

The Einstein-Weyl Equations,

Scattering Maps, &

Holomorphic Disks

(Lecture VI)

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Stony Brook University

Autumn School on Holomorphic Disks
Schloss Rauschholzhausen, November 17, 2018

Joint work with

Lionel Mason
Oxford University

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The Einstein-Weyl Equations, Scattering Maps,
and Holomorphic Disks,
Math. Res. Lett. 16 (2009) 291–301.

Weyl's 1918 gauge theory

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$$\nabla_v g \propto g \quad \forall v$$

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for some 1-form α .

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where $\nu = d \log u$.

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where $n = \dim M$.

Hermann Weyl



$$\mathfrak{B} = (\mathfrak{G} + \alpha \mathbf{I}) + \frac{\varepsilon^2}{4} V \bar{g} \{1 - 3 (\varphi_i \varphi^i)\},$$

$$\Gamma_{ik}^r = \left\{ \begin{matrix} ik \\ r \end{matrix} \right\} + \frac{1}{2} \varepsilon^2 (\delta_i^r \varphi_k + \delta_k^r \varphi_i - g_{ik} \varphi^r).$$

Unter Vernachlässigung der winzigen kosmologischen Terme erhalten wir hier also genau die klassische Maxwell-Einsteinsche Theorie der Elektrizität und Gravitation. Um Übereinstimmung mit den in § 34 verwendeten

Ricci tensor

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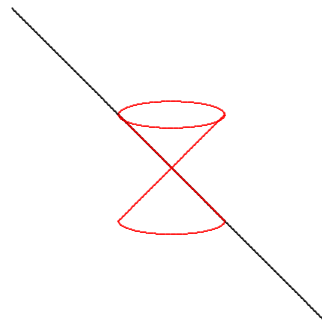
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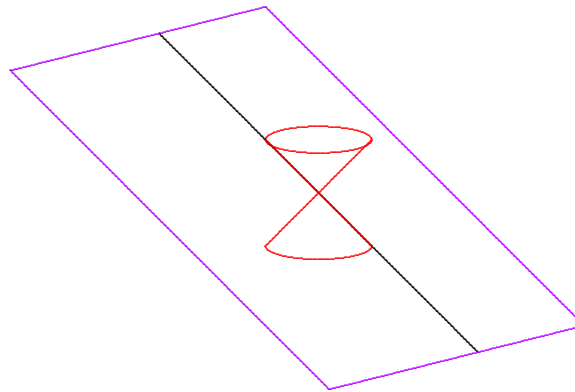
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THÉORÈME. — Les espaces de Weyl à trois dimensions qui admettent ∞^2 plans isotropes dépendent essentiellement de quatre fonctions arbitraires de deux arguments.

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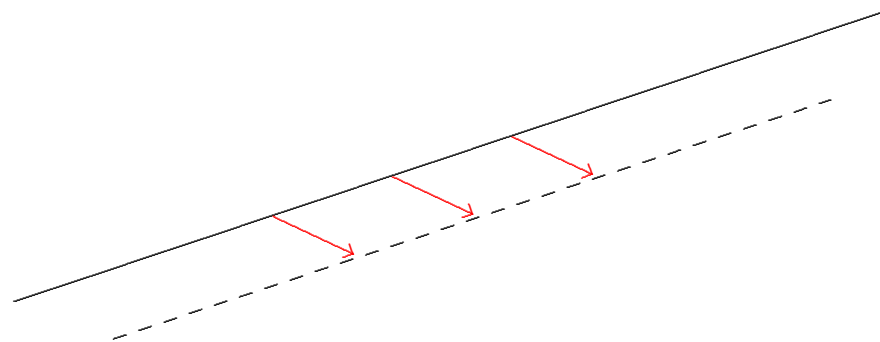
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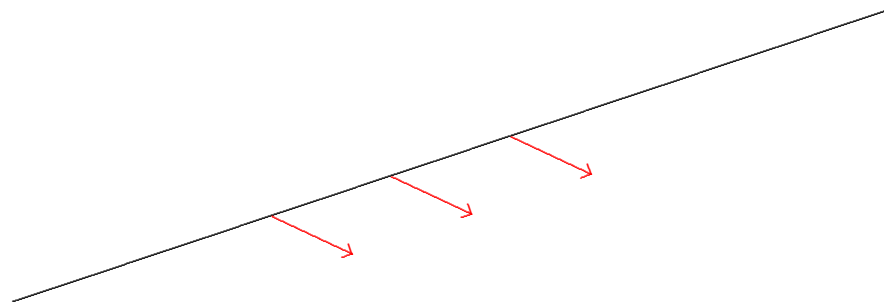
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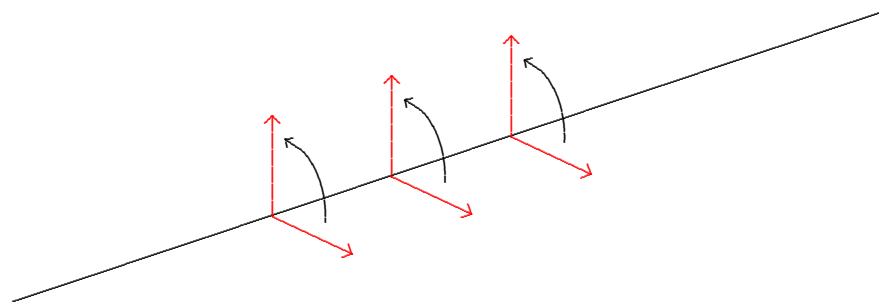
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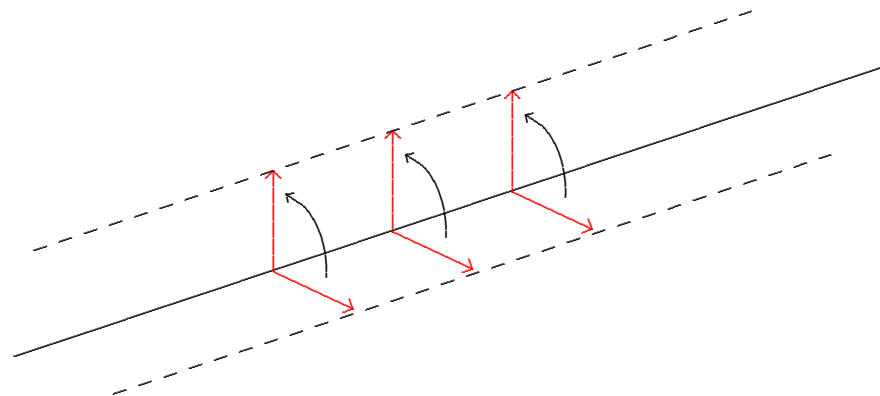
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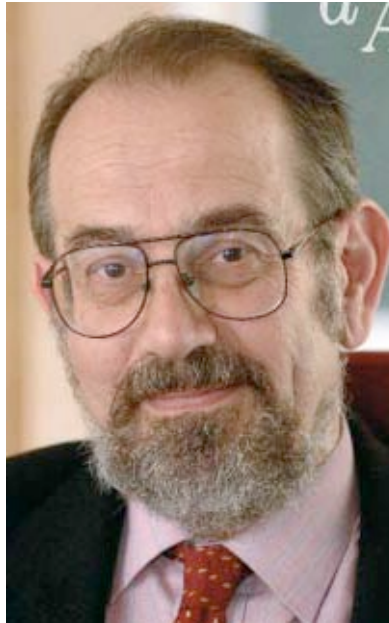
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Nigel Hitchin



and so if

$$R(U \times V, U)U = U \times R(V, U)U, \quad (2.2)$$

then we can define a linear map

$$J(V) = U \times V \quad (2.3)$$

which satisfies

$$J^2(V) = U \times (U \times V) = (U, V)U - (U, U)V = -V$$

We thus have a real complex surface G with a family of real lines of self-intersection number 2. It can be shown that any such surface may be obtained by the above geodesic construction, but using a Weyl structure rather than a Riemannian structure. The integrability condition (2.2) is then the analogue of Einstein's equations ($R_{(ij)} = \Lambda g_{ij}$) for the Weyl structure (see [10]). This is the

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Theorem. *There is a natural one-to-one correspondence between*

- *smooth, space-time-oriented, conformally compact, globally hyperbolic Lorentzian Einstein-Weyl 3-manifolds $(M, [g], \nabla)$; and*
- *orientation-reversing diffeomorphisms*

$$\psi : \mathbb{CP}_1 \rightarrow \mathbb{CP}_1.$$

space-time oriented

Conformal Lorentzian n -manifold $(M, [g])$ called
space-time oriented

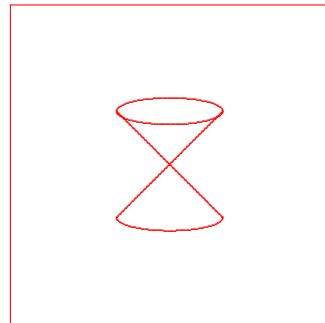
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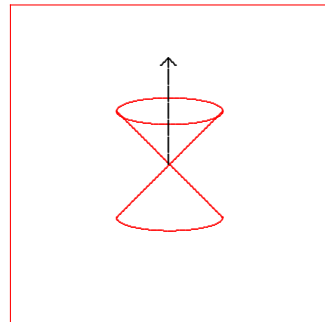
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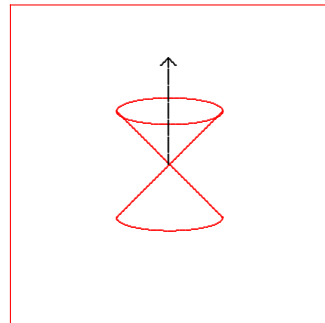
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$\implies M$ also oriented, in usual sense.

