

On

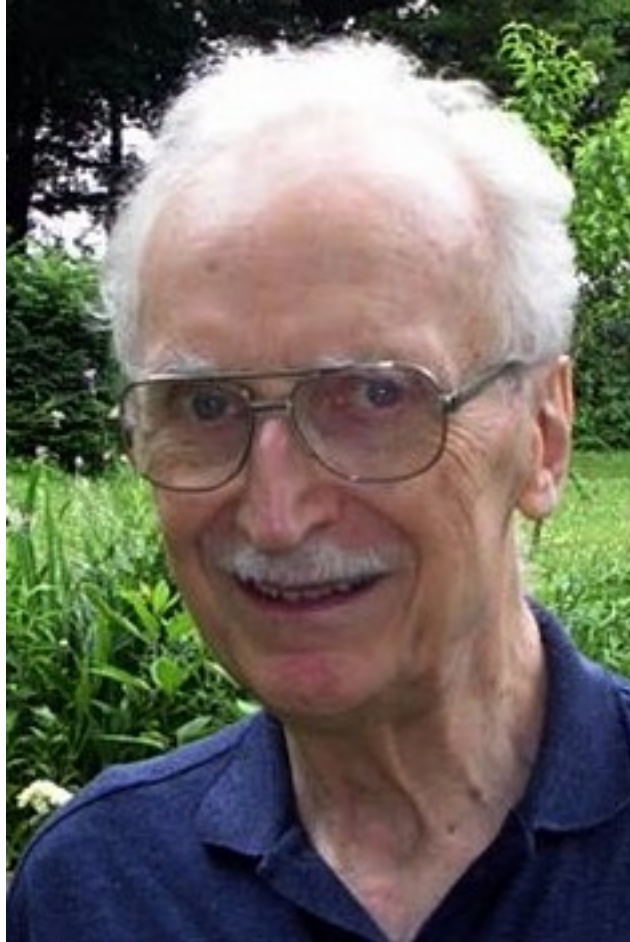
Hermitian, Einstein

4-Manifolds

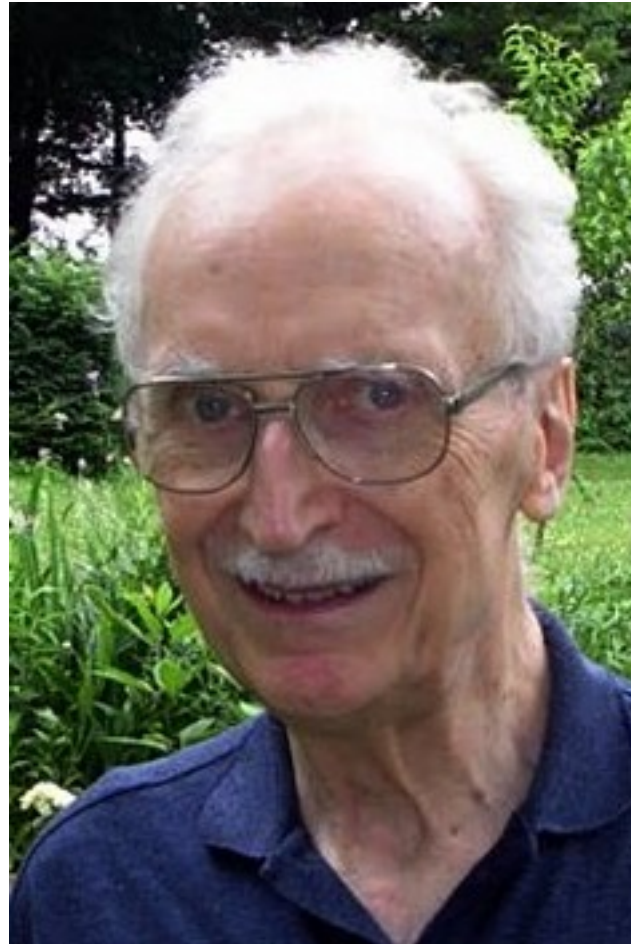
Claude LeBrun

Stony Brook University

For Eugenio Calabi



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who discovered the magic link between
Einstein manifolds and complex geometry.

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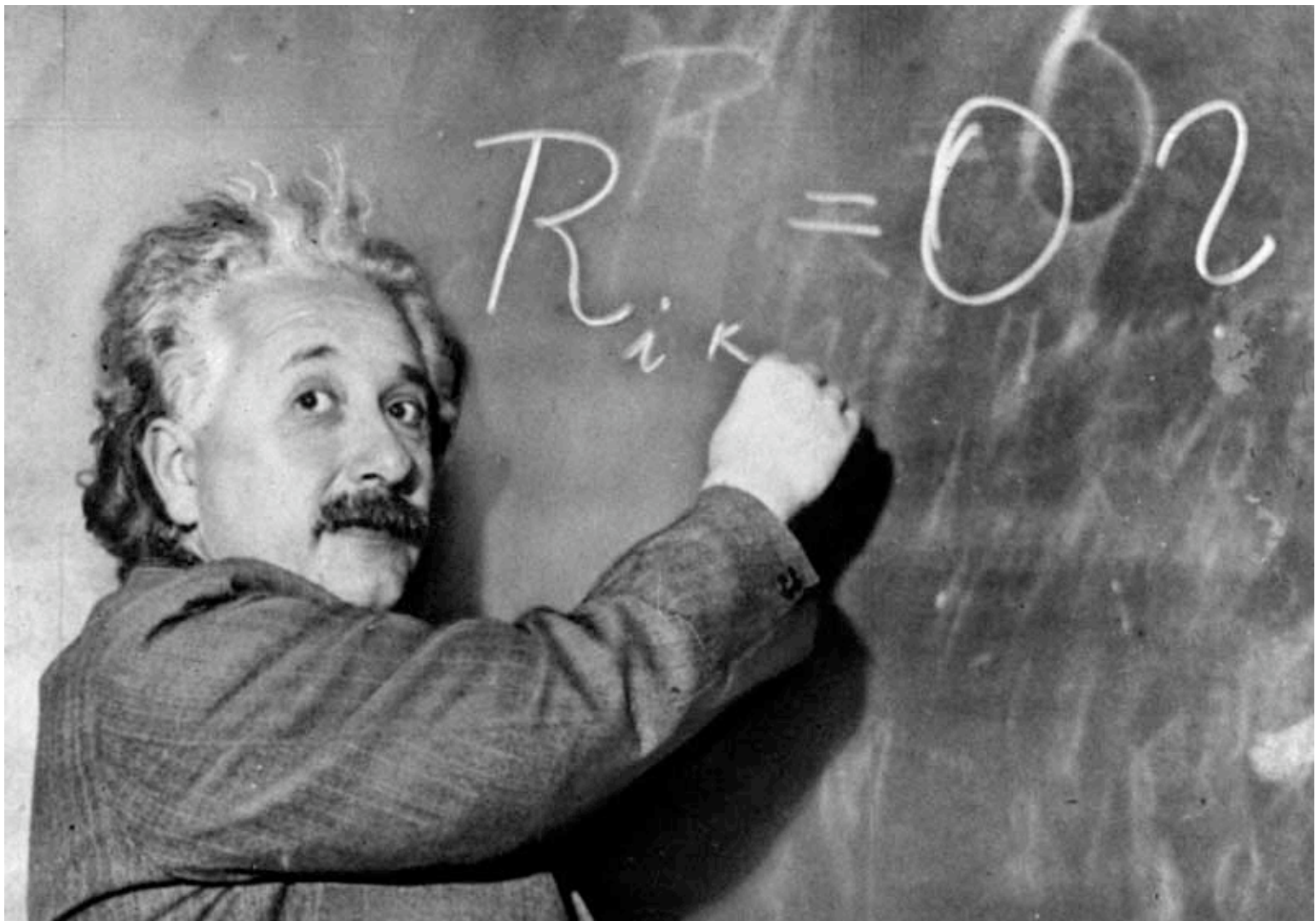
“... the greatest blunder of my life!”

— A. Einstein, to G. Gamow

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Has same sign as the *scalar curvature*

$$s = r_j^j = \mathcal{R}^{ij}_{ij}.$$

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where $[h_{j\bar{k}}]$ Hermitian matrix at each point.

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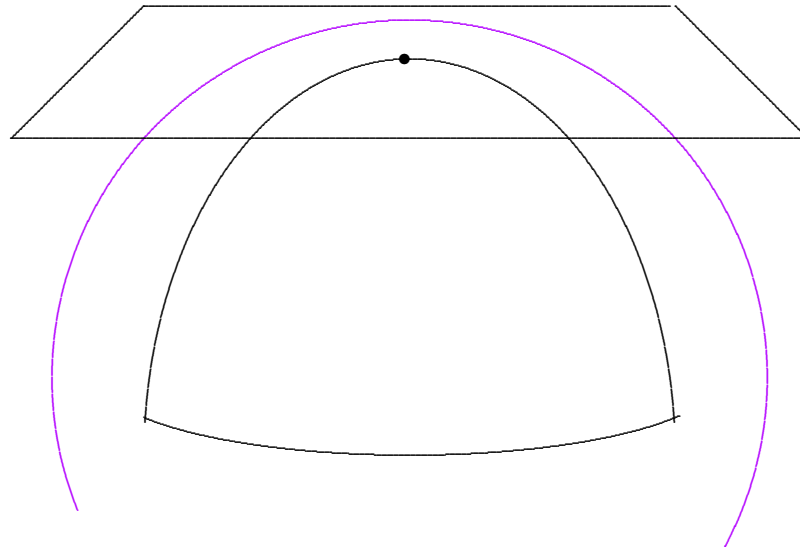
$[\omega] \in H^2(M, \mathbb{R})$ called the Kähler class.

