

The Einstein-Weyl Equations,

Scattering Maps,

and

Holomorphic Disks

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SUNY Stony Brook

Joint work with

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University of Oxford

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Some overlap
with results of

Fuminori Nakata

Tokyo Institute of Technology

Weyl's 1918 gauge theory

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based on Weyl connections $([g], \nabla)$:

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$$\nabla_v g \propto g \quad \forall v$$

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where $\nu = d \log u$.

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where $n = \dim M$.

Hermann Weyl



$$\mathfrak{B} = (\mathfrak{G} + \alpha \mathbf{I}) + \frac{\varepsilon^2}{4} V \bar{g} \{ \mathbf{1} - 3 (\varphi_i \varphi^i) \},$$

$$\Gamma_{ik}^r = \left\{ \begin{matrix} ik \\ r \end{matrix} \right\} + \frac{1}{2} \varepsilon^2 (\delta_i^r \varphi_k + \delta_k^r \varphi_i - g_{ik} \varphi^r).$$

Unter Vernachlässigung der winzigen kosmologischen Terme erhalten wir hier also genau die klassische Maxwell-Einsteinsche Theorie der Elektrizität und Gravitation. Um Übereinstimmung mit den in § 34 verwendeten

Ricci tensor

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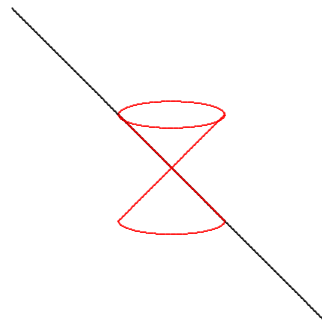
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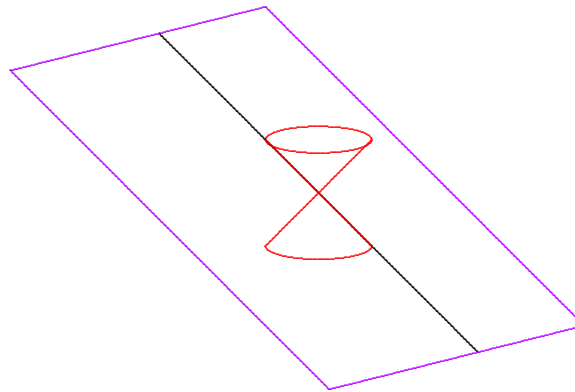
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THÉORÈME. — Les espaces de Weyl à trois dimensions qui admettent ∞^2 plans isotropes dépendent essentiellement de quatre fonctions arbitraires de deux arguments.

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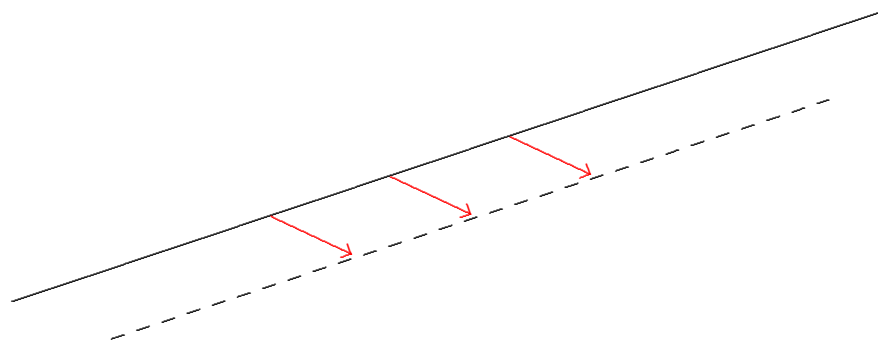
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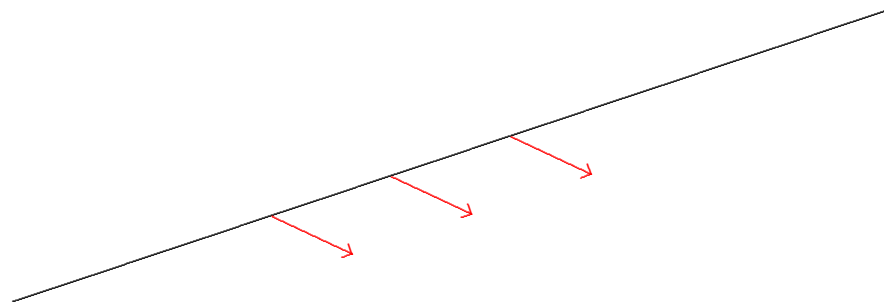
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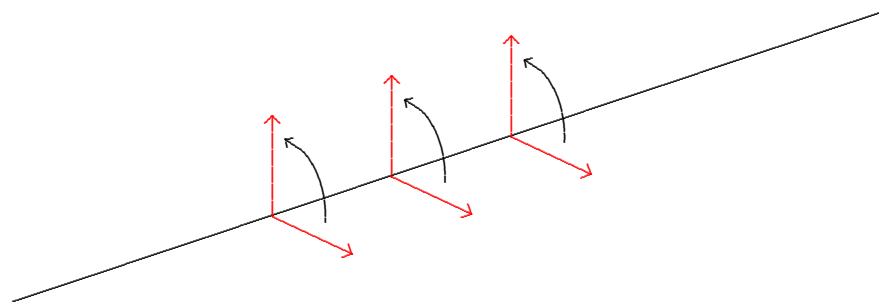
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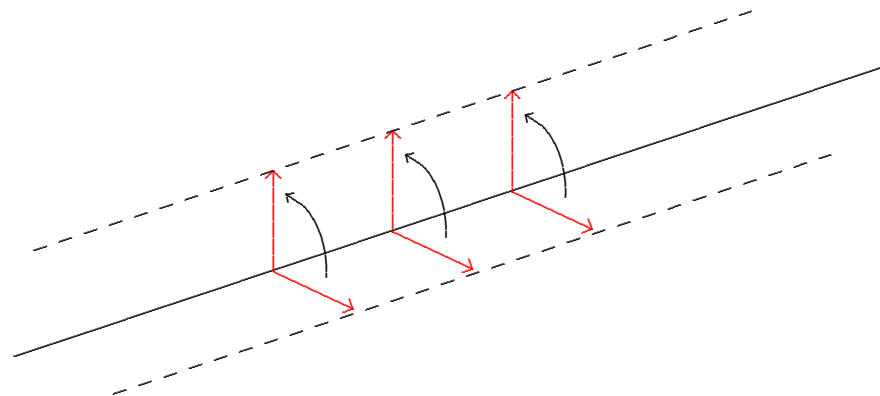
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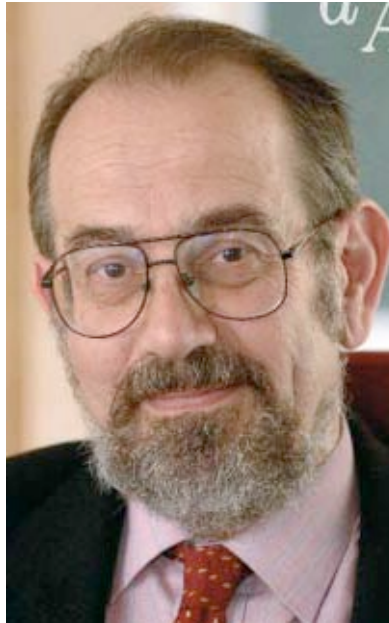
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Nigel Hitchin