globally hyperbolic

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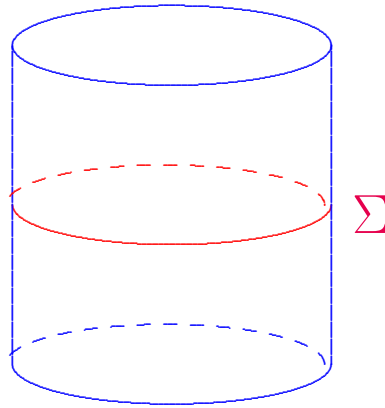
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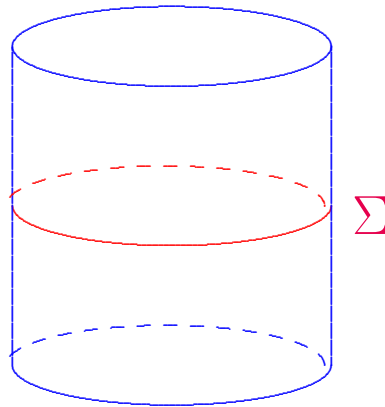
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$$\implies M \approx \Sigma \times \mathbb{R}$$

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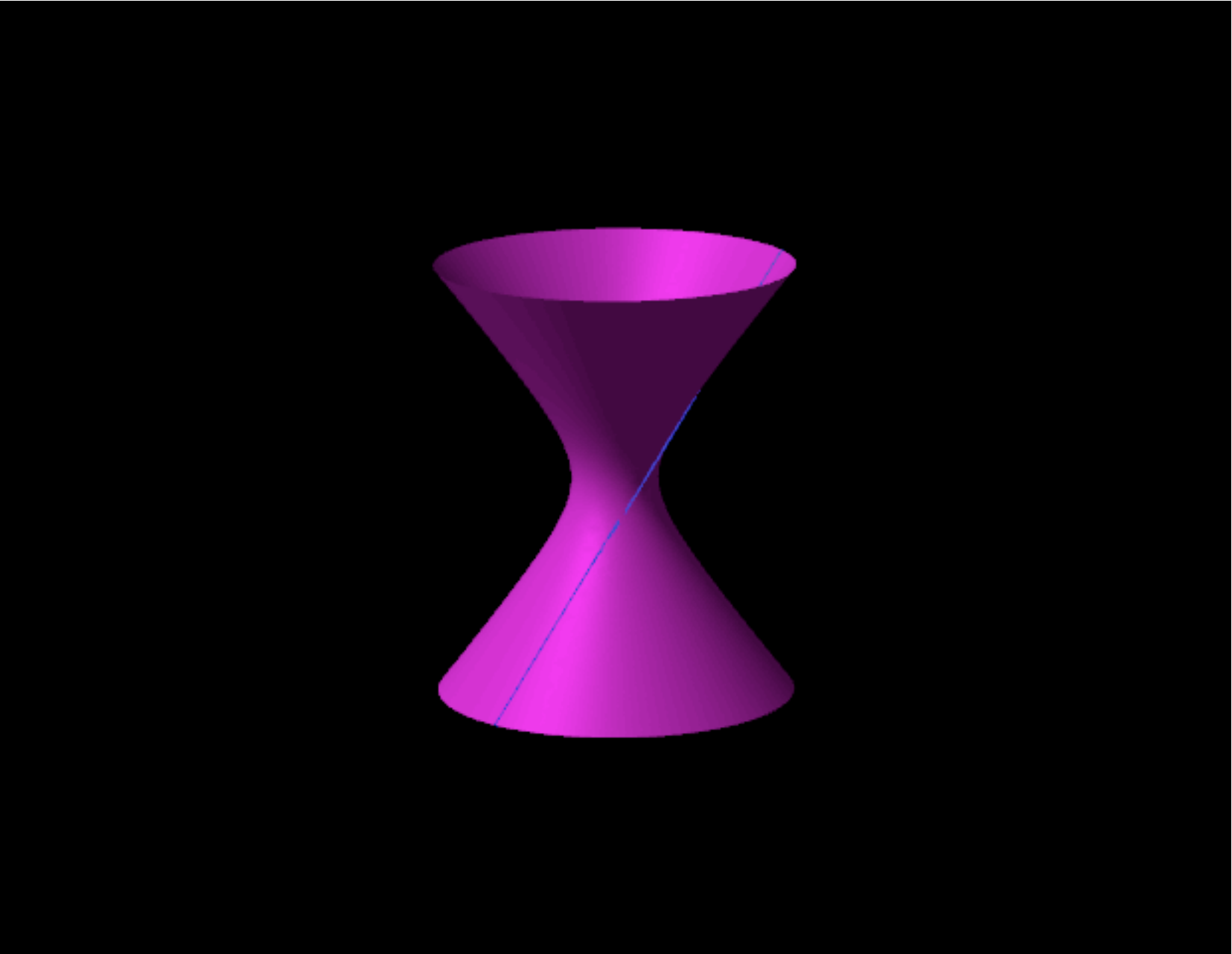