$(M^n, g):$

holonomy

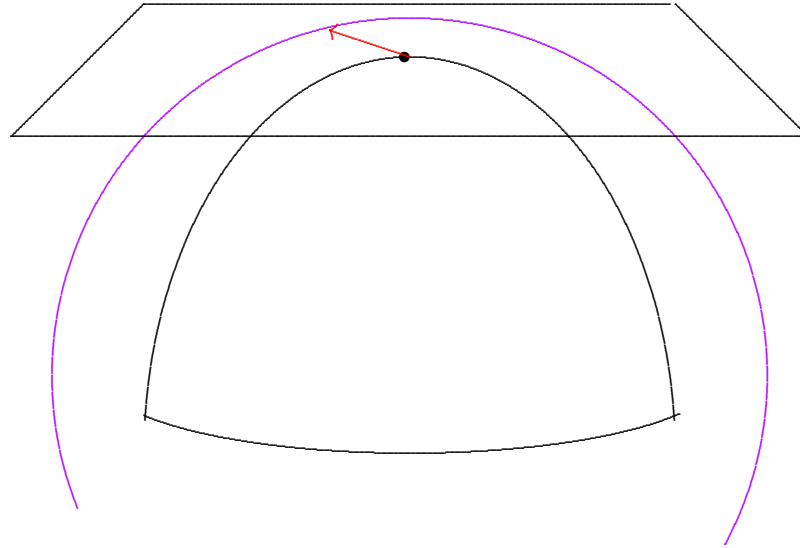
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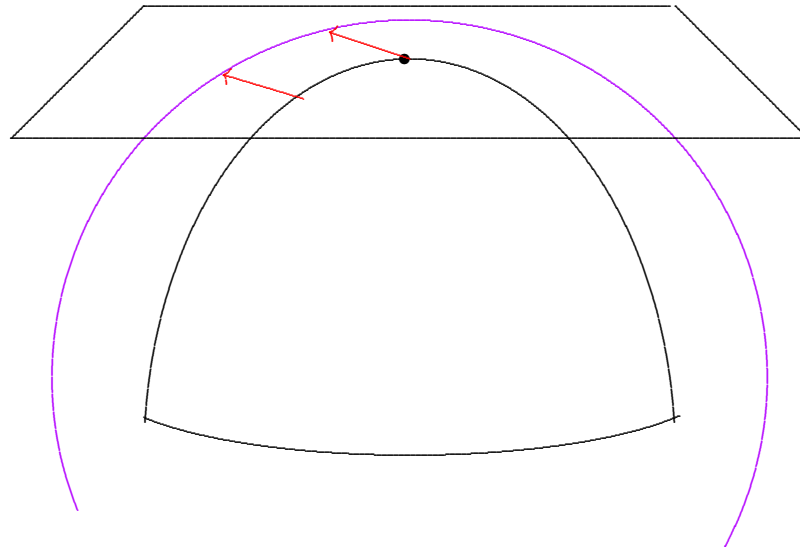
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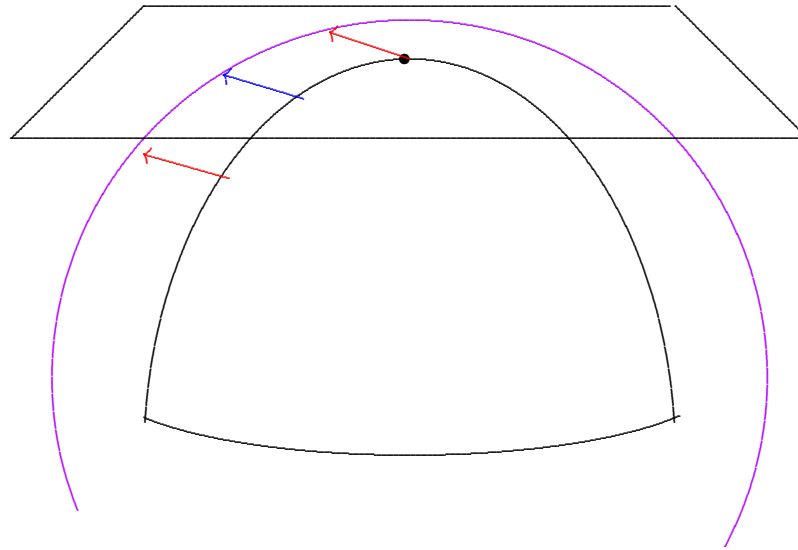
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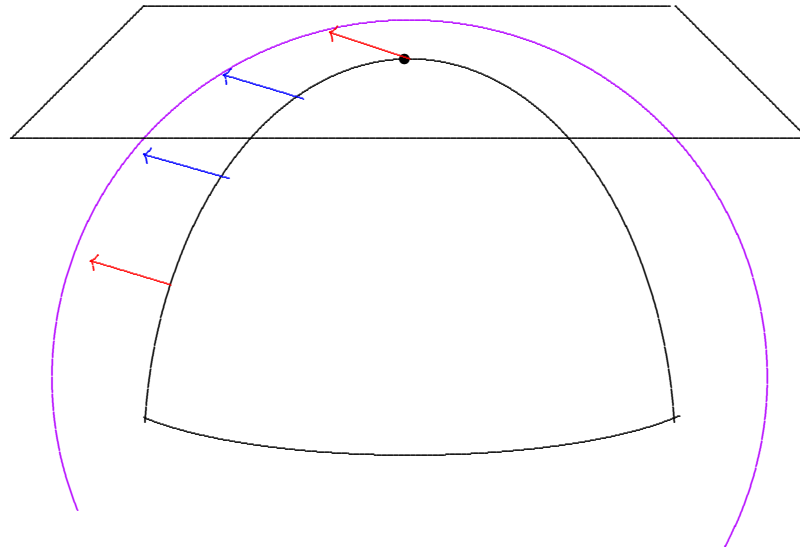
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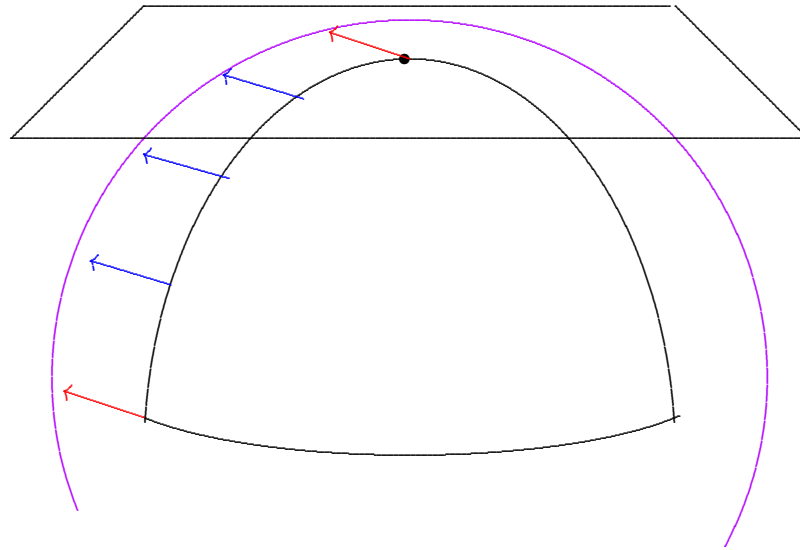
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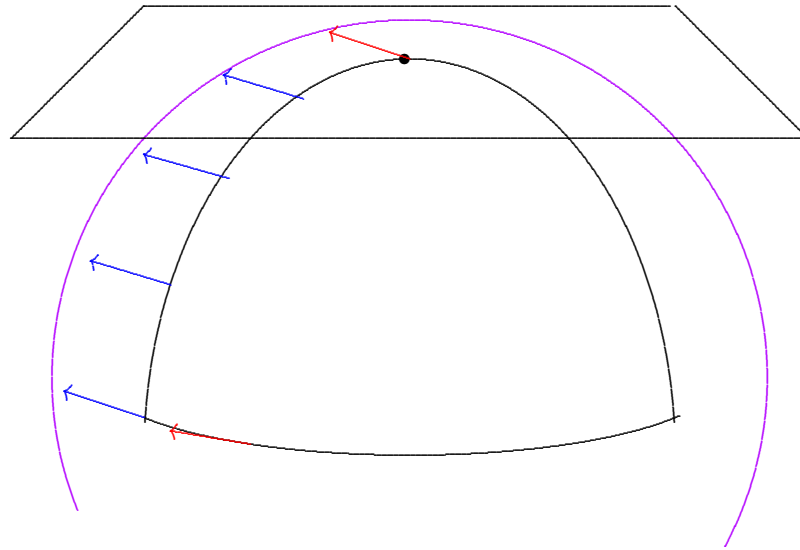
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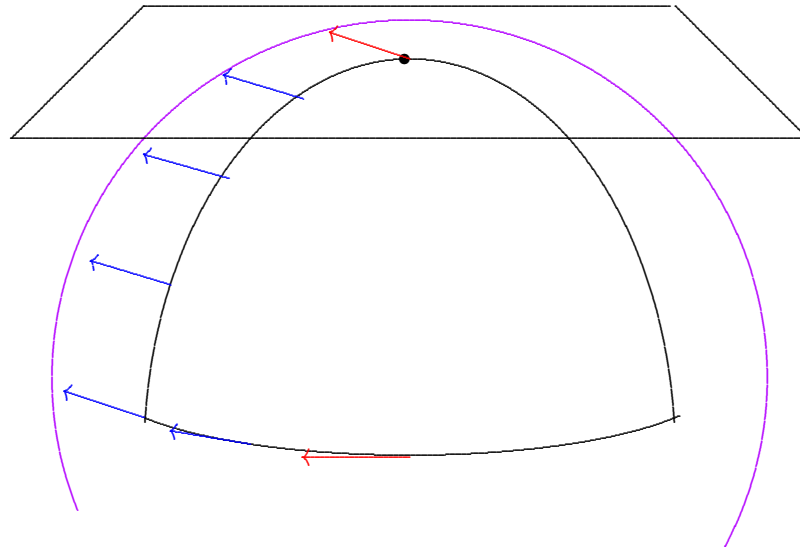
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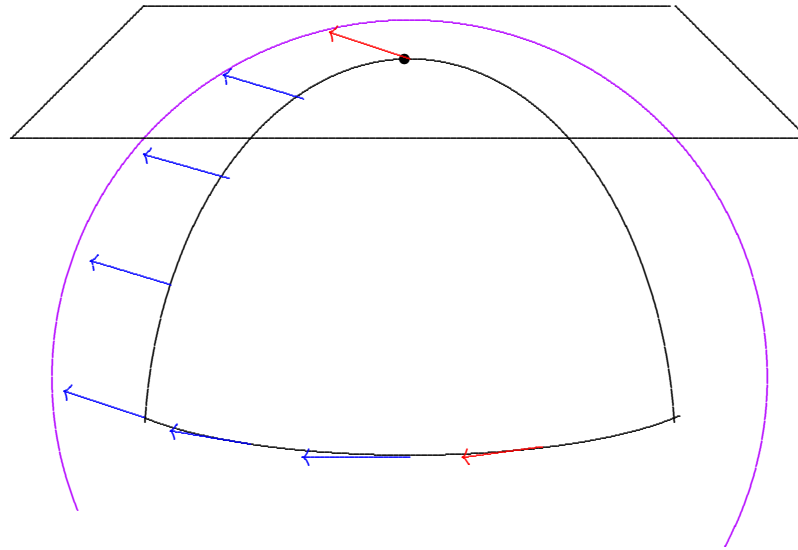