and so if

$$R(U \times V, U)U = U \times R(V, U)U, \quad (2.2)$$

then we can define a linear map

$$J(V) = U \times V \quad (2.3)$$

which satisfies

$$J^2(V) = U \times (U \times V) = (U, V)U - (U, U)V = -V$$

We thus have a real complex surface G with a family of real lines of self-intersection number 2. It can be shown that any such surface may be obtained by the above geodesic construction, but using a Weyl structure rather than a Riemannian structure. The integrability condition (2.2) is then the analogue of Einstein's equations $(R_{(ij)} = \Lambda g_{ij})$ for the Weyl structure (see [10]). This is the

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- *smooth, space-time-oriented, conformally compact, globally hyperbolic Lorentzian Einstein-Weyl 3-manifolds $(M, [g], \nabla)$; and*
- *orientation-reversing diffeomorphisms*

$$\psi : \mathbb{C}P_1 \rightarrow \mathbb{C}P_1.$$

space-time oriented

Conformal Lorentzian n -manifold $(M, [g])$ called
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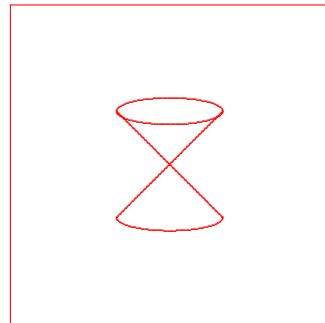
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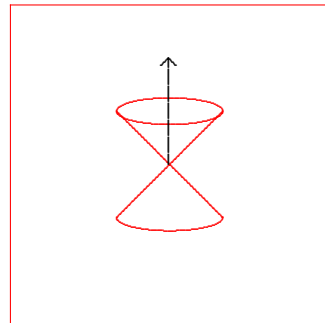
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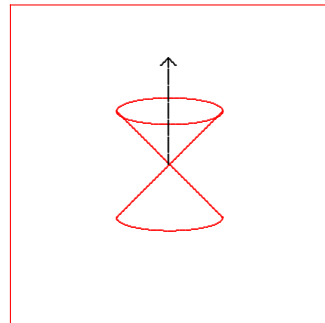
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$\implies M$ also oriented, in usual sense.

globally hyperbolic

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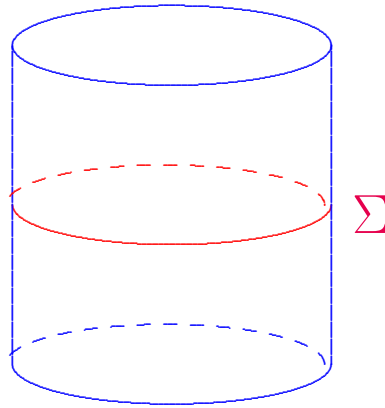
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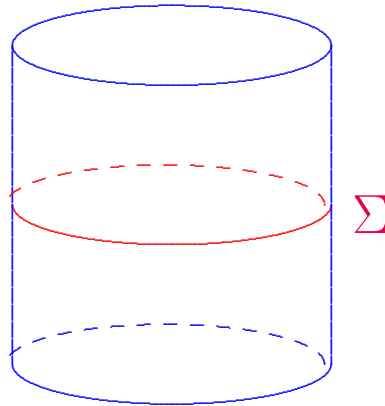
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$$\implies M \approx \Sigma \times \mathbb{R}$$