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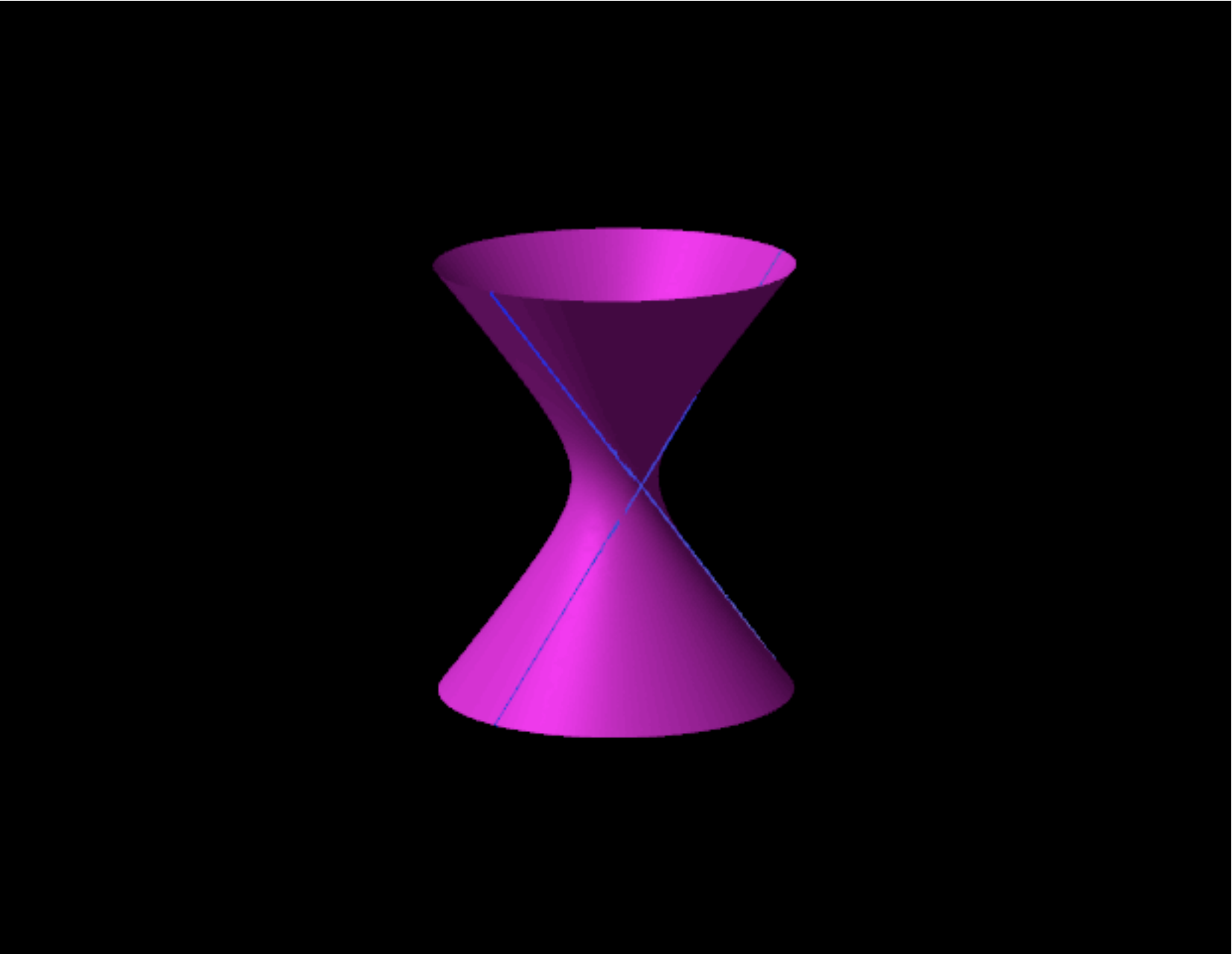
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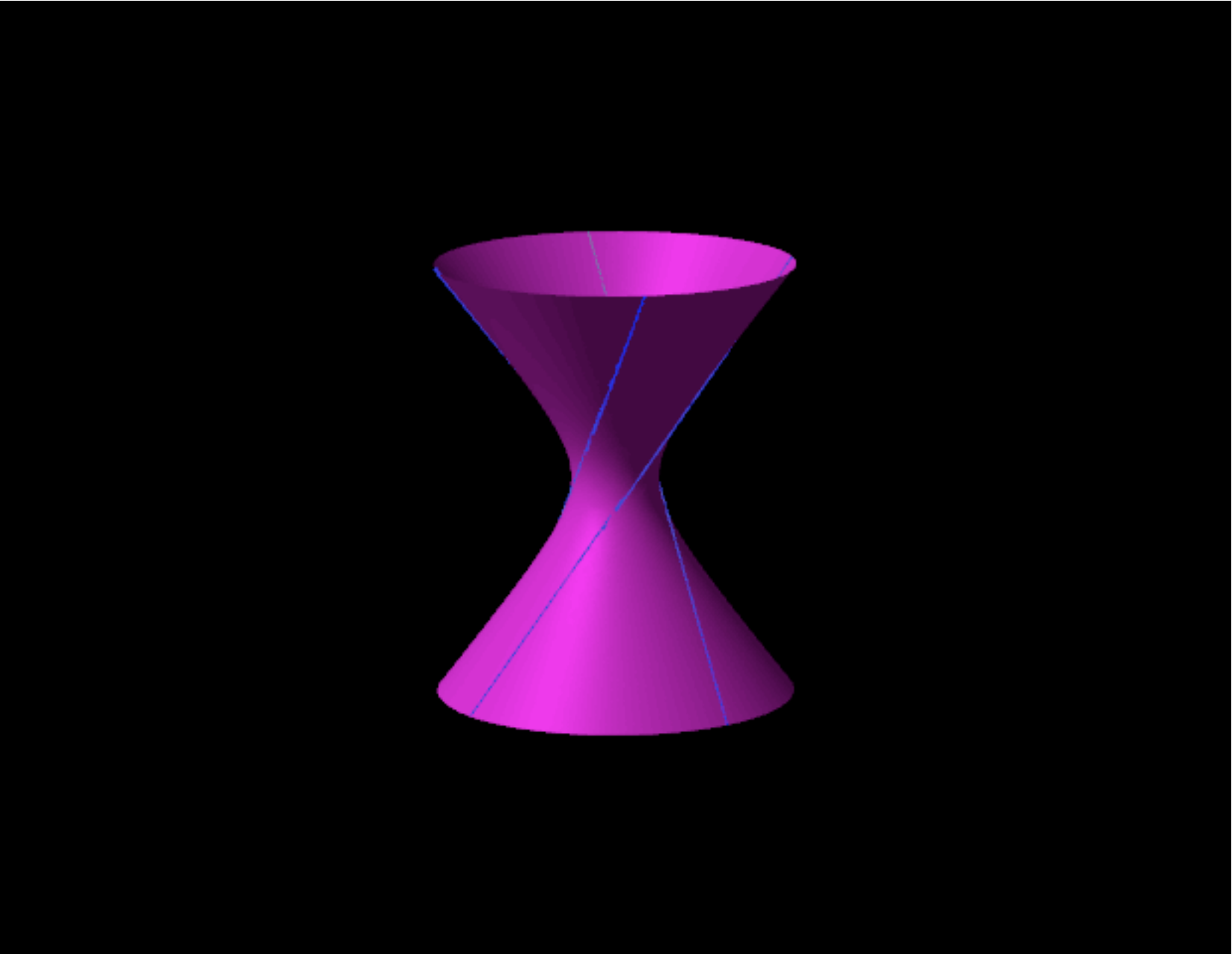
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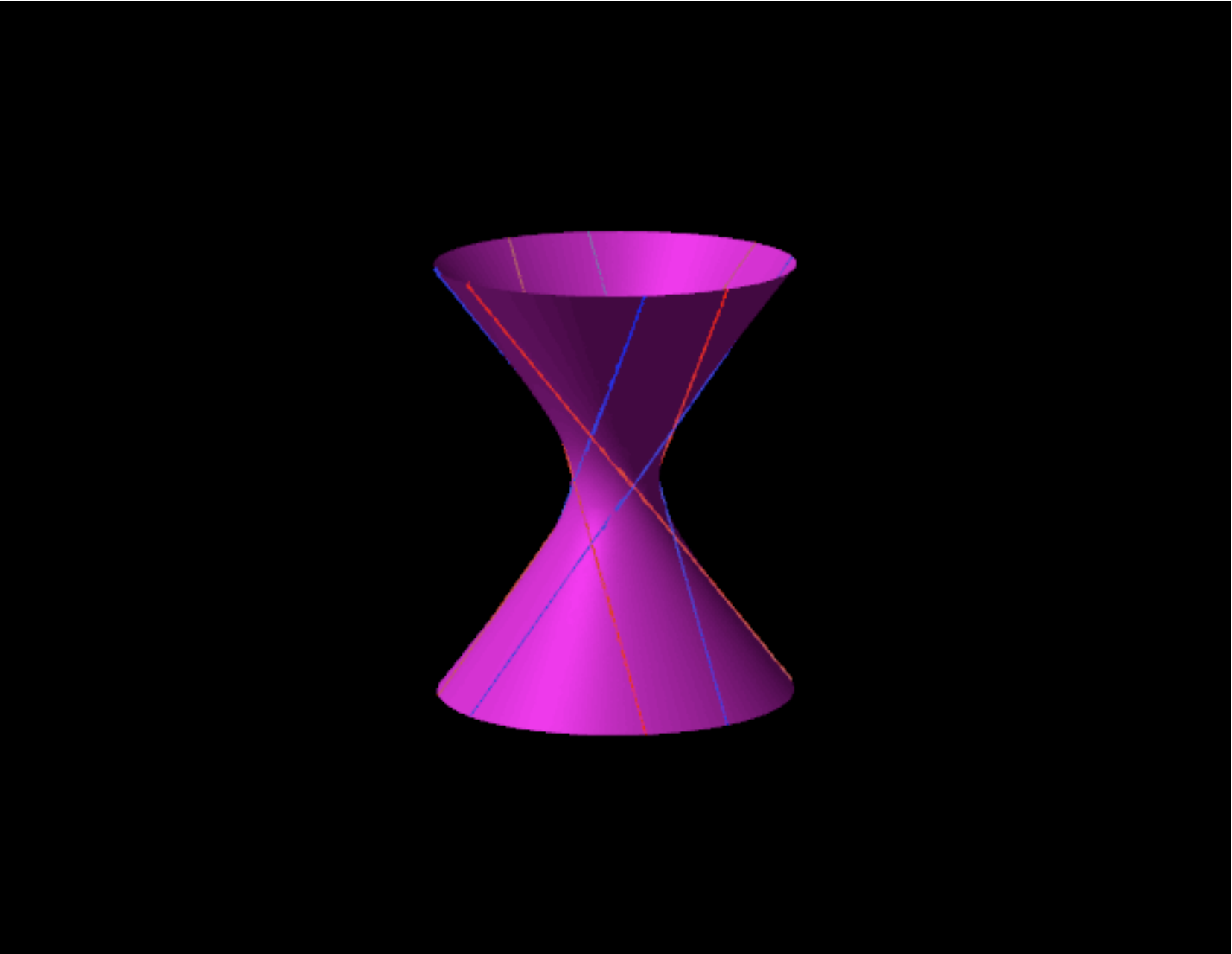
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$$M = SL(2, \mathbb{C}) / SL(2, \mathbb{R})$$