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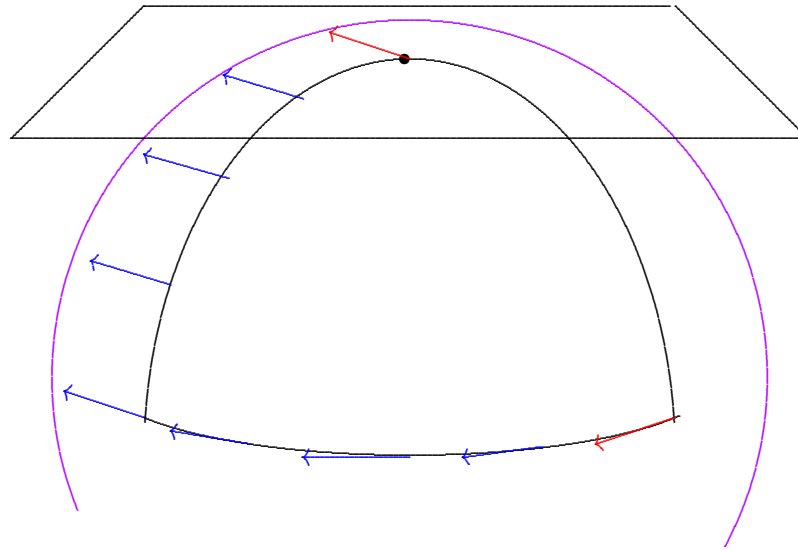
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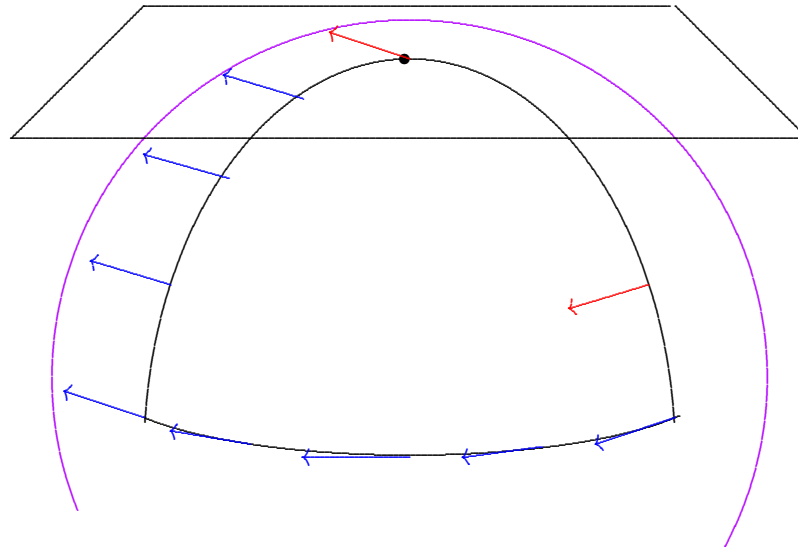
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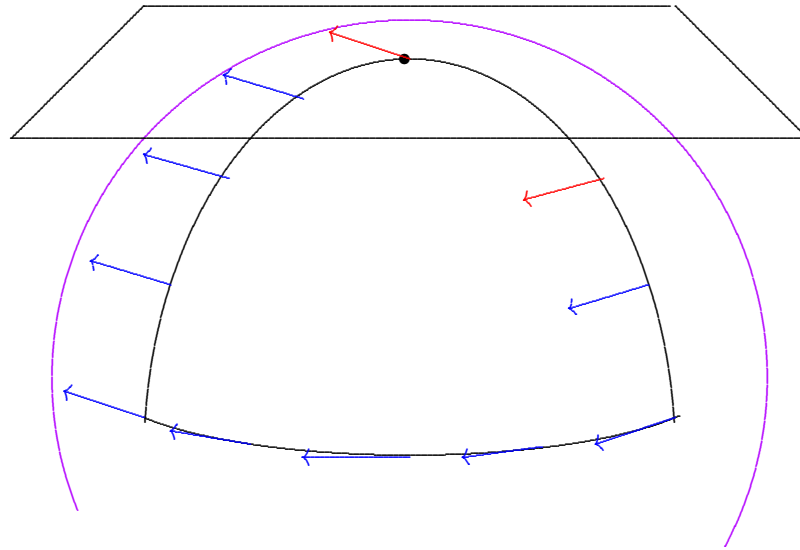
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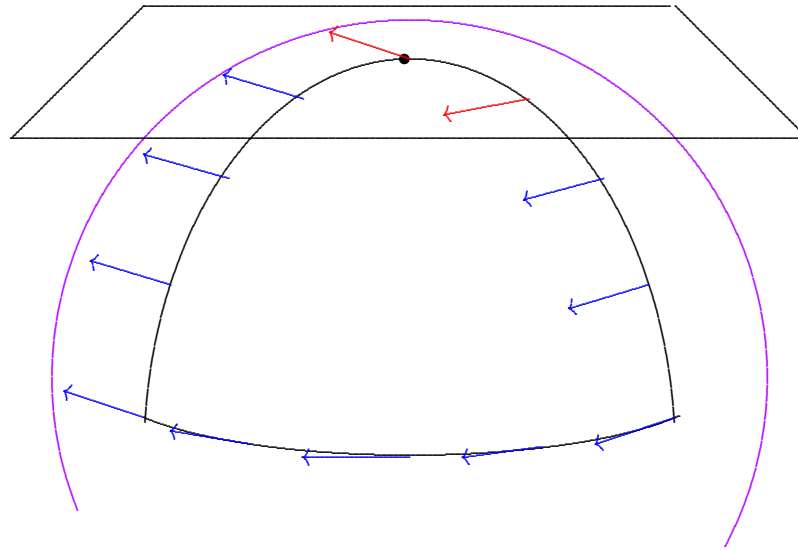
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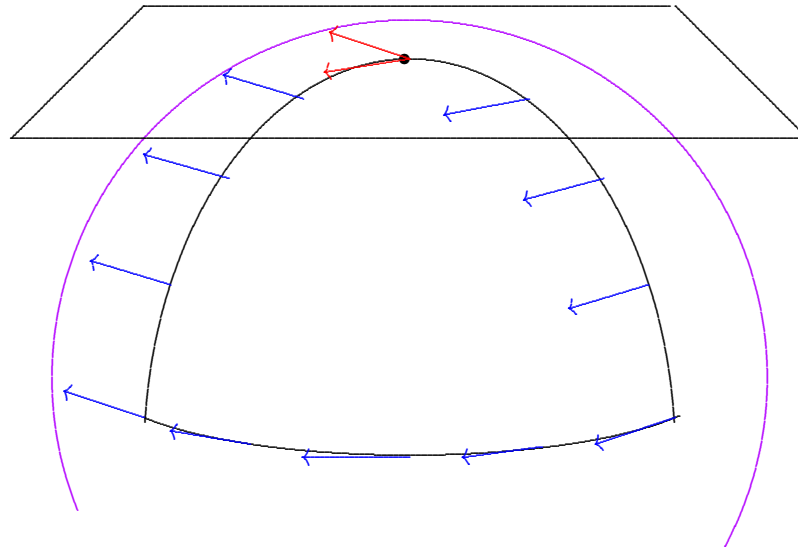
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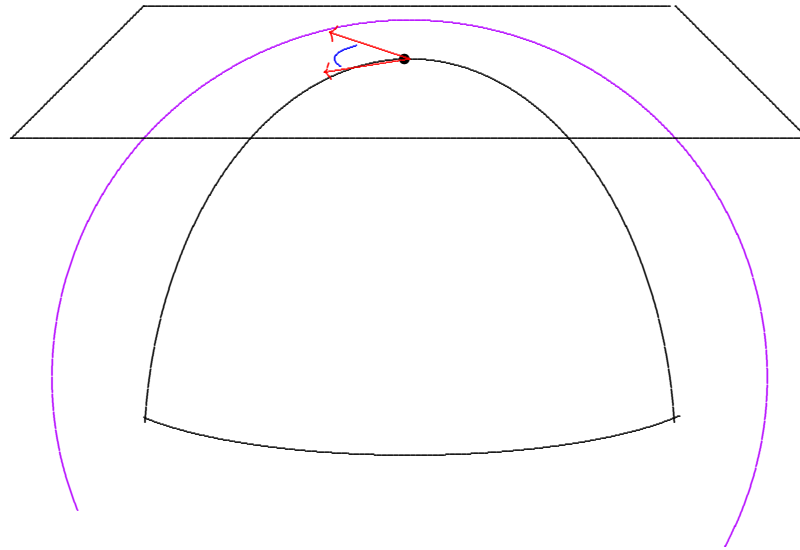
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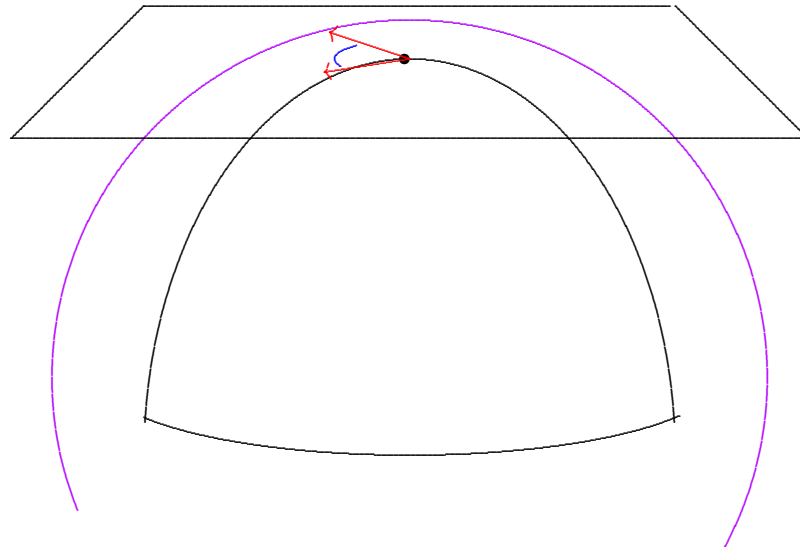
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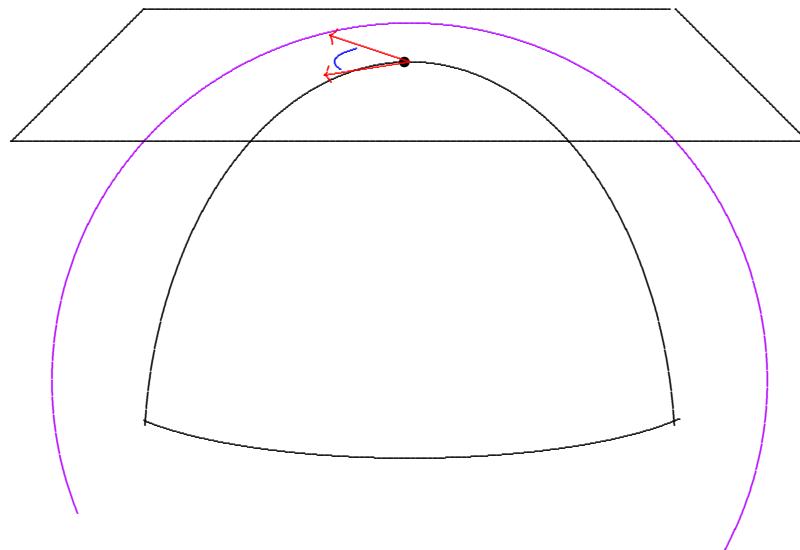
holonomy $\subset O(n)$