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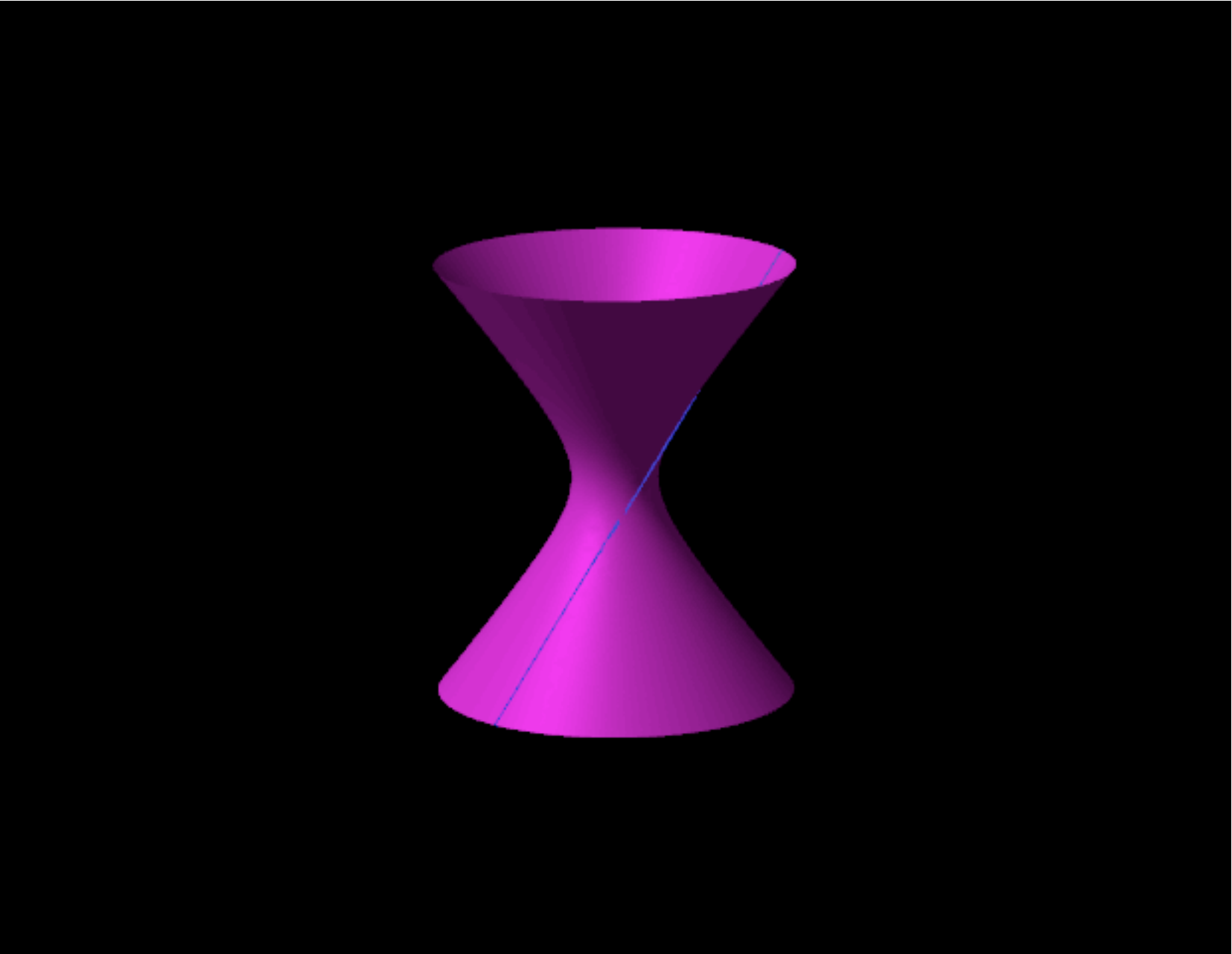
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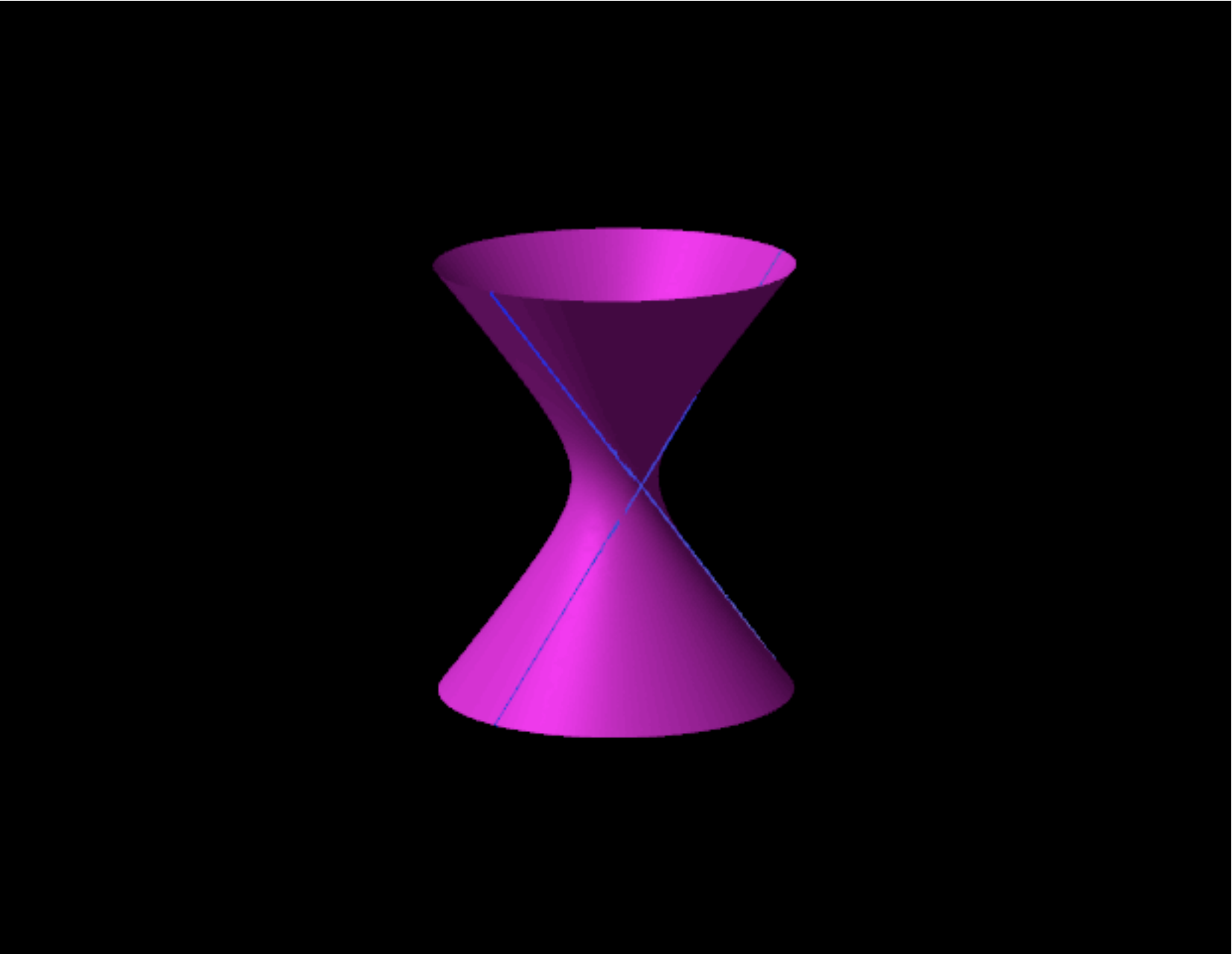
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