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where h = standard metric on S^2 .

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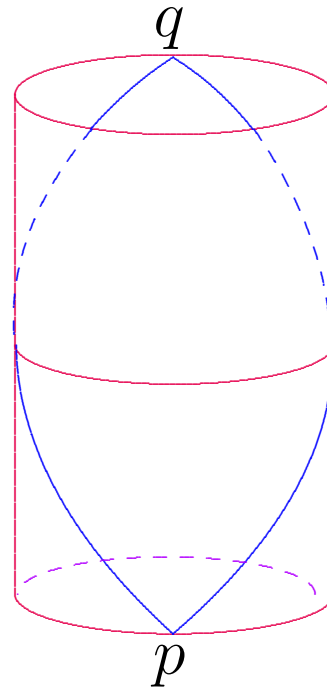
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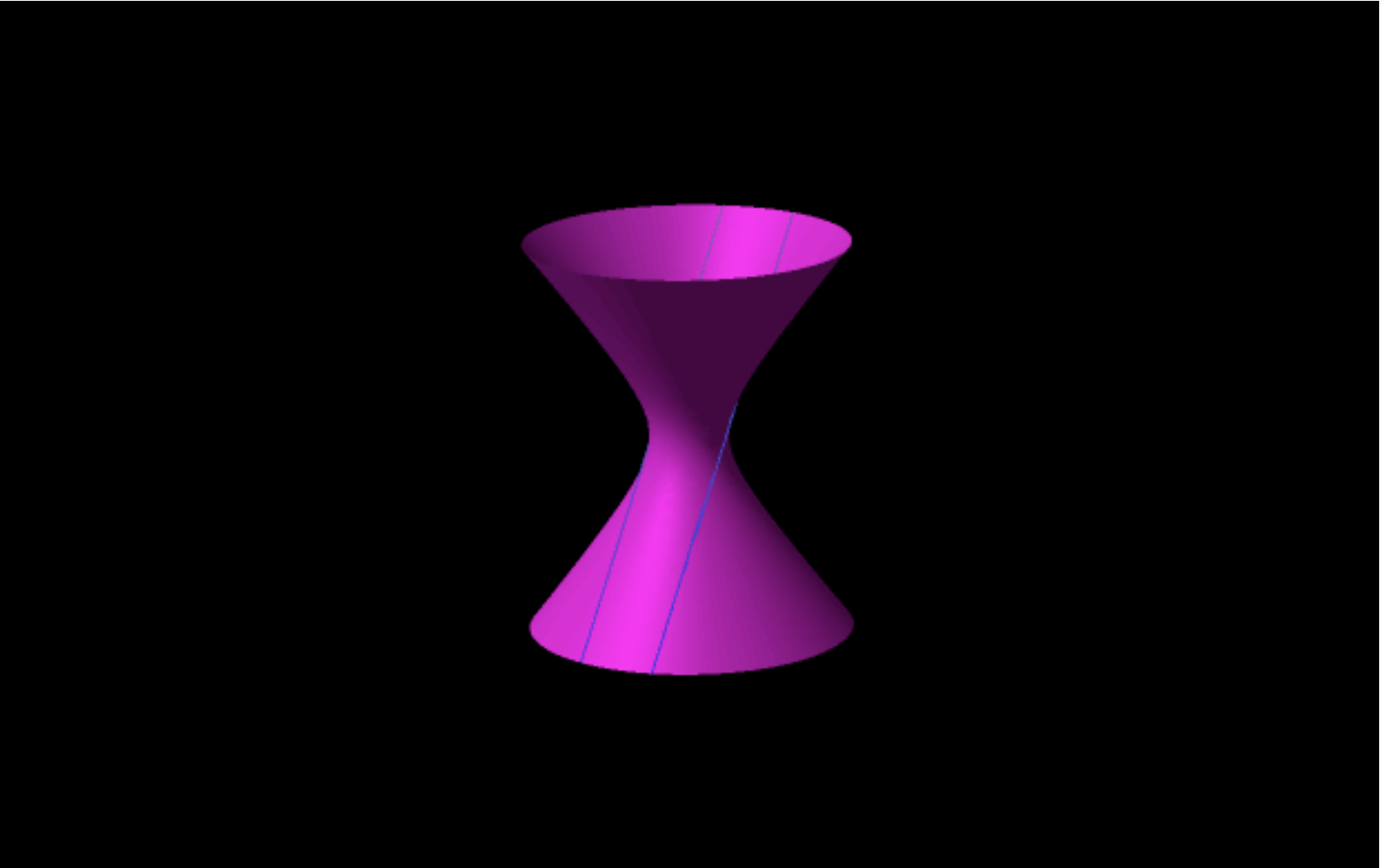
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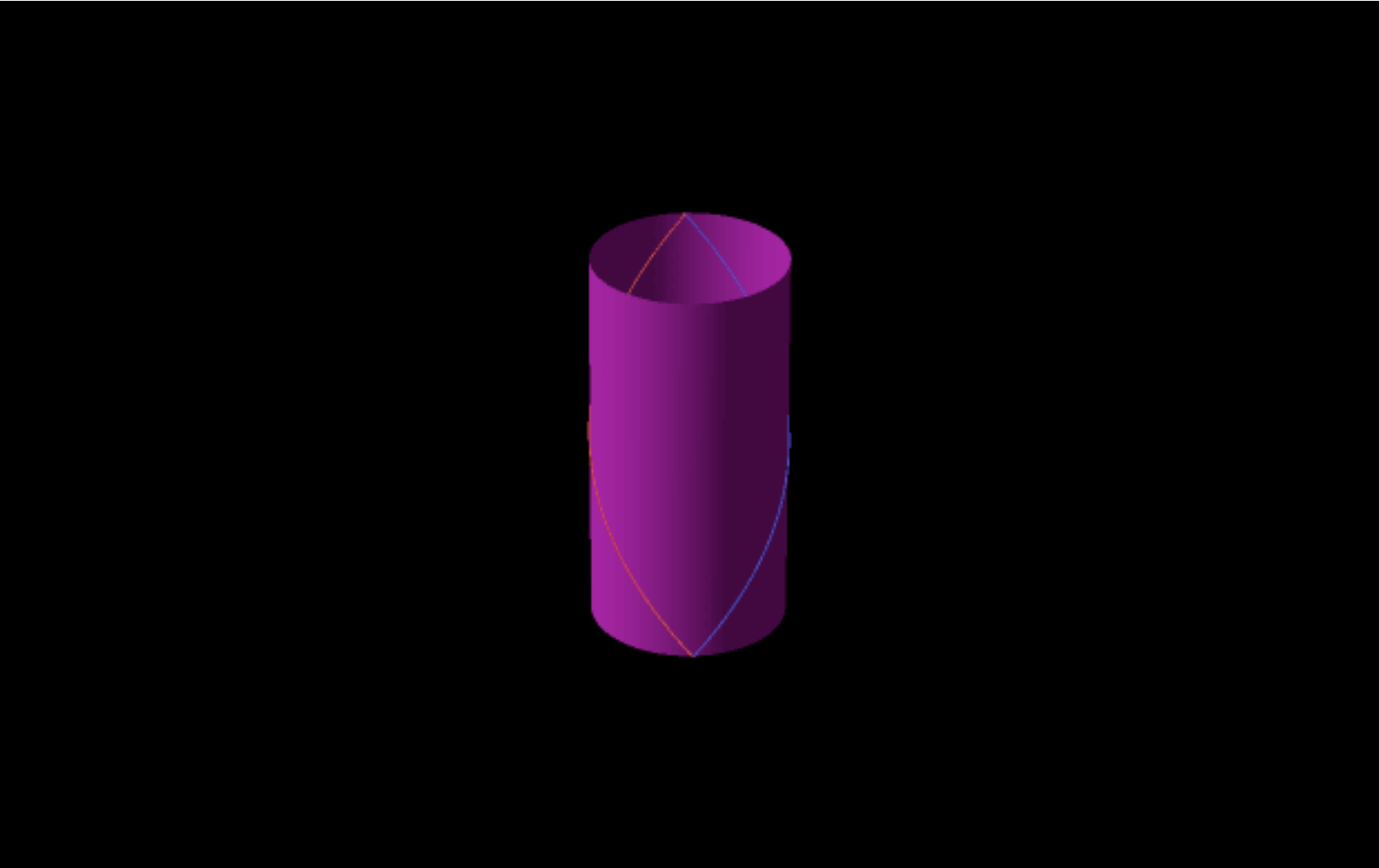
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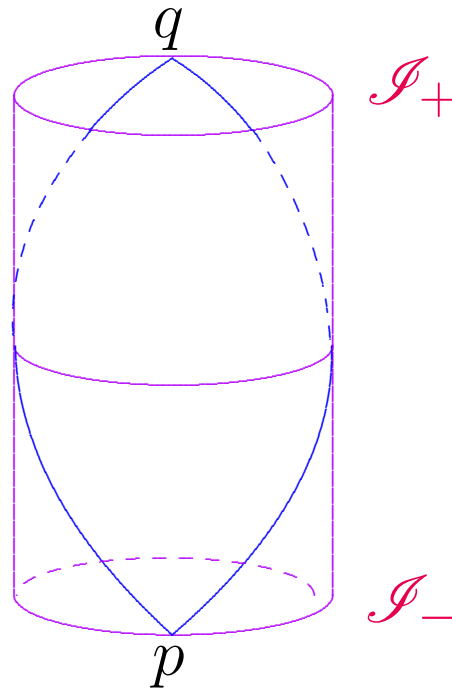


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$$\psi_1, \psi_2 : \mathbb{CP}_1 \rightarrow \mathbb{CP}_1$$

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for Möbius transformations $\varphi, \phi \in PSL(2, \mathbb{C})$.

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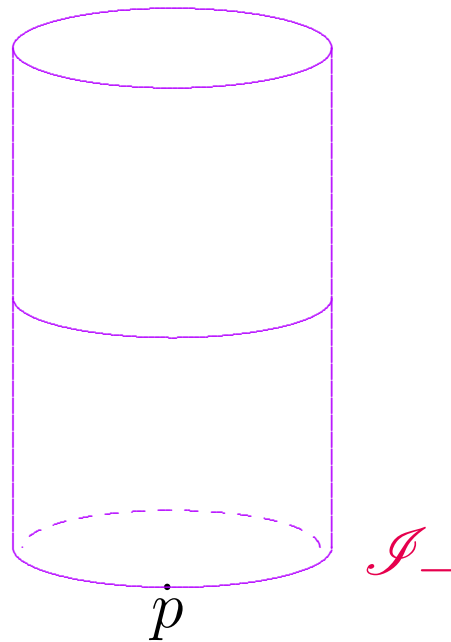
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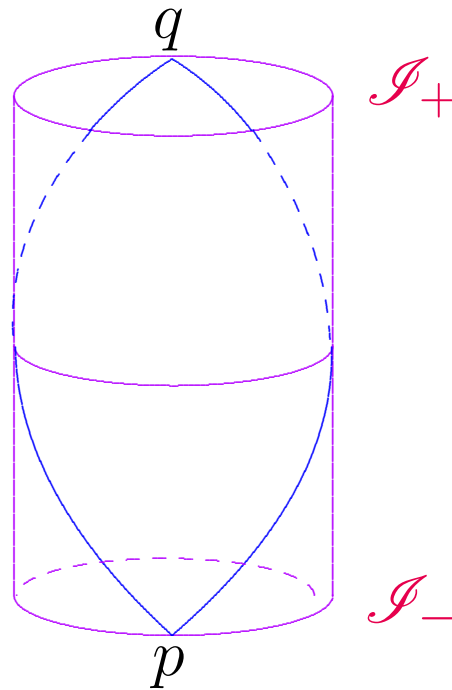
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to Einstein-Weyl $(M^3, [g], \nabla)$ satisfying hypotheses.