Kähler metrics:

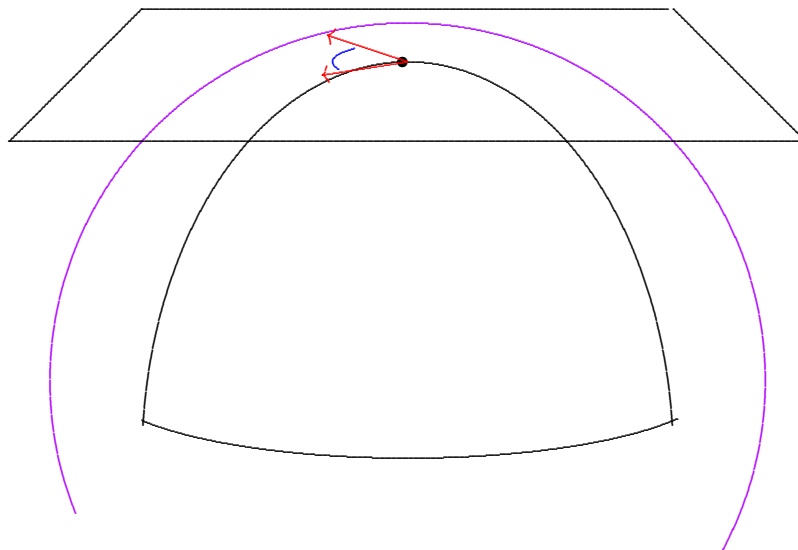
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is curvature of canonical line bundle $K = \Lambda^{m,0}$.

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$$r_{j\bar{k}} = -\frac{\partial^2}{\partial z^j \partial \bar{z}^k} \log \det[h_{\ell\bar{m}}]$$

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Both are actually Hermitian.

Theorem A.

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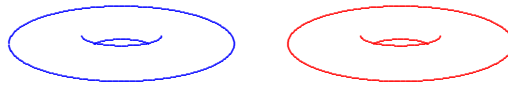
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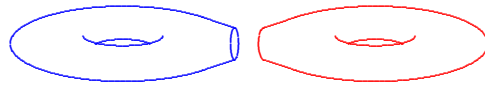
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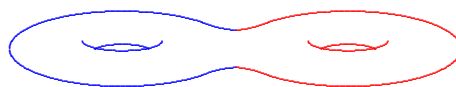
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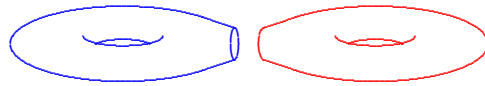
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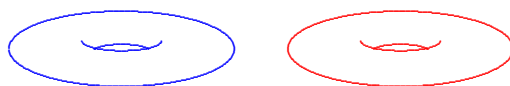
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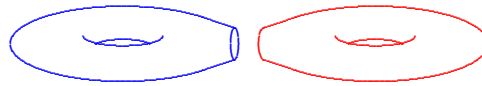
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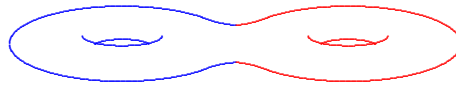
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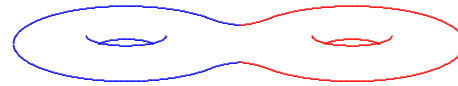
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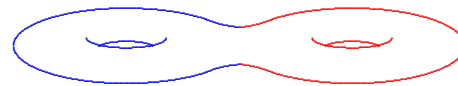


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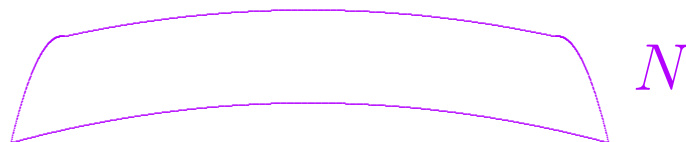
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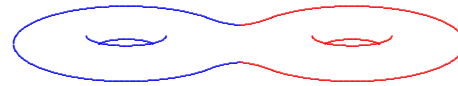
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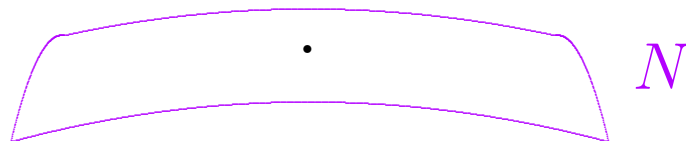
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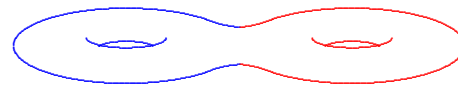
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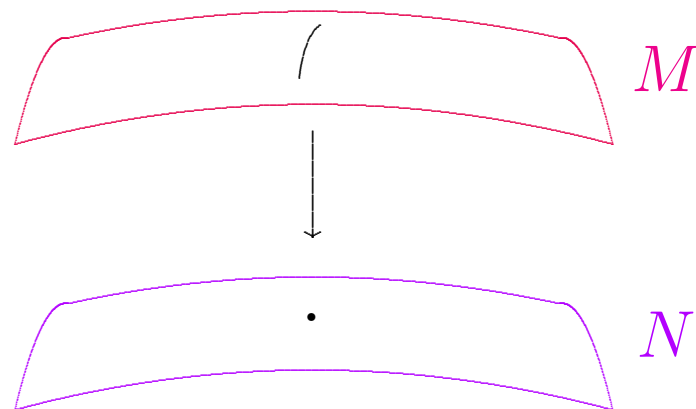
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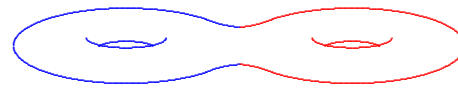
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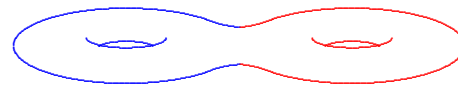
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in which new $\mathbb{C}P_1$ has self-intersection -1 .

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Then either

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- *$M \approx \mathbb{C}P_2 \# \overline{\mathbb{C}P}_2$, and h is a constant times the Page metric; or*
- *$M \approx \mathbb{C}P_2 \# 2\overline{\mathbb{C}P}_2$ and h is a constant times the Chen-LeBrun-Weber metric.*

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Exceptional cases: $\mathbb{C}P_2$ blown up at 1 or 2 points.

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Warning: when h is non-Kähler, its relation to ω is surprisingly complicated!