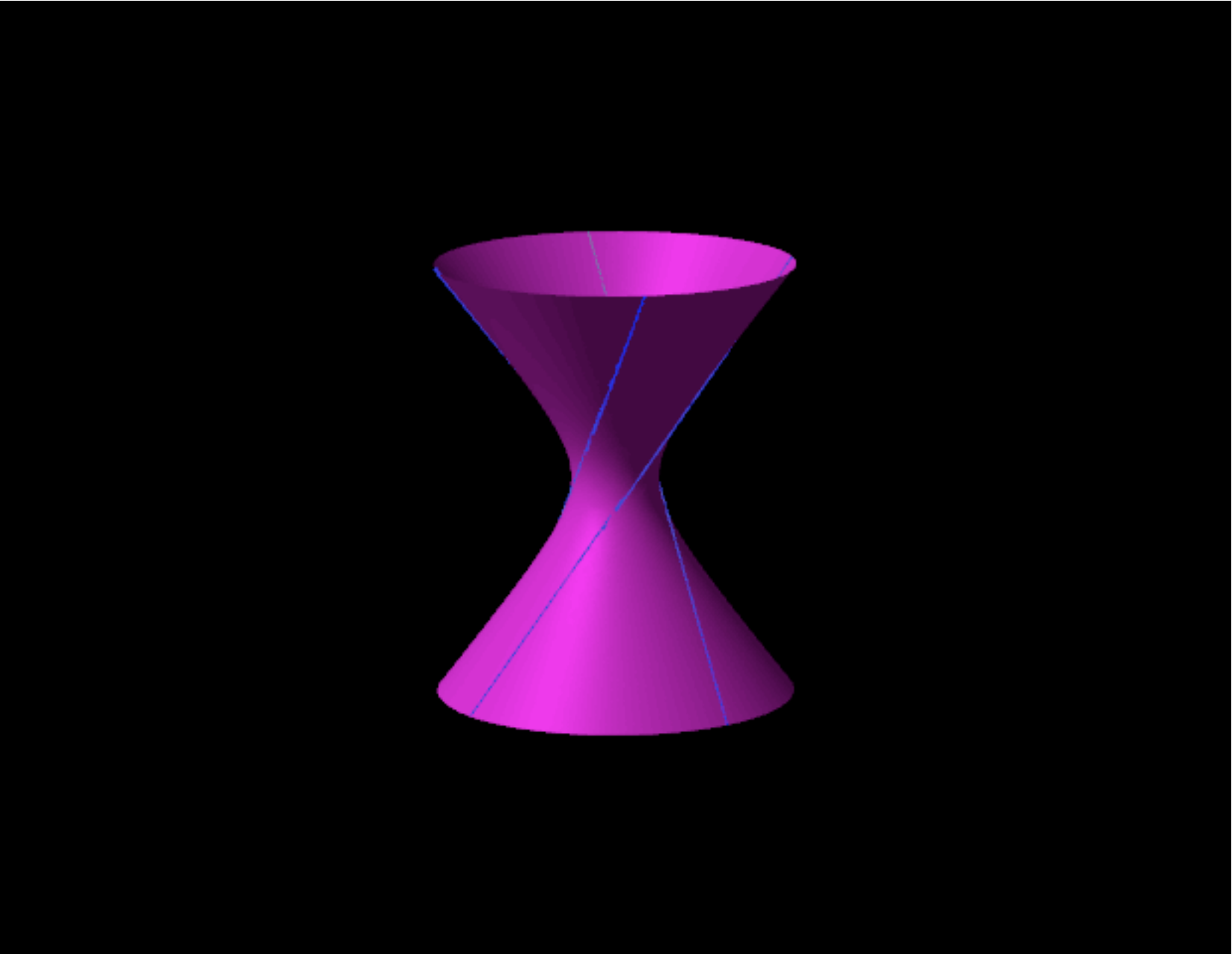
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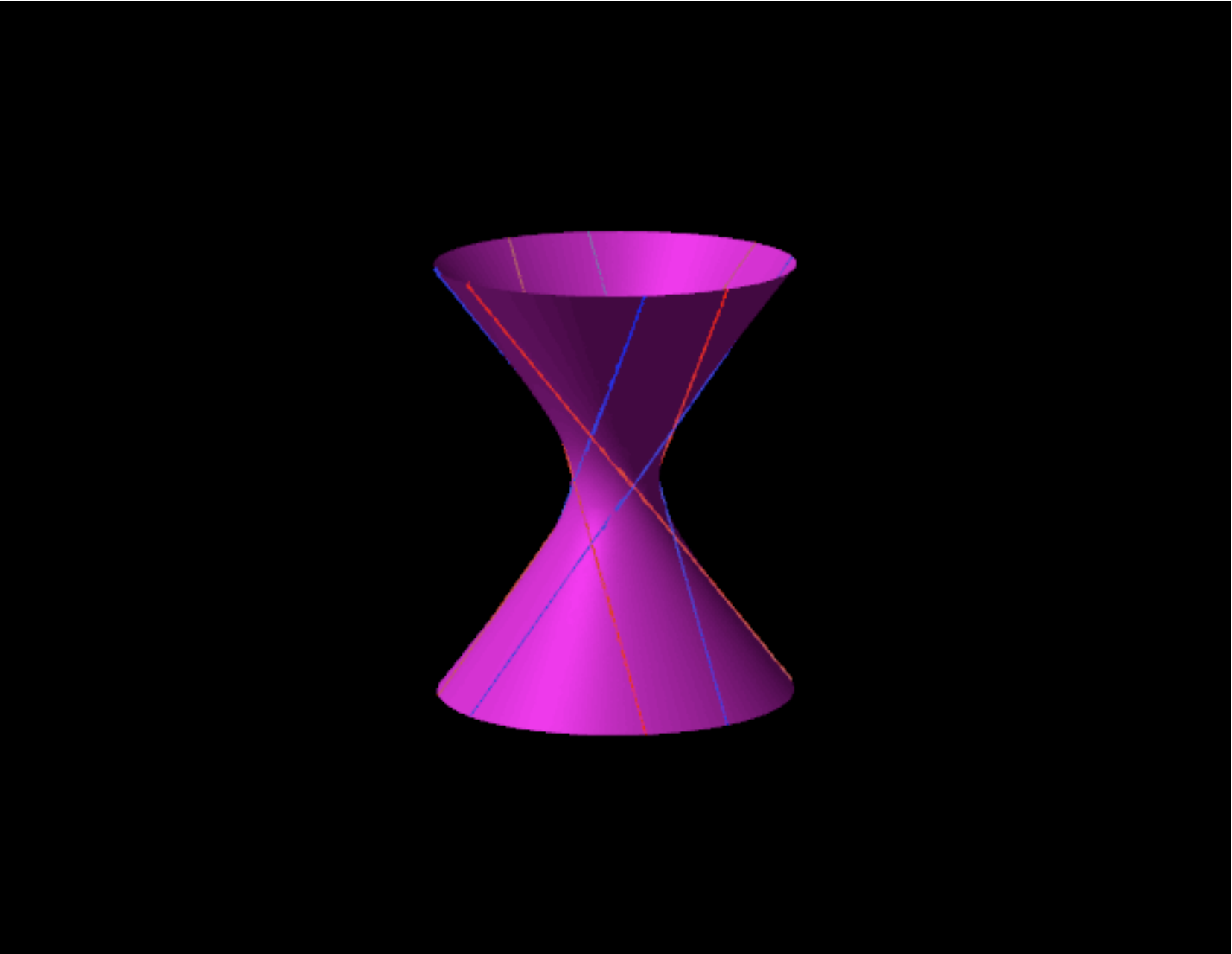
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where $h =$ standard metric on S^2 .