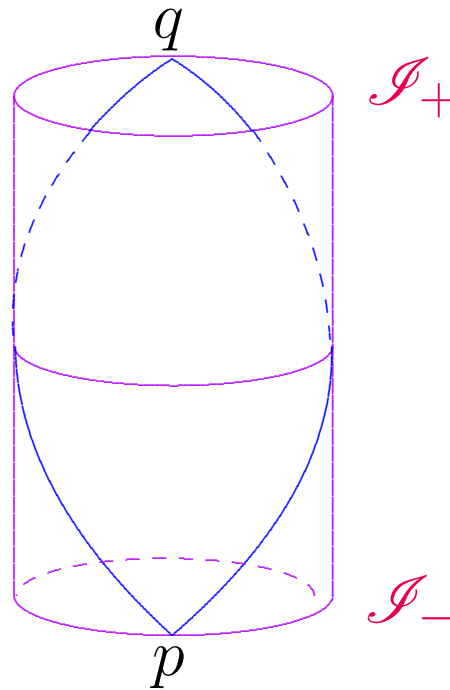
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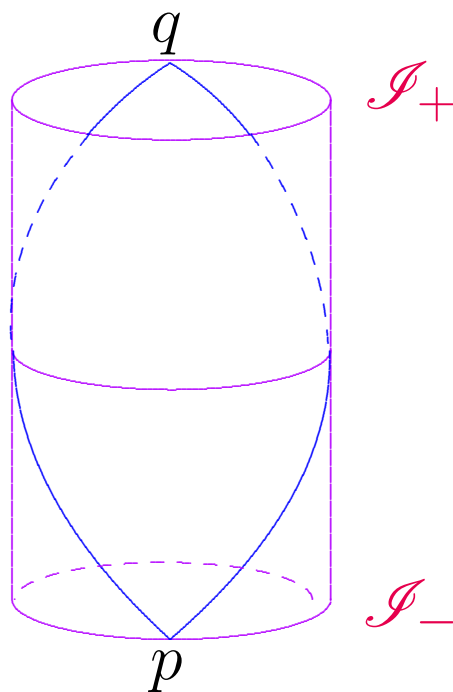
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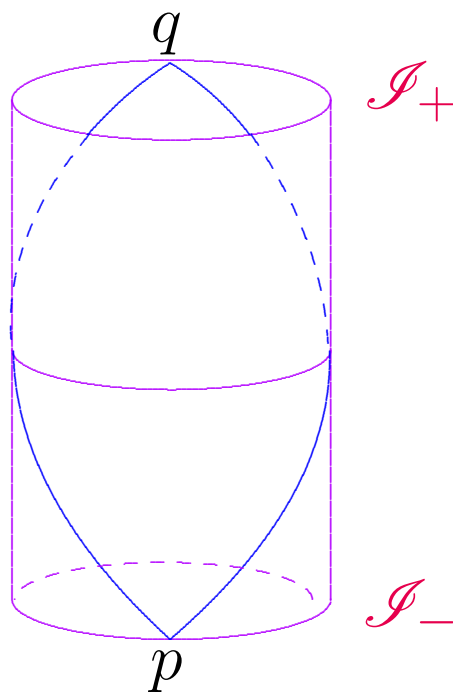
Lemma. For an Einstein-Weyl manifold as above, let $p \in \mathcal{I}_-$ be any point of past infinity. Then all the null geodesics emanating from p refocus at a unique point $q \in \mathcal{I}_+$. Moreover, $\mathcal{I}_\pm \approx S^2$.



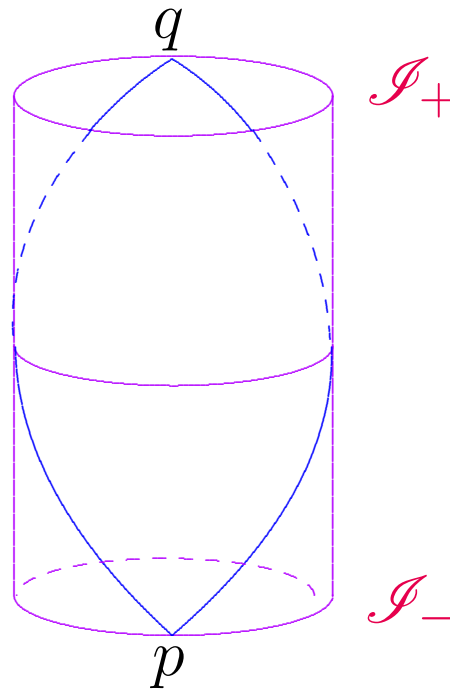
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twistor disk construction

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Graph of orientation-reversing diffeomorphism

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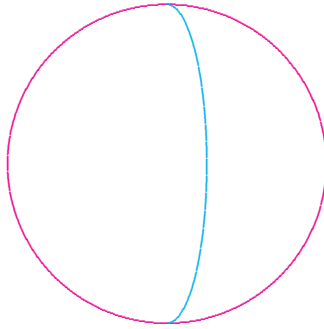
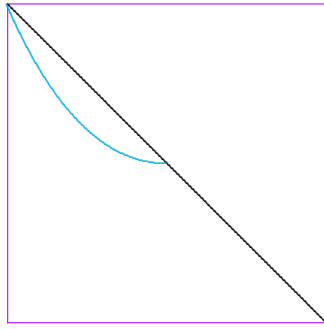
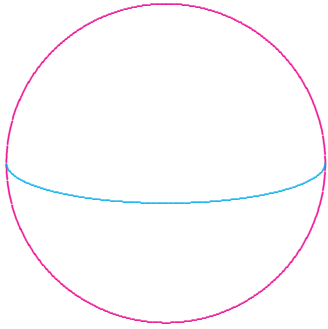
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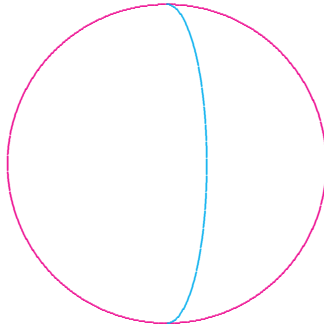
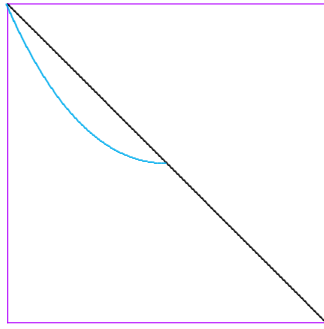
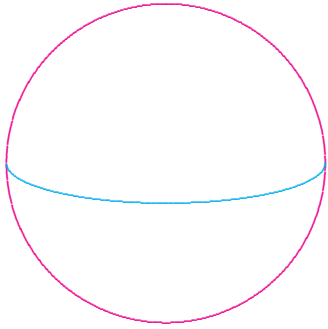
Strategy: construct 3-manifold $M = M_\psi$
as moduli space of holomorphic disks D
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When ψ is the antipodal map,
disks are explicitly given by

$$\zeta \longmapsto ([a\zeta + b : c\zeta + d], [-\bar{d}\zeta - \bar{c} : \bar{b}\zeta + \bar{a}])$$

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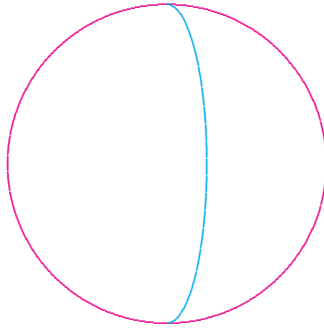
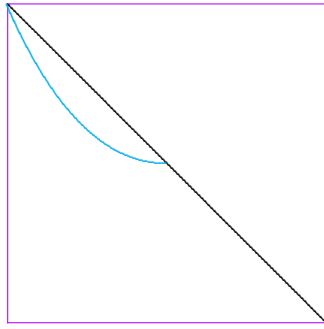
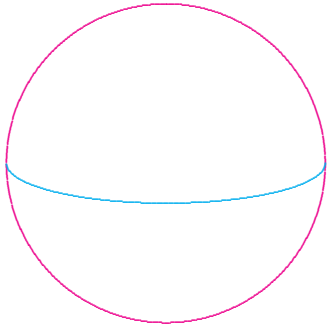
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Moduli space M of disks mod reparameterization:

de Sitter space $SL(2, \mathbb{C})/SL(2, \mathbb{R})$.

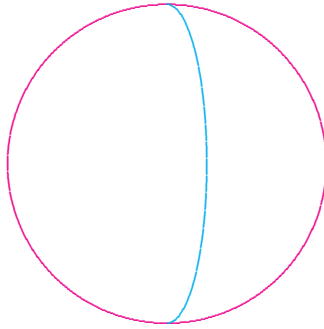
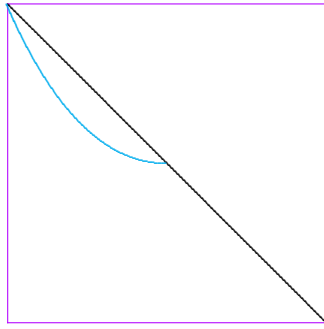
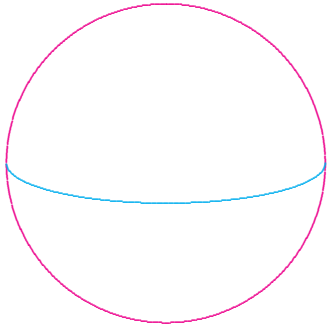


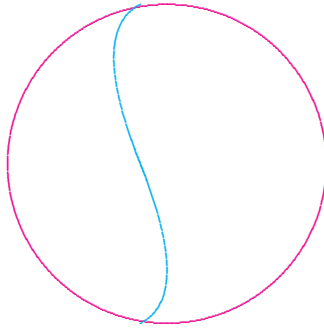
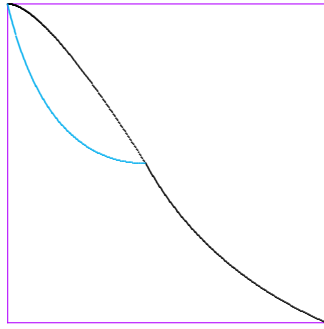
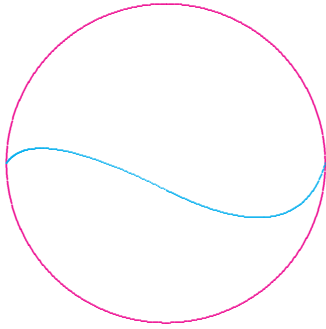
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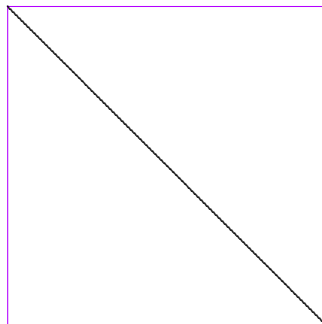
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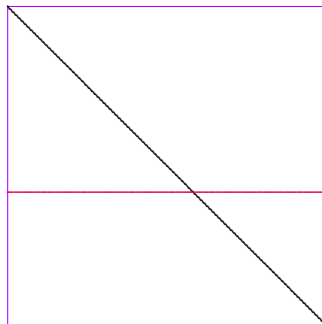
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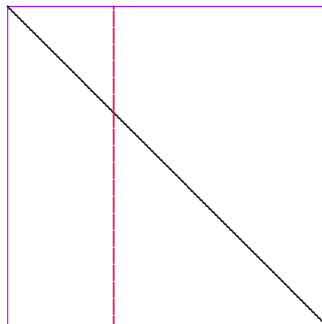
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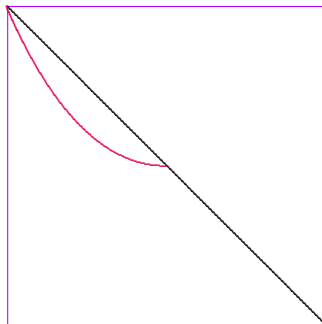
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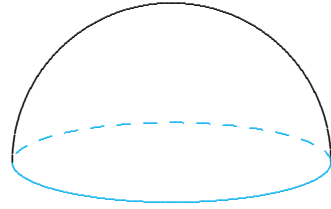
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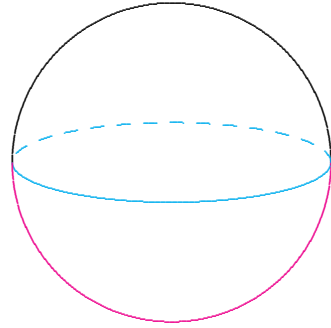
If $(\Sigma, \partial\Sigma) \rightarrow (Z, P)$ is any holomorphic curve with boundary representing \mathbf{a} , then Σ is either a holomorphic disk as above, or is a factor $\mathbb{C}\mathbb{P}_1$ of $Z = \mathbb{C}\mathbb{P}_1 \times \mathbb{C}\mathbb{P}_1$.

Proof. Consider abstract double $D \cup_{\partial D} \overline{D}$.

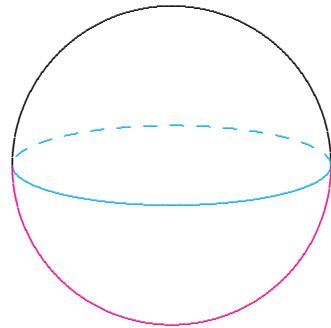
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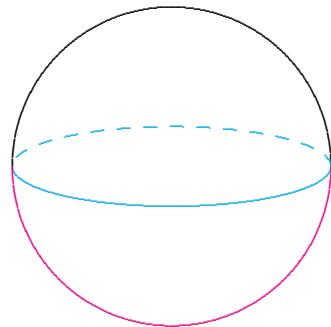
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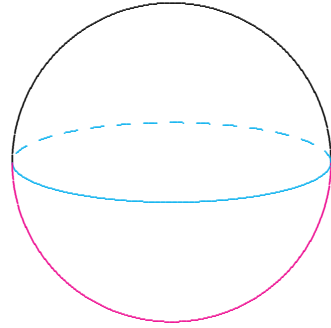
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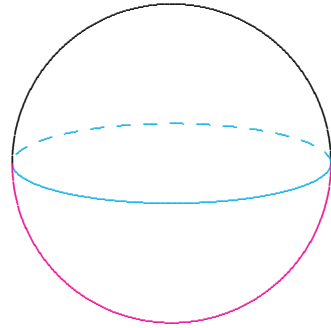
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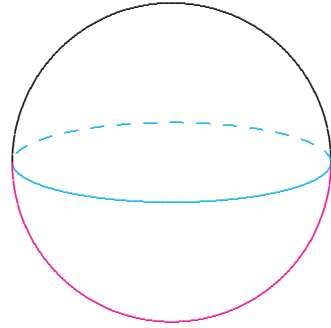
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So $\deg \Phi = 1$.

Proof. Consider abstract double $D \cup_{\partial D} \overline{D}$.



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Φ orientation-preserving; \implies homeomorphism.

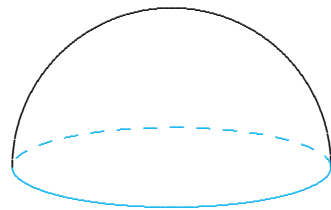
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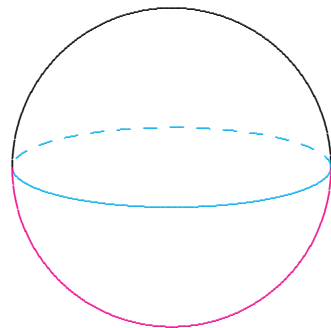
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Equals 2 in our case:

$$E \cong \mathcal{O}(2).$$

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$$\begin{aligned}h^1(\mathbb{C}\mathbb{P}_1, \mathcal{O}(2)) &= 0 \\h^0(\mathbb{C}\mathbb{P}_1, \mathcal{O}(2)) &= 3\end{aligned}$$

cf. Kodaira's Theorem
on deformation of complex submanifolds

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any sequence has convergent subsequence...

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Tricky point: disks can degenerate to factor $\mathbb{C}\mathbb{P}_1$.

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Since ψ is continuous deformation of antipodal,

Continuity method \Rightarrow each level set **non-empty!**

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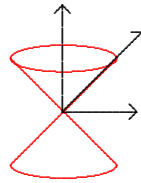
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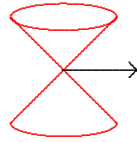
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\Rightarrow up to homothety $T_D M$ carries Lorentz metric, modelled on Killing form of $\mathfrak{sl}(2, \mathbb{R})$.

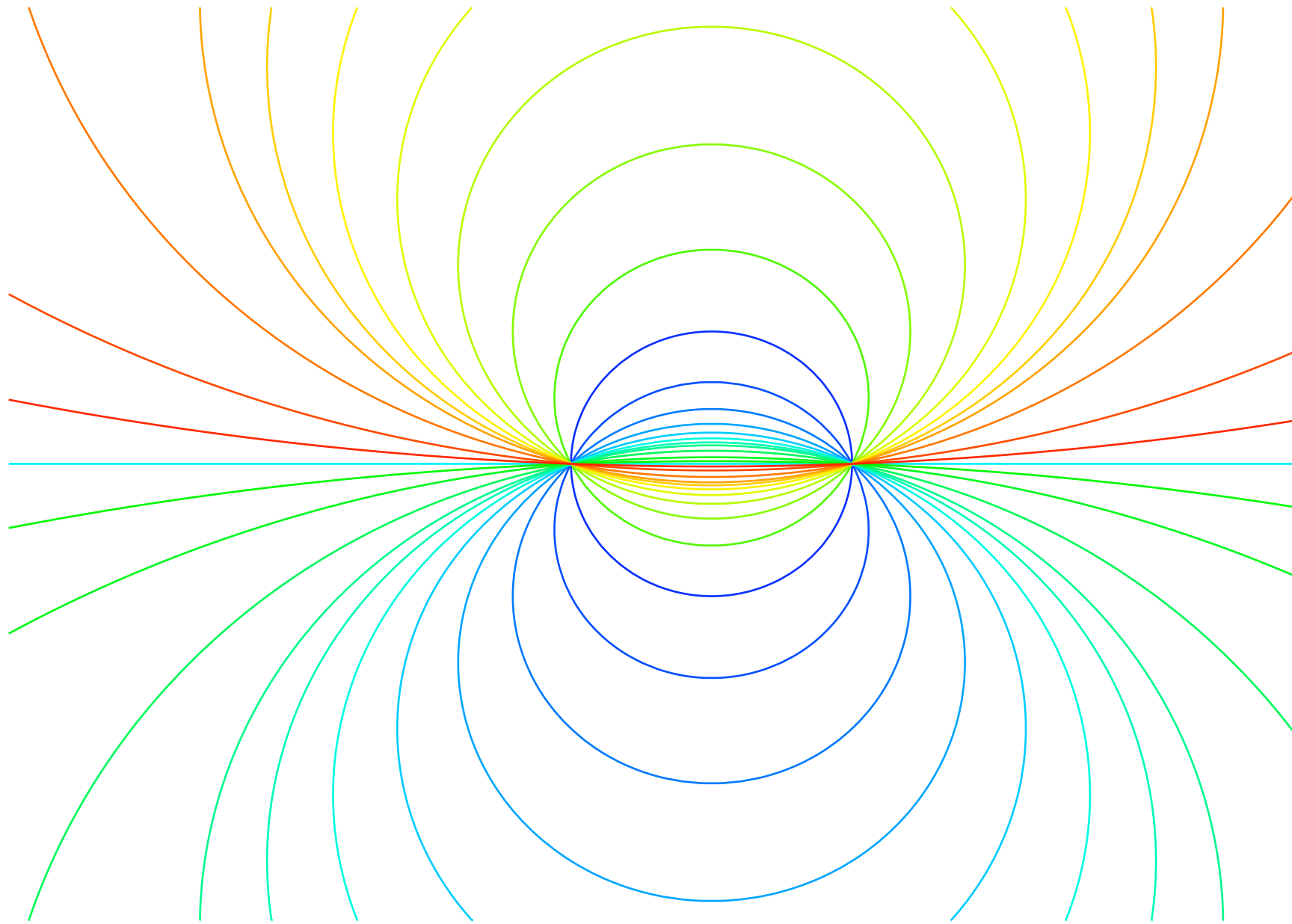
Trichotomy:

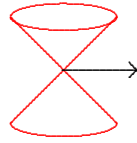
TM	$\mathfrak{sl}(2, \mathbb{R})$
space-like	hyperbolic
null	parabolic
time-like	elliptic



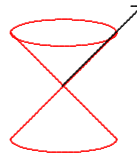


Space-like vector = infinitesimal variation with
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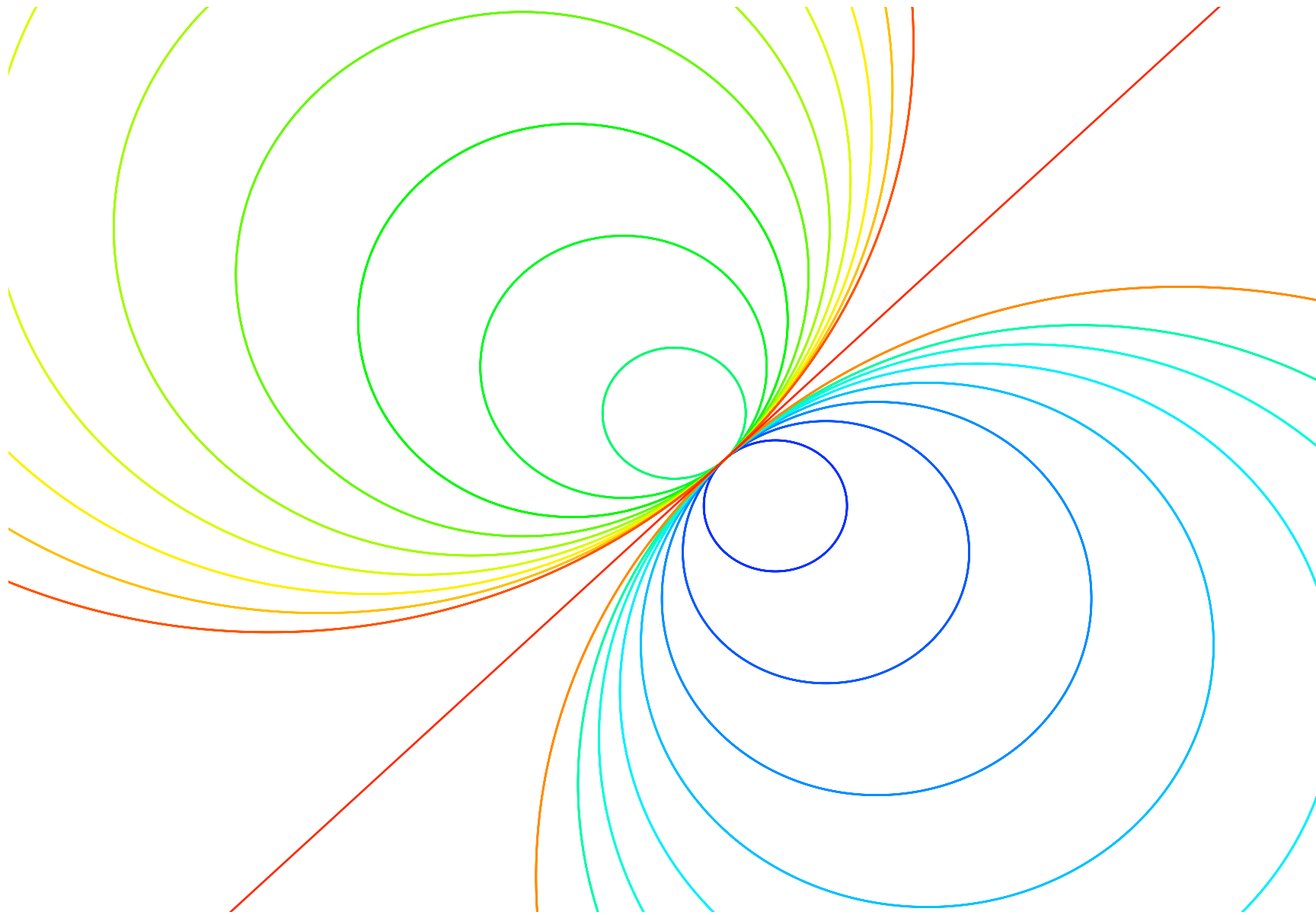


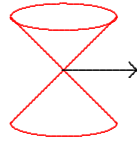


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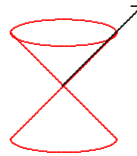


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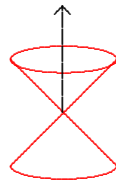




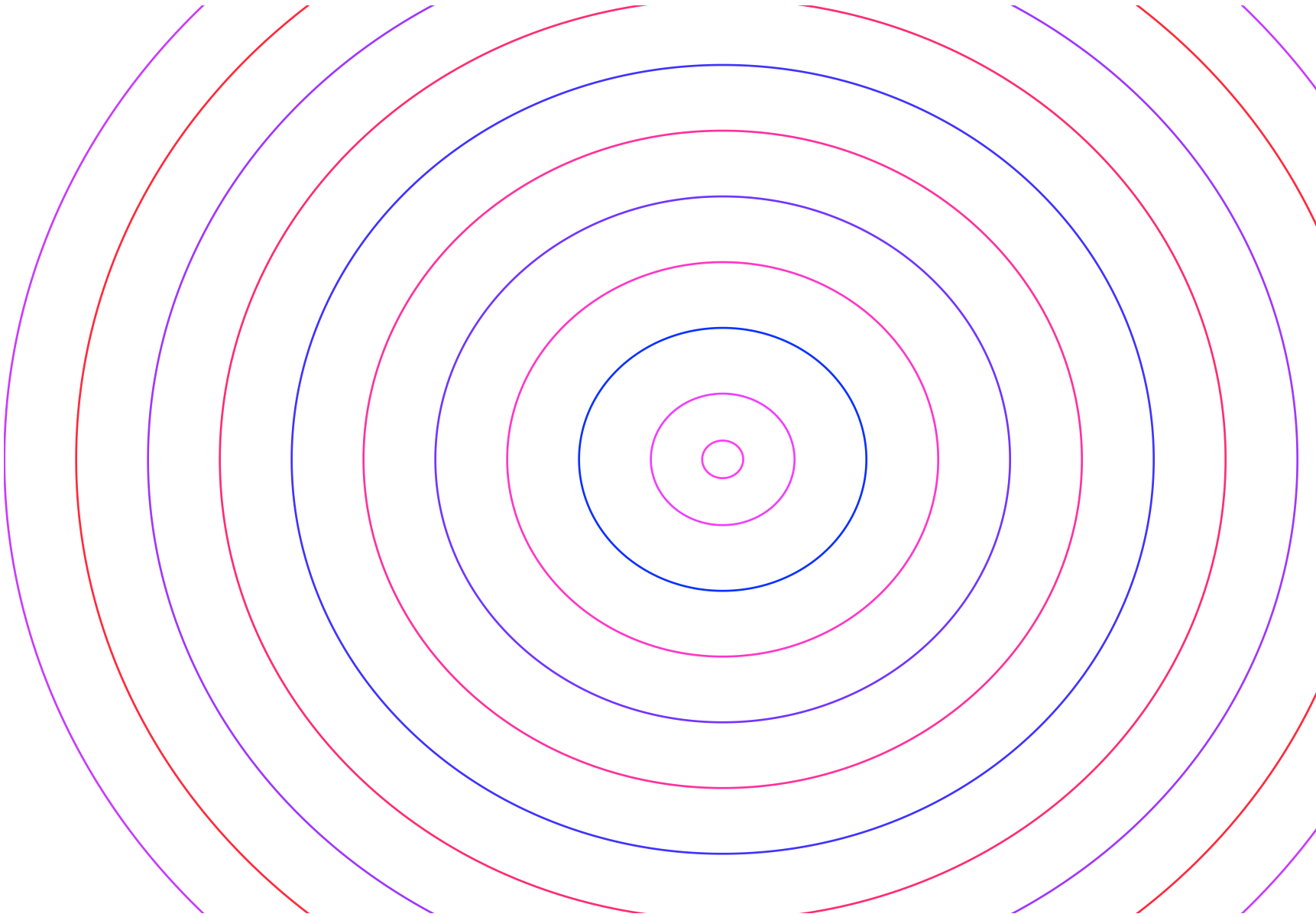
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By deformation: Cauchy surface topologically S^2 .

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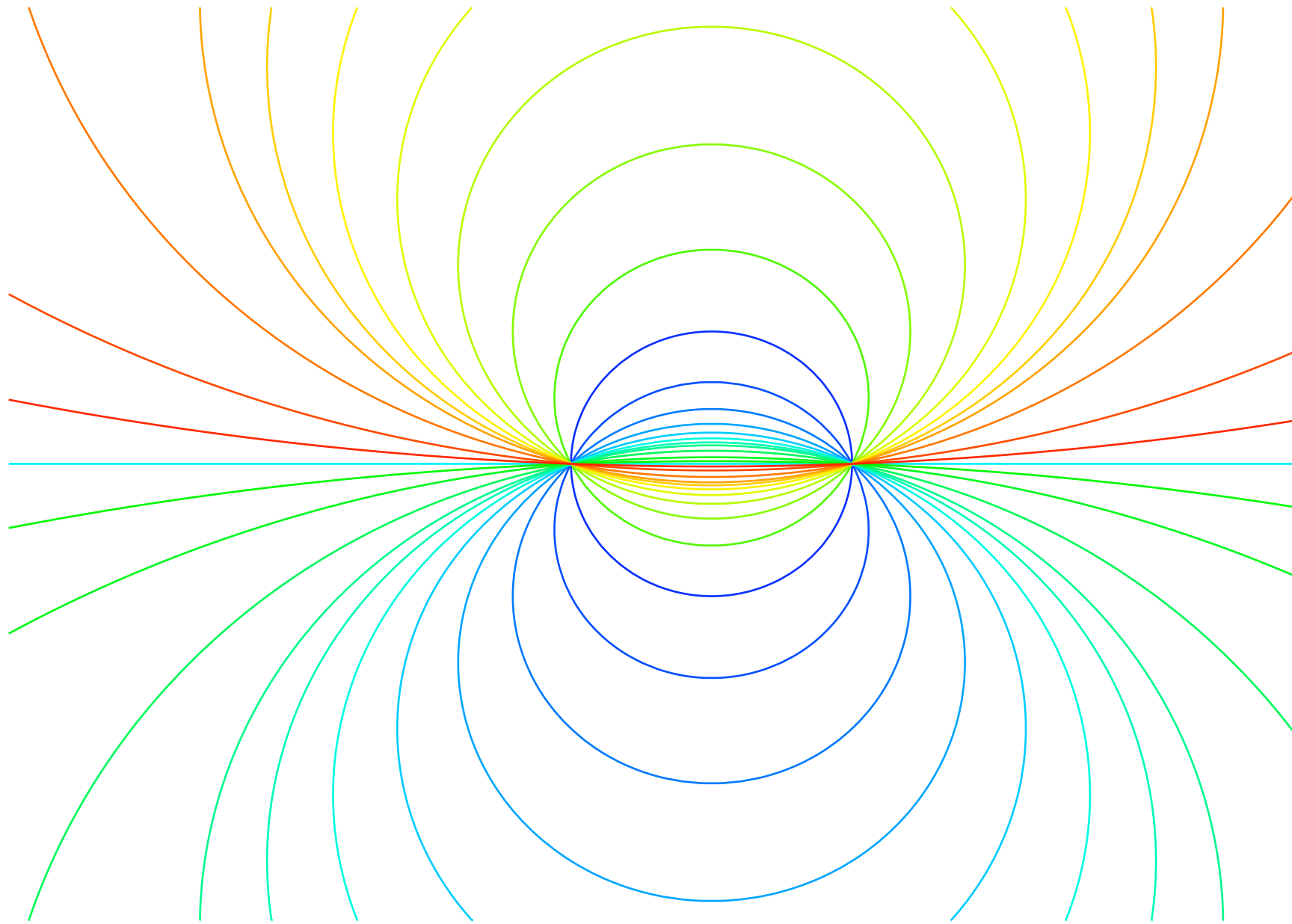
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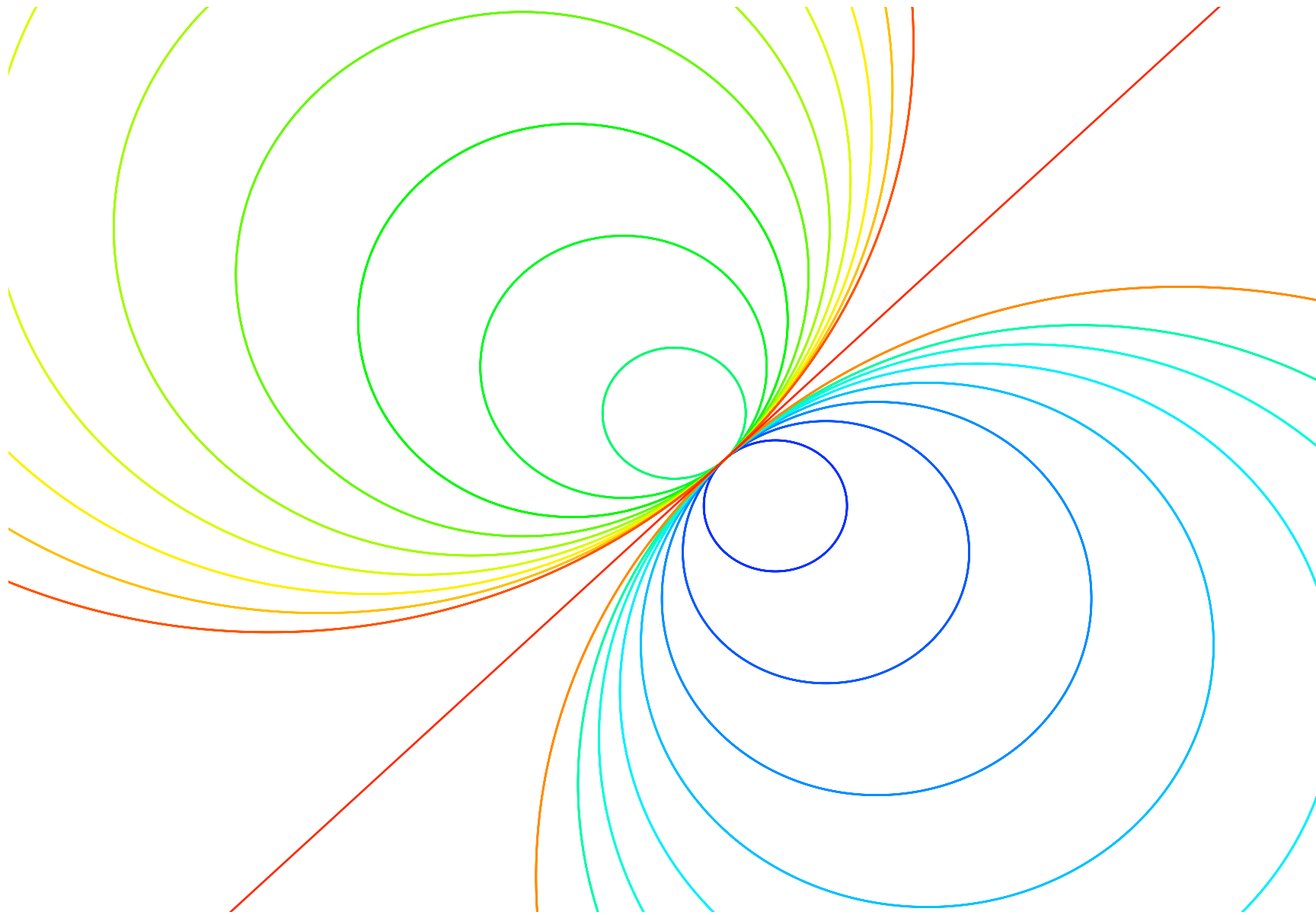
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Also gives direct proof of conformal compactness.

End, Part VI