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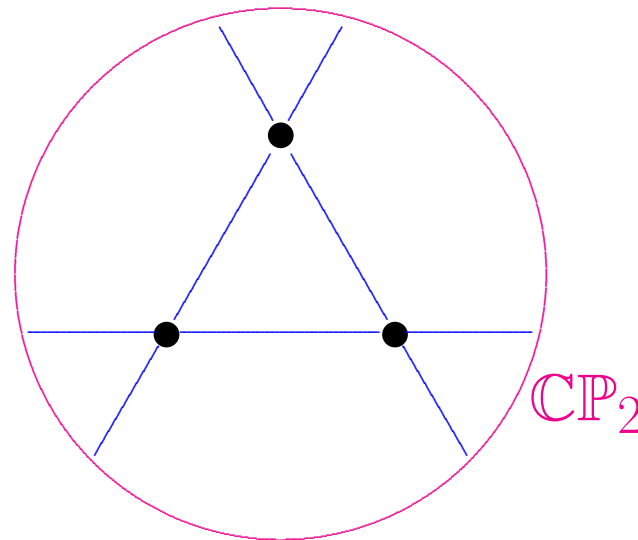
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In other words,

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\exists Kähler metric g , smooth function $f : M \rightarrow \mathbb{R}^+$.

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Similarly for $S^{2n+1} \times S^{2m+1}$.

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X.X. Chen: always minimizers.

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Donaldson/Mabuchi/Chen-Tian:

unique in Kähler class, modulo bihomorphisms.

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Λ^+ self-dual 2-forms.

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Notice that W_+ has a repeated eigenvalue.

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$\nabla \cdot W_+ = 0$, while $T^{1,0}M$ isotropic & involutive.

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When $W_+ \equiv 0$, use global results of Boyer et al.

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1-parameter family of metrics

$$g_t := g + t\dot{g} + O(t^2)$$

First variation

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$$\text{Conformally Einstein} \implies B = 0$$

Restriction of \mathcal{W}_+ to Kähler metrics?

On Kähler metrics,

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In fact, for Kähler metrics,

$$B = \frac{1}{12} \left[2s\dot{r} + \text{Hess}_0(s) + 3J^* \text{Hess}_0(s) \right]$$

where Hess_0 denotes trace-free part of $\nabla\nabla$.

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So the critical metrics of restriction of \mathcal{W}_+ to $\{\text{Kähler metrics}\}$ are Bach-flat Kähler metrics.

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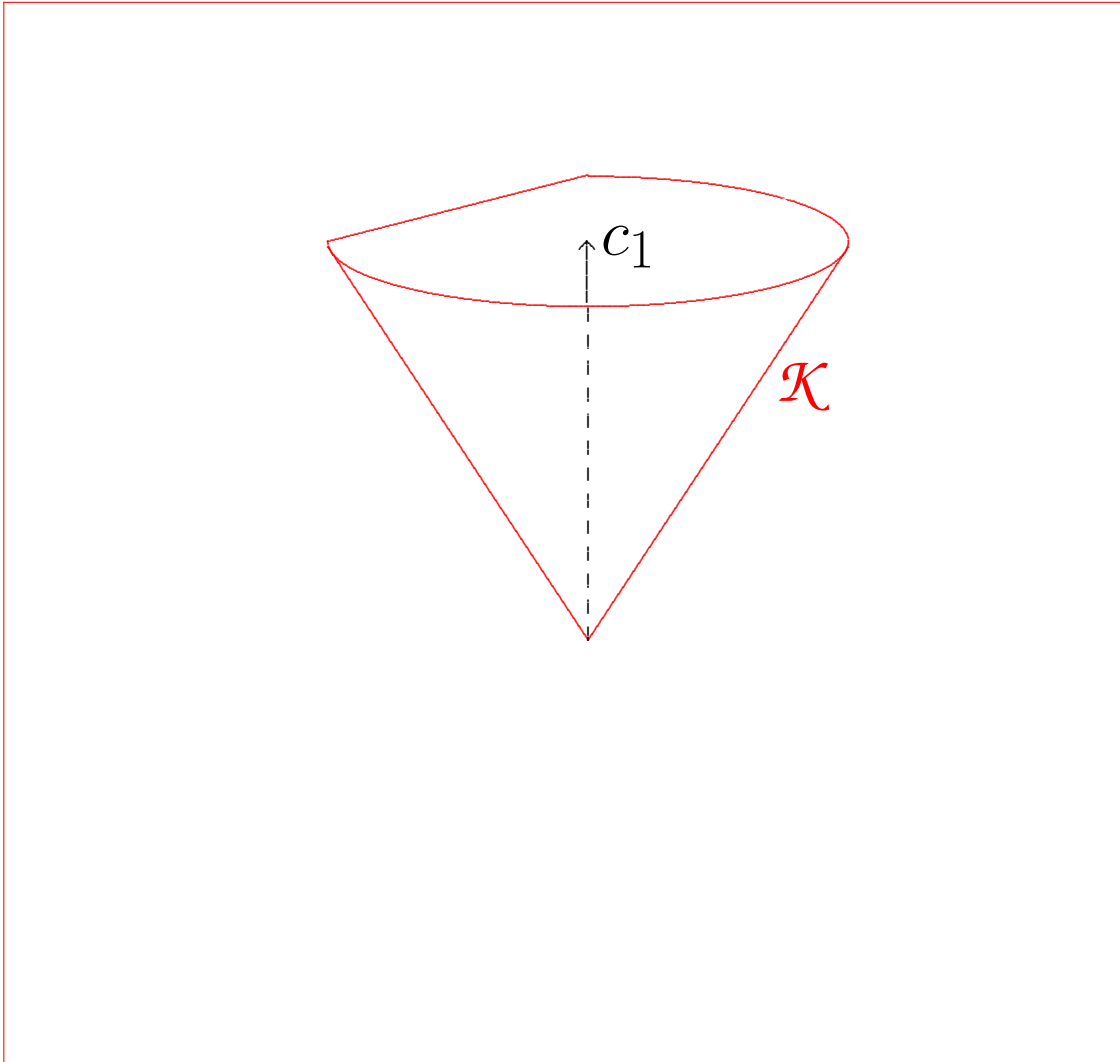
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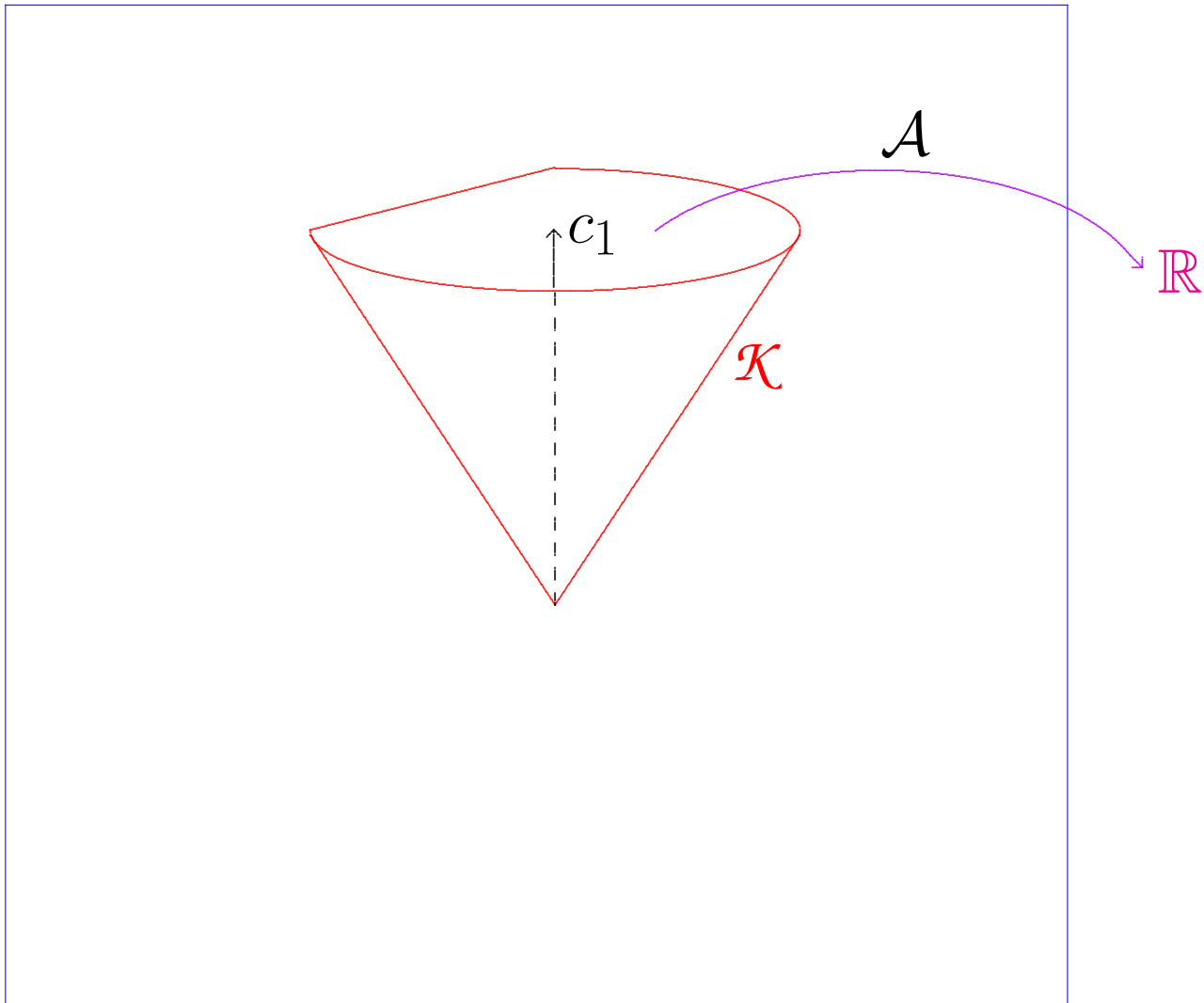
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$$\rho + 2i\partial\bar{\partial}\log s > 0.$$

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and uniqueness [Theorem A](#) follows.

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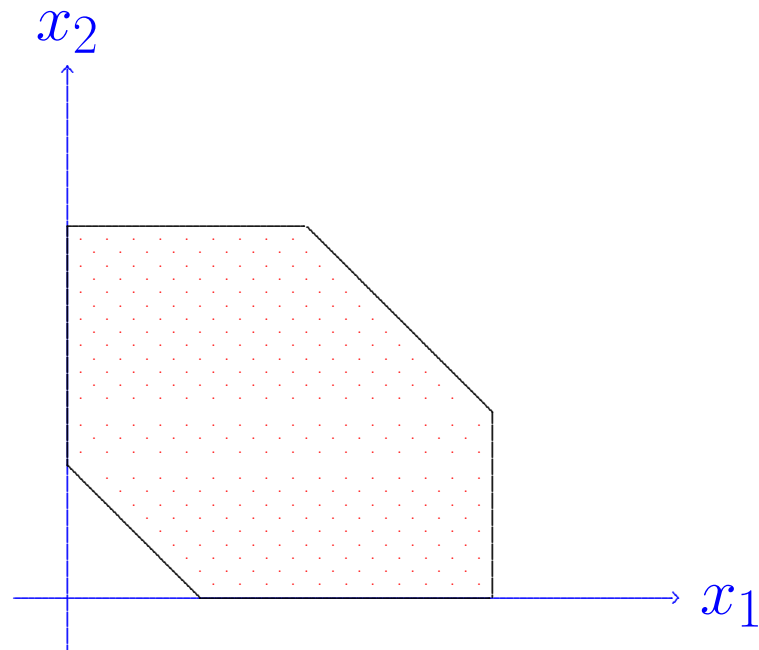
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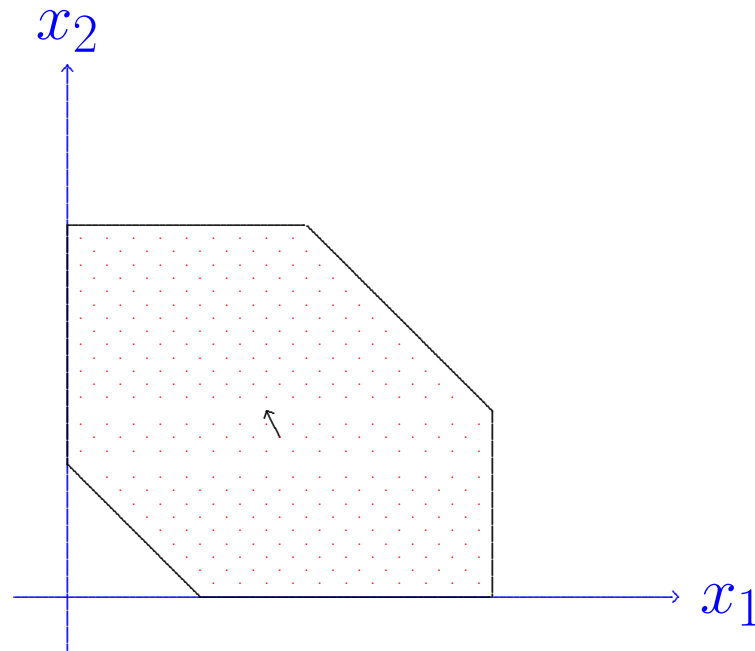
Only three cases are non-trivial:

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$$\mathcal{A}([\omega]) = \frac{|\partial P|^2}{2} \left(\frac{1}{|P|} + \vec{\mathfrak{D}} \cdot \Pi^{-1} \vec{\mathfrak{D}} \right)$$

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$$\begin{aligned}
& 3[3 + 28\gamma + 96\gamma^2 + 168\gamma^3 + 164\gamma^4 + 80\gamma^5 + 16\gamma^6 + 16\beta^6(1 + \gamma)^4 + 16\alpha^6(1 + \beta + \gamma)^4 + 16\beta^5(5 + 24\gamma + 43\gamma^2 + 37\gamma^3 + 15\gamma^4 + 2\gamma^5) + 4\beta^4(41 + 228\gamma + 478\gamma^2 + 496\gamma^3 + 263\gamma^4 + \\
& 60\gamma^5 + 4\gamma^6) + 8\beta^3(21 + 135\gamma + 326\gamma^2 + 392\gamma^3 + 248\gamma^4 + 74\gamma^5 + 8\gamma^6) + 4\beta(7 + 58\gamma + 176\gamma^2 + 270\gamma^3 + 228\gamma^4 + 96\gamma^5 + 16\gamma^6) + 4\beta^2(24 + 176\gamma + 479\gamma^2 + 652\gamma^3 + 478\gamma^4 + \\
& 172\gamma^5 + 24\gamma^6) + 16\alpha^5(5 + 2\beta^5 + 24\gamma + 43\gamma^2 + 37\gamma^3 + 15\gamma^4 + 2\gamma^5 + \beta^4(15 + 14\gamma) + \beta^3(37 + 70\gamma + 30\gamma^2) + \beta^2(43 + 123\gamma + 108\gamma^2 + 30\gamma^3) + \beta(24 + 92\gamma + 123\gamma^2 + 70\gamma^3 + \\
& 14\gamma^4)) + 4\alpha^4(41 + 4\beta^6 + 228\gamma + 478\gamma^2 + 496\gamma^3 + 263\gamma^4 + 60\gamma^5 + 4\gamma^6 + \beta^5(60 + 56\gamma) + \beta^4(263 + 476\gamma + 196\gamma^2) + 8\beta^3(62 + 169\gamma + 139\gamma^2 + 35\gamma^3) + 2\beta^2(239 + 876\gamma + 1089\gamma^2 + \\
& 556\gamma^3 + 98\gamma^4) + 4\beta(57 + 263\gamma + 438\gamma^2 + 338\gamma^3 + 119\gamma^4 + 14\gamma^5)) + 8\alpha^3(21 + 135\gamma + 326\gamma^2 + 392\gamma^3 + 248\gamma^4 + 74\gamma^5 + 8\gamma^6 + 8\beta^6(1 + \gamma) + 2\beta^5(37 + 70\gamma + 30\gamma^2) + 4\beta^4(62 + \\
& 169\gamma + 139\gamma^2 + 35\gamma^3) + 4\beta^3(98 + 353\gamma + 428\gamma^2 + 210\gamma^3 + 35\gamma^4) + 2\beta^2(163 + 735\gamma + 1179\gamma^2 + 856\gamma^3 + 278\gamma^4 + 30\gamma^5) + \beta(135 + 736\gamma + 1470\gamma^2 + 1412\gamma^3 + 676\gamma^4 + 140\gamma^5 + \\
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& [1 + 10\gamma + 36\gamma^2 + 64\gamma^3 + 60\gamma^4 + 24\gamma^5 + 24\beta^5(1 + \gamma)^5 + 24\alpha^5(1 + \beta + \gamma)^5 + 12\beta^4(1 + \gamma)^2(5 + 20\gamma + 23\gamma^2 + 10\gamma^3) + 16\beta^3(4 + 28\gamma + 72\gamma^2 + 90\gamma^3 + 57\gamma^4 + 15\gamma^5) + \\
& 12\beta^2(3 + 24\gamma + 69\gamma^2 + 96\gamma^3 + 68\gamma^4 + 20\gamma^5) + 2\beta(5 + 45\gamma + 144\gamma^2 + 224\gamma^3 + 180\gamma^4 + 60\gamma^5) + 12\alpha^4(1 + \beta + \gamma)^2(5 + 20\gamma + 23\gamma^2 + 10\gamma^3 + 10\beta^3(1 + \gamma) + \beta^2(23 + 46\gamma + \\
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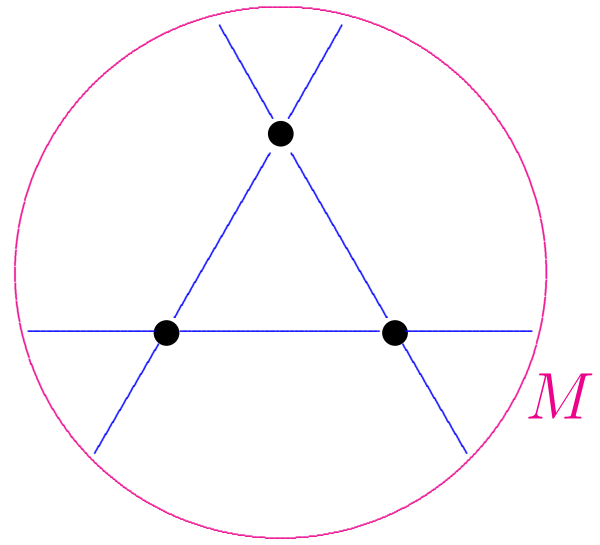
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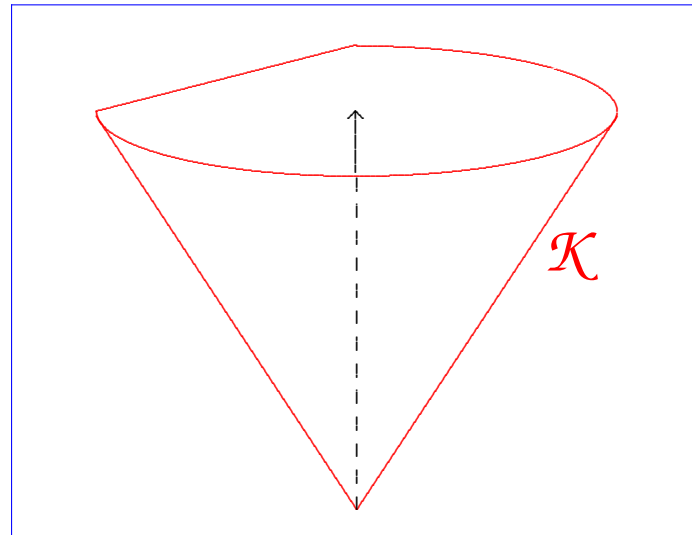
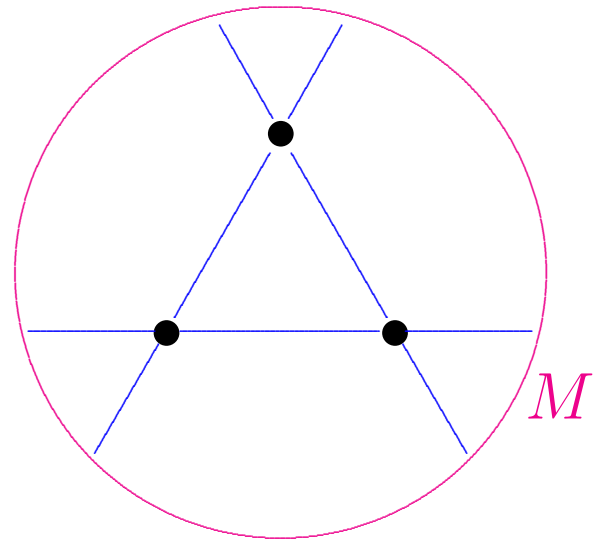
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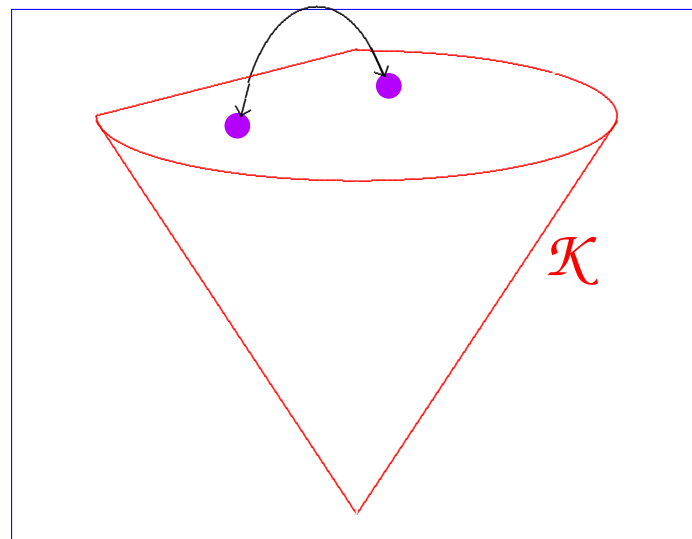
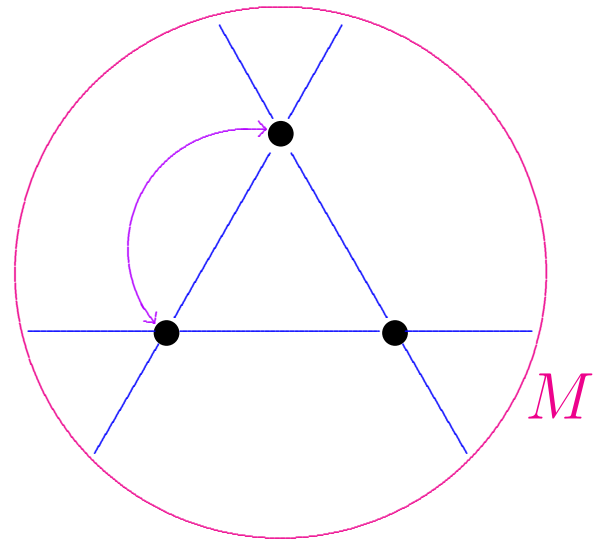
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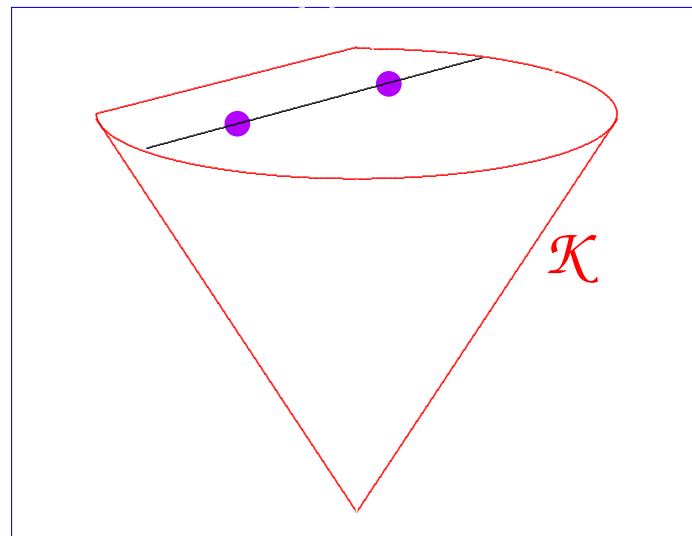
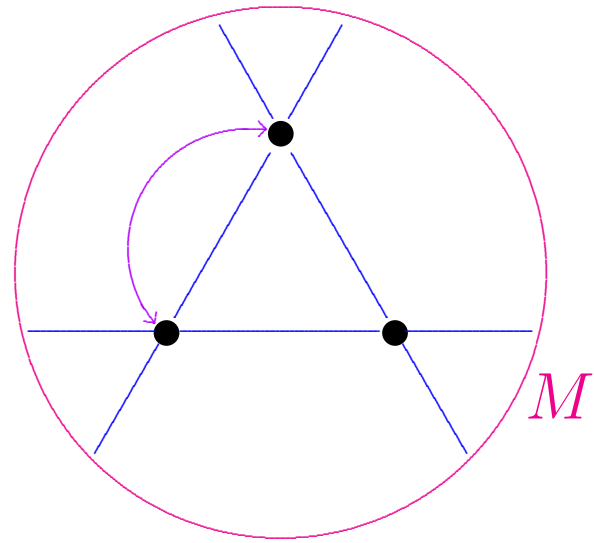
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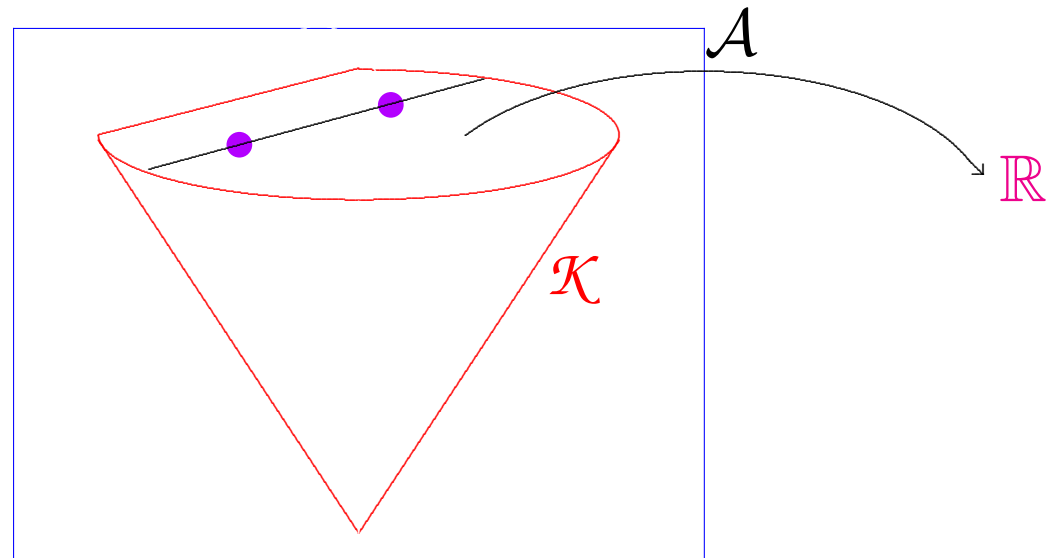
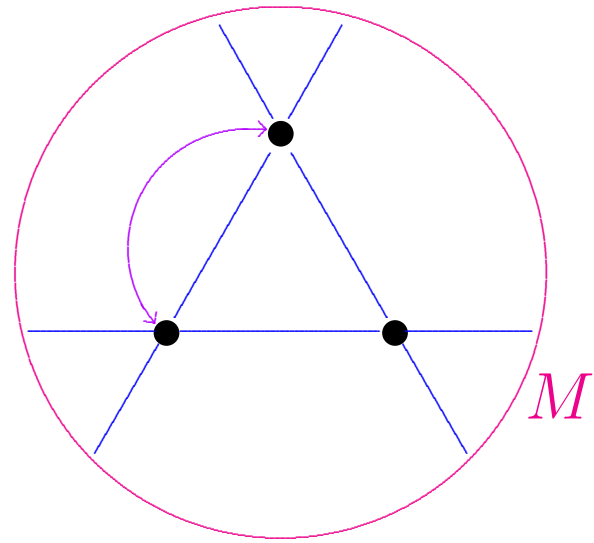
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Final step then just calculus in one variable...

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Similar calculations also led to new existence proof. . .

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Could also reconstruct Page metric this way...

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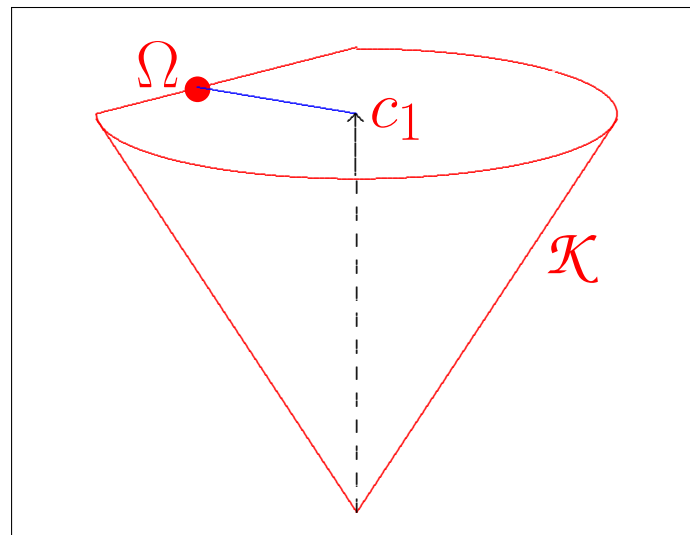
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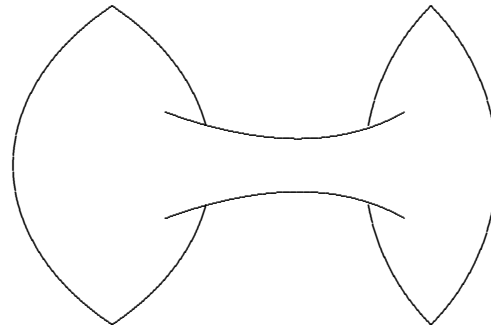
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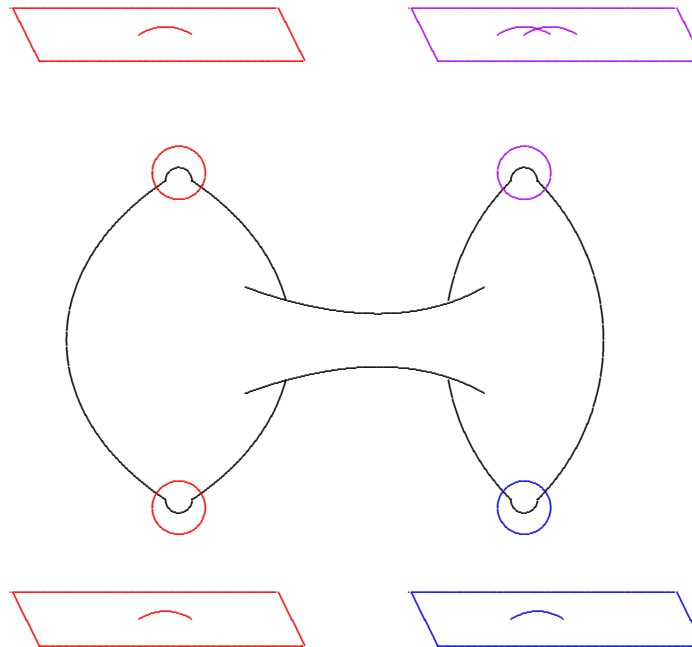
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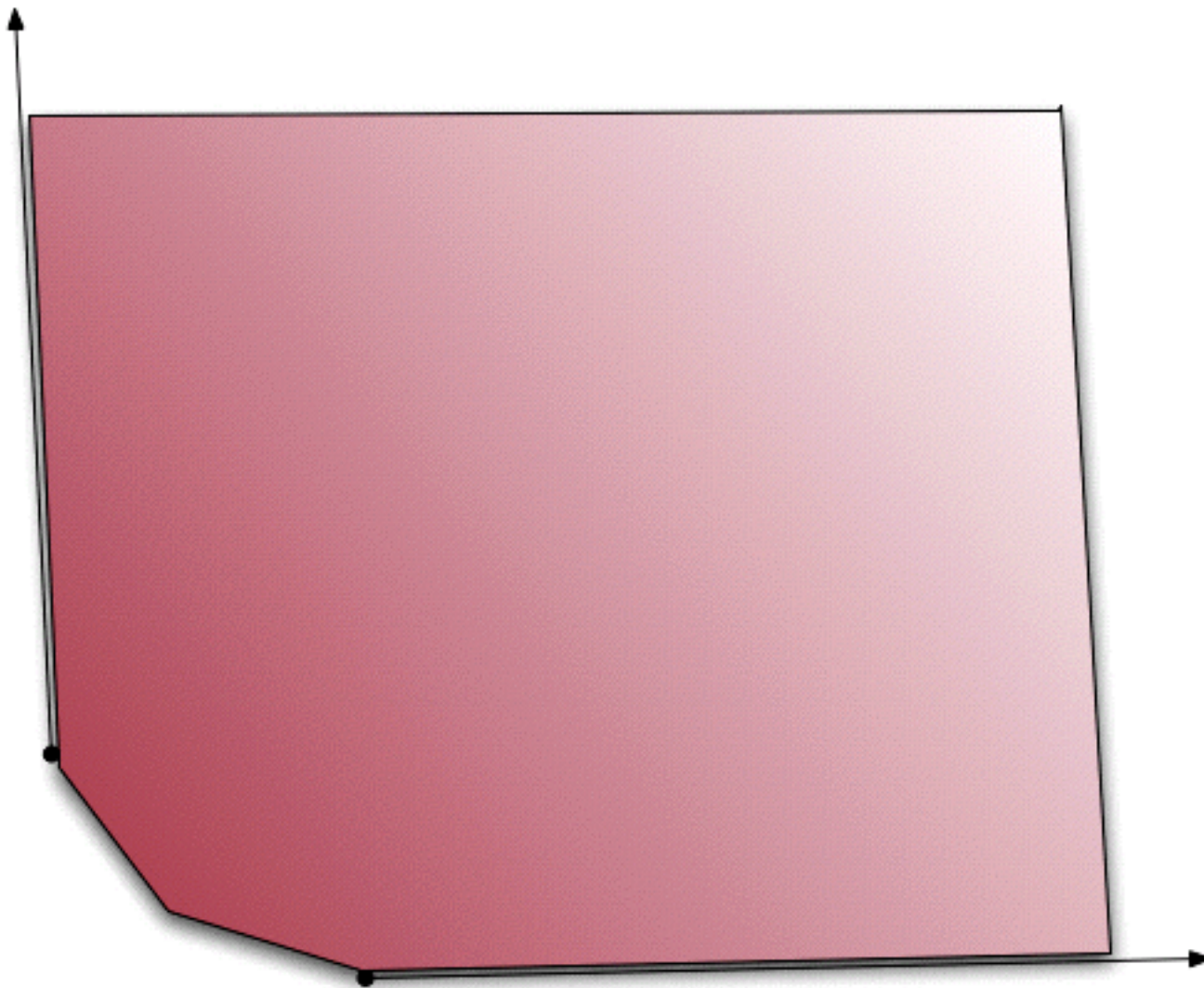
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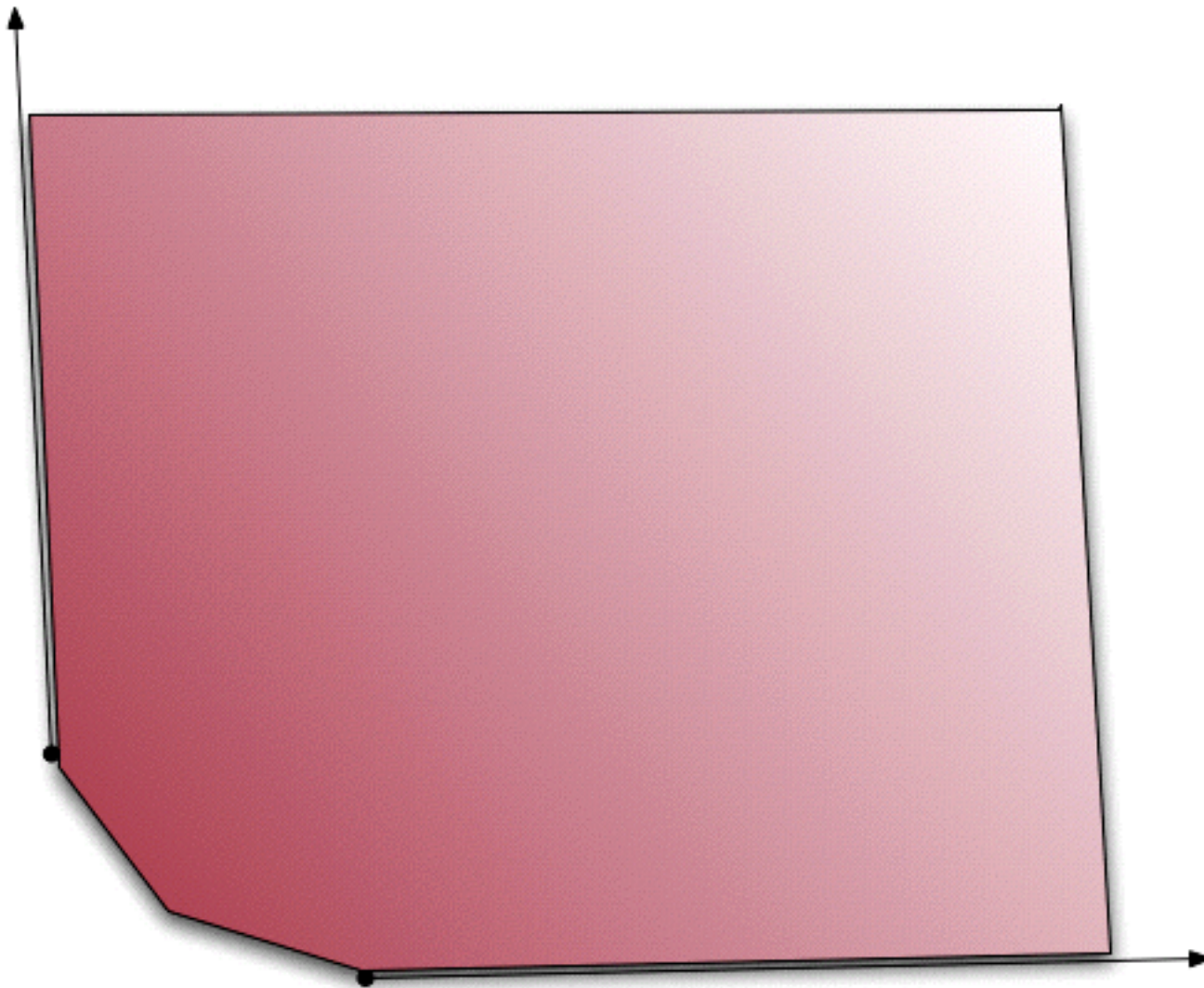
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But every such class in M represented by a holomorphic curve! So $\Omega_{t_\infty} = \Omega_1$, and we have just bubbled off a (-1) -curve, as desired!

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