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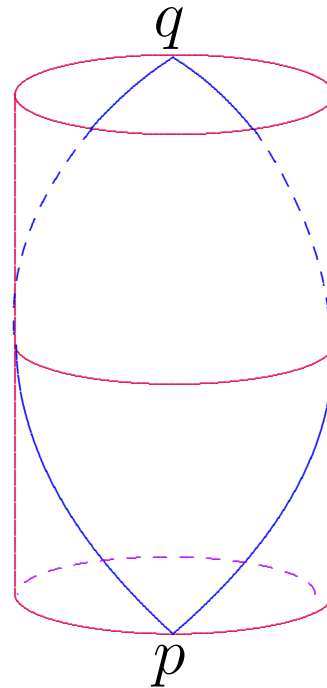
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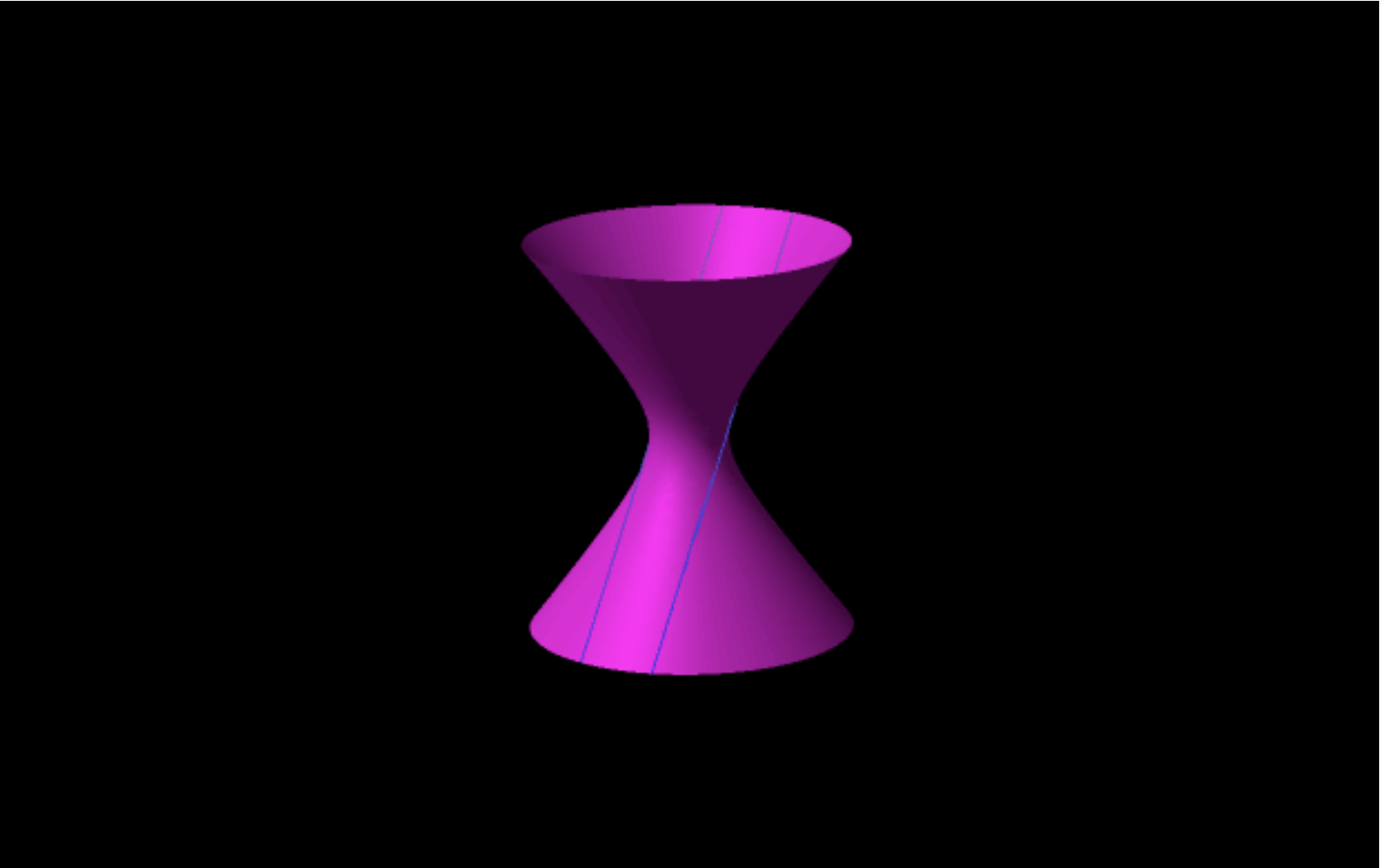
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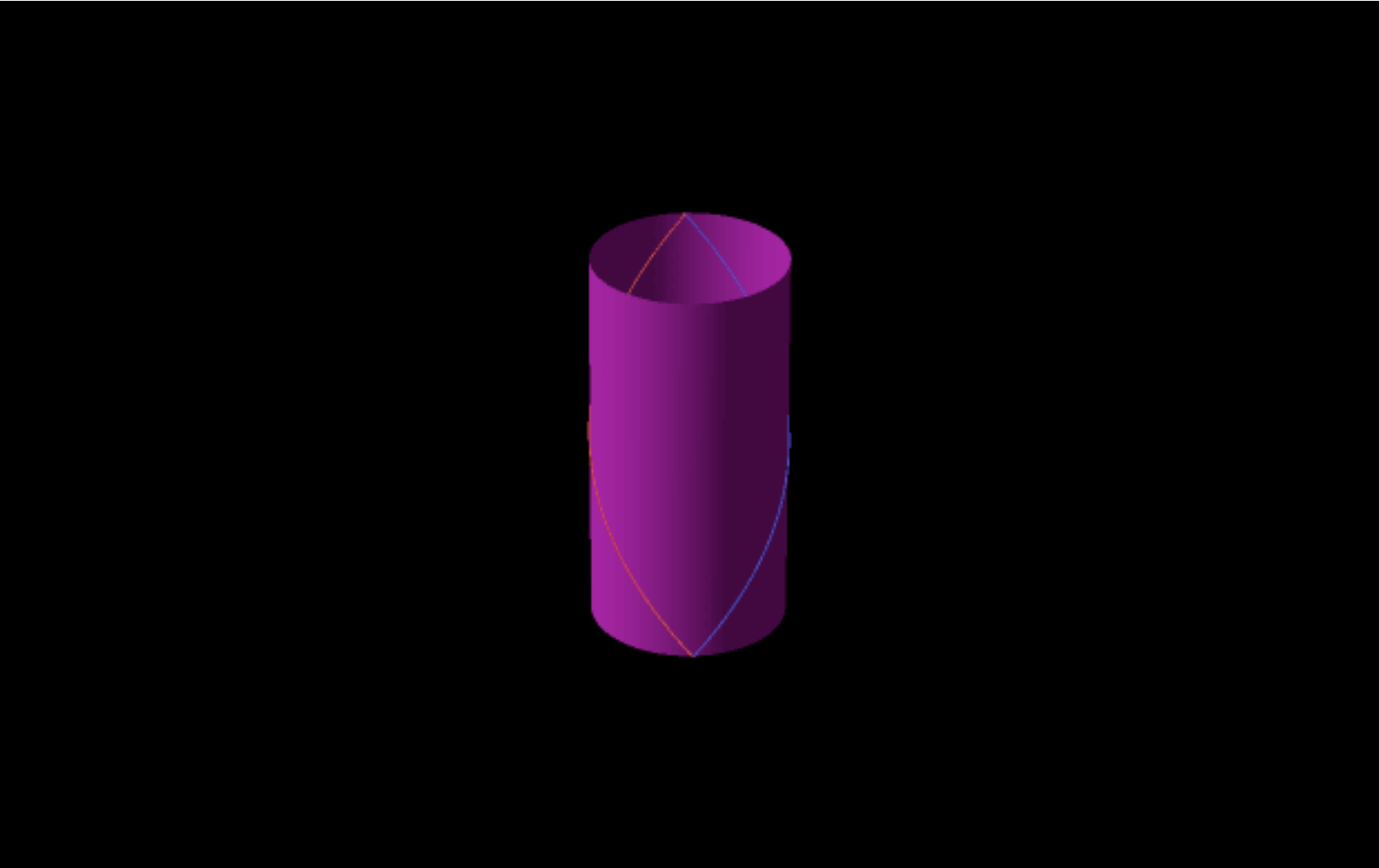
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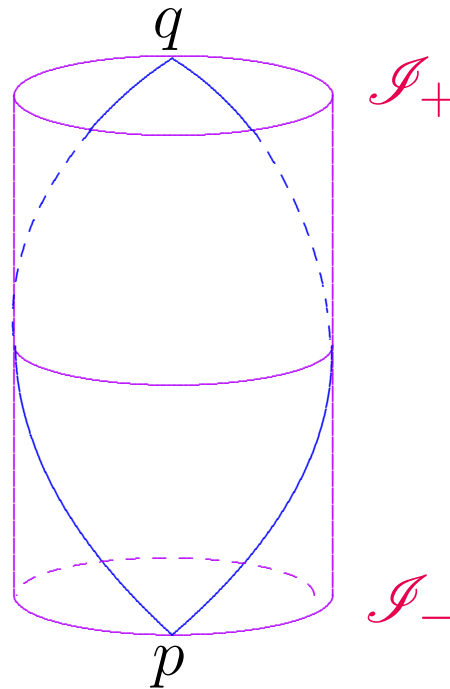


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$$\psi_1, \psi_2 : \mathbb{CP}_1 \rightarrow \mathbb{CP}_1$$

considered same iff

$$\psi_1 = \varphi \circ \psi_2 \circ \phi^{-1}$$

for Möbius transformations $\varphi, \phi \in PSL(2, \mathbb{C})$.

Theorem. *There is a natural one-to-one correspondence between*

- *smooth, space-time-oriented, conformally compact, globally hyperbolic Lorentzian Einstein-Weyl 3-manifolds $(M, [g], \nabla)$; and*
- *orientation-reversing diffeomorphisms*

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Example: de Sitter \longleftrightarrow antipodal map of \mathbb{CP}_1 .

In one direction, direct geometrical interpretation of correspondence in terms of scattering maps.

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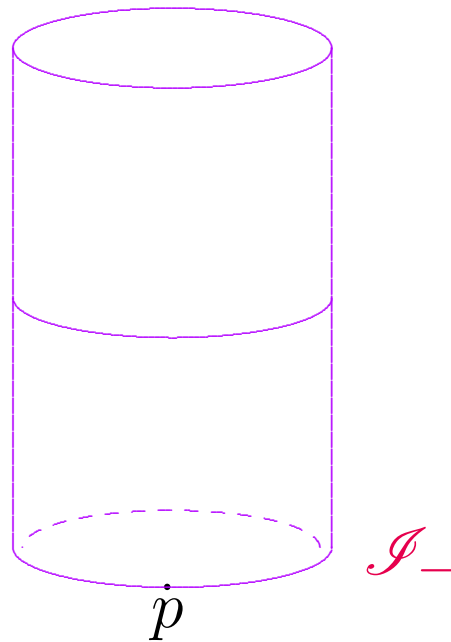
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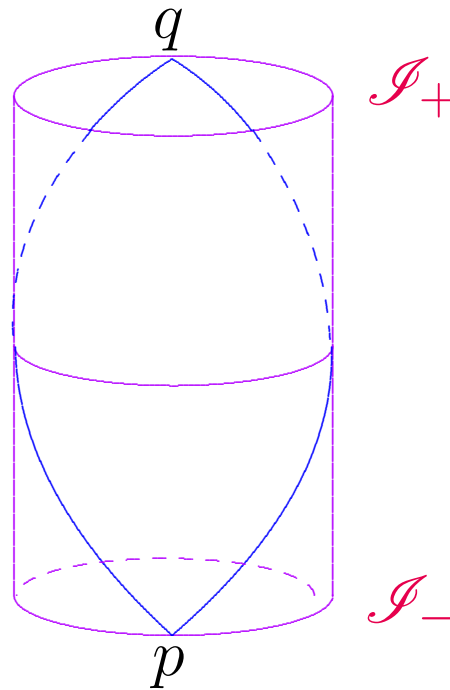
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to Einstein-Weyl $(M^3, [g], \nabla)$ satisfying hypotheses.

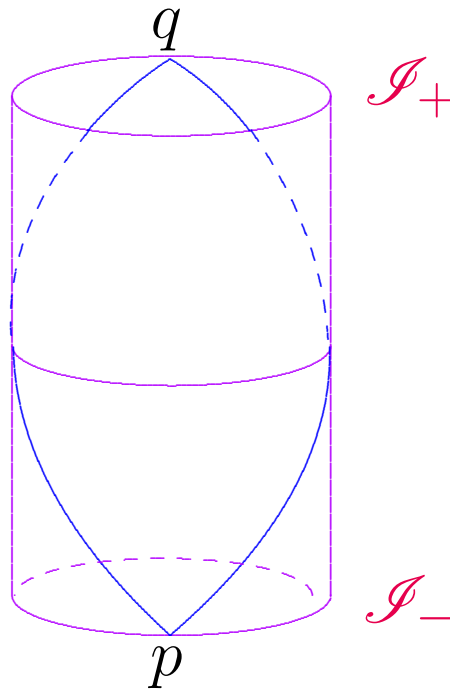
Lemma. *For an Einstein-Weyl manifold as above, let $p \in \mathcal{I}_-$ be any point of past infinity.*



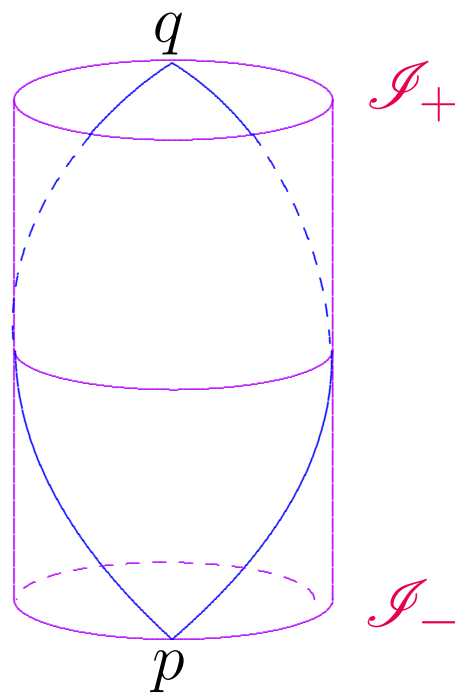
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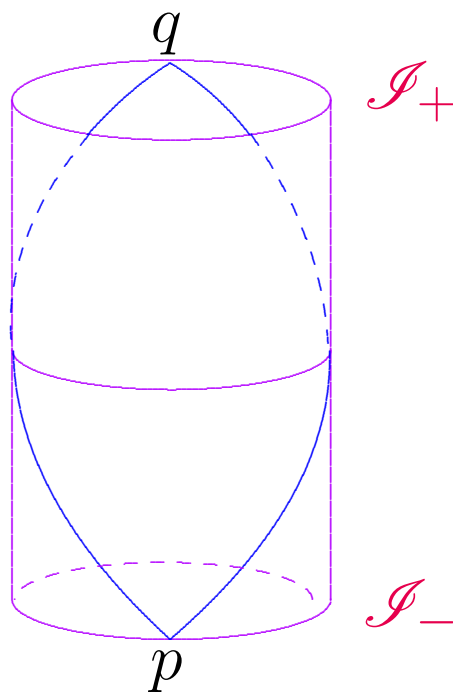
Lemma. For an Einstein-Weyl manifold as above, let $p \in \mathcal{I}_-$ be any point of past infinity. Then all the null geodesics emanating from p refocus at a unique point $q \in \mathcal{I}_+$. Moreover, $\mathcal{I}_\pm \approx S^2$.



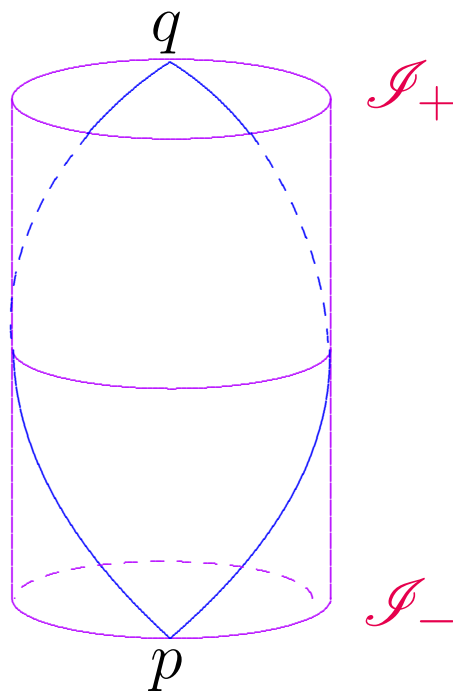
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Inverting correspondence:

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twistor disk construction

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Graph of orientation-reversing diffeomorphism

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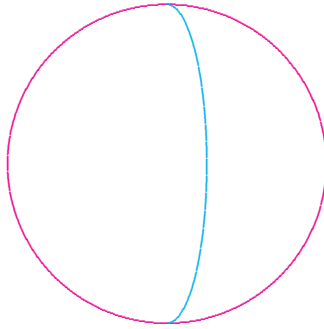
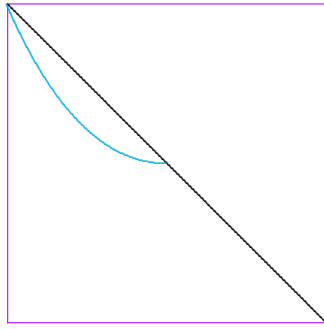
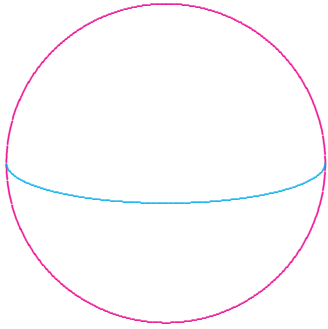
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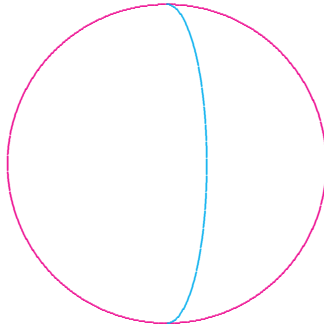
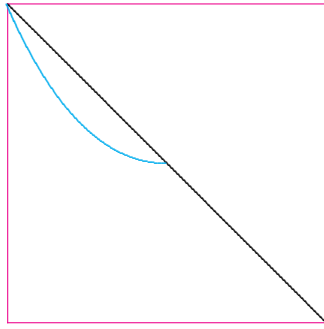
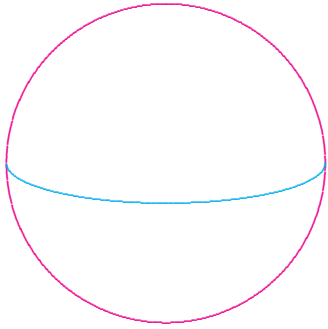
Strategy: construct 3-manifold $M = M_\psi$
as moduli space of holomorphic disks D
in $Z = \mathbb{CP}_1 \times \mathbb{CP}_1$ with ∂D on $P \subset Z$.



When ψ is the antipodal map,
disks are explicitly given by

$$\zeta \longmapsto ([a\zeta + b : c\zeta + d], [-\bar{d}\zeta - \bar{c} : \bar{b}\zeta + \bar{a}])$$

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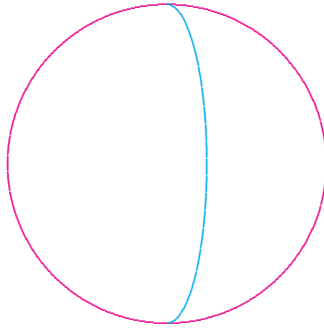
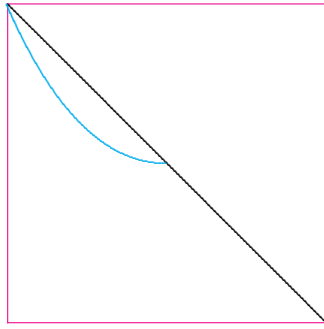
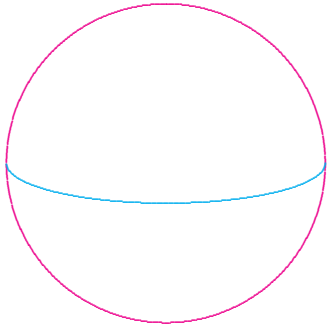
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Moduli space M of disks mod reparameterization:
de Sitter space $SL(2, \mathbb{C})/SL(2, \mathbb{R})$.

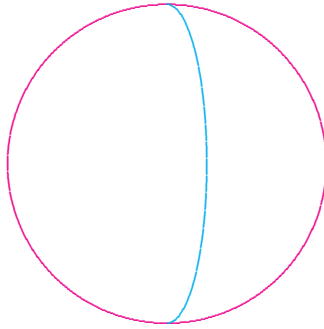
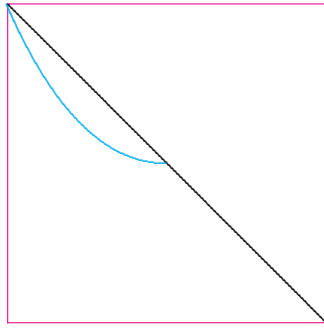
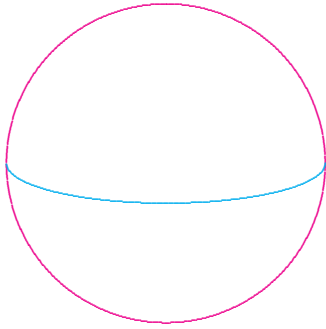


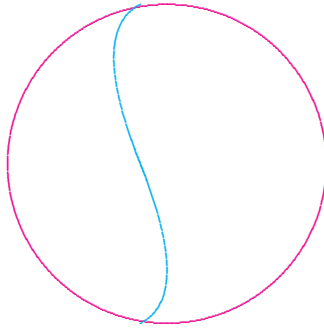
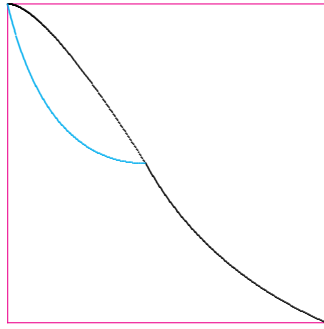
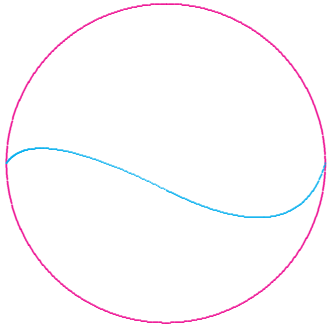
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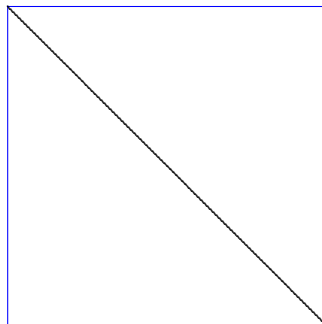
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