} = O(\varrho^{-2}), \quad \dots$$

Theorem A. *Let (M, g_0) be one of the ALF toric Hermitian gravitational instantons featured in Biquard-Gauduchon classification. Then any other Ricci-flat Riemannian metric g on M which is sufficiently C_1^3 -close to g_0*

Theorem A. *Let (M, g_0) be one of the ALF toric Hermitian gravitational instantons featured in Biquard-Gauduchon classification. Then any other Ricci-flat Riemannian metric g on M which is sufficiently C_1^3 -close to g is conformal to some strictly extremal Kähler metric h ,*

Theorem A. *Let (M, g_0) be one of the ALF toric Hermitian gravitational instantons featured in Biquard-Gauduchon classification. Then any other Ricci-flat Riemannian metric g on M which is sufficiently C_1^3 -close to g is conformal to some strictly extremal Kähler metric h , and so is, in particular, Hermitian.*

Theorem A. *Let (M, g_0) be one of the ALF toric Hermitian gravitational instantons featured in Biquard-Gauduchon classification. Then any other Ricci-flat Riemannian metric g on M which is sufficiently C_1^3 -close to g_0 is conformal to some strictly extremal Kähler metric h , and so is, in particular, Hermitian. Moreover, every such g carries at least one Killing field.*

Proposition.

Proposition. *In the setting of Theorem A,*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group,*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$.*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic,*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus,*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric ALF*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric ALF Hermitian*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric *ALF Hermitian* gravitational instantons*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric *ALF Hermitian* gravitational instantons classified by Biquard-Gauduchon.*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric *ALF Hermitian* gravitational instantons classified by Biquard-Gauduchon.*

Corollary. *Let (M, g_0) be a toric *Hermitian ALF* gravitational instanton*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric *ALF Hermitian* gravitational instantons classified by Biquard-Gauduchon.*

Corollary. *Let (M, g_0) be a toric *Hermitian ALF* gravitational instanton for which the corresponding vector field T on Σ is *not periodic*.*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric *ALF Hermitian* gravitational instantons classified by Biquard-Gauduchon.*

Corollary. *Let (M, g_0) be a toric *Hermitian ALF* gravitational instanton for which the corresponding vector field T on Σ is not periodic. Then any Ricci-flat metric g on M*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric *ALF Hermitian* gravitational instantons classified by Biquard-Gauduchon.*

Corollary. *Let (M, g_0) be a toric *Hermitian ALF* gravitational instanton for which the corresponding vector field T on Σ is not periodic. Then any Ricci-flat metric g on M which is sufficiently C_1^3 close to g_0 must be*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric *ALF Hermitian* gravitational instantons classified by Biquard-Gauduchon.*

Corollary. *Let (M, g_0) be a toric *Hermitian ALF* gravitational instanton for which the corresponding vector field T on Σ is not periodic. Then any Ricci-flat metric g on M which is sufficiently C_1^3 close to g_0 must be one of the toric gravitational instantons*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric *ALF Hermitian* gravitational instantons classified by Biquard-Gauduchon.*

Corollary. *Let (M, g_0) be a toric *Hermitian ALF* gravitational instanton for which the corresponding vector field T on Σ is not periodic. Then any Ricci-flat metric g on M which is sufficiently C_1^3 close to g_0 must be one of the toric gravitational instantons classified by Biquard-Gauduchon.*

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric *ALF Hermitian* gravitational instantons classified by Biquard-Gauduchon.*

This is suggestive, but not quite definitive.

Proposition. *In the setting of Theorem A, the identity component $Iso_0(M, g)$ of the isometry group is a compact connected Lie group, and the closure of the exponential of the extremal vector field defines a sub-torus $\mathbb{T} \subset Iso_0(M, g)$. If the extremal vector field is non-periodic, then \mathbb{T} is a 2-torus, and (M, g) is another one of the toric *ALF Hermitian* gravitational instantons classified by Biquard-Gauduchon.*

This is suggestive, but not quite definitive.

However, if g_0 is Kerr or Taub-bolt, we can prove a more definitive rigidity result, because g_0 then has both $\det(W^+) > 0$ and $\det(W^-) > 0$.

Theorem B.

Theorem B. *Let (M, g_0) be*

Theorem B. *Let (M, g_0) be a Taub-bolt*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton,*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 .*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 . Then (M, g) is once again*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 . Then (M, g) is once again a Taub-bolt*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 . Then (M, g) is once again a Taub-bolt or Kerr*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 . Then (M, g) is once again a Taub-bolt or Kerr gravitational instanton.*

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 . Then (M, g) is once again a Taub-bolt or Kerr gravitational instanton.*

M admits complex structures J_{\pm} for 2 orientations,

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 . Then (M, g) is once again a Taub-bolt or Kerr gravitational instanton.*

M admits complex structures J_{\pm} for 2 orientations, and Kähler metrics h_{\pm} in conformal class $[g]$.

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 . Then (M, g) is once again a Taub-bolt or Kerr gravitational instanton.*

M admits complex structures J_{\pm} for 2 orientations, and Kähler metrics h_{\pm} in conformal class $[g]$.

Apostolov-Calderbank-Gauduchon '16:

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 . Then (M, g) is once again a Taub-bolt or Kerr gravitational instanton.*

M admits complex structures J_{\pm} for 2 orientations, and Kähler metrics h_{\pm} in conformal class $[g]$.

Apostolov-Calderbank-Gauduchon '16:

“ambi-Kähler” structure.

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 . Then (M, g) is once again a Taub-bolt or Kerr gravitational instanton.*

M admits complex structures J_{\pm} for 2 orientations, and Kähler metrics h_{\pm} in conformal class $[g]$.

Apostolov-Calderbank-Gauduchon '16:

“ambi-Kähler” structure.

Ambi-Kähler Einstein metrics are ambi-toric!

Theorem B. *Let (M, g_0) be a Taub-bolt or Kerr gravitational instanton, and let g be another Ricci-flat metric on M that is sufficiently C_1^3 -close to g_0 . Then (M, g) is once again a Taub-bolt or Kerr gravitational instanton.*

Breaking News...

Breaking News...

In [arXiv:2310.13197](#), Mingyang Li asserts that any Hermitian ALF gravitational instanton is toric!

Breaking News...

In [arXiv:2310.13197](#), Mingyang Li asserts that any Hermitian ALF gravitational instanton is toric!

Assuming this is correct, our methods then prove:

Breaking News...

In [arXiv:2310.13197](https://arxiv.org/abs/2310.13197), Mingyang Li asserts that any Hermitian ALF gravitational instanton is toric!

Assuming this is correct, our methods then prove:

Conjecture. *Let (M, g_0) be any toric Hermitian ALF gravitational instanton. Then any Ricci-flat metric g on M which is sufficiently C_1^3 close to g_0 must be another one of the gravitational instantons classified by Biquard-Gauduchon.*





Thank you for inviting me!





It's a pleasure to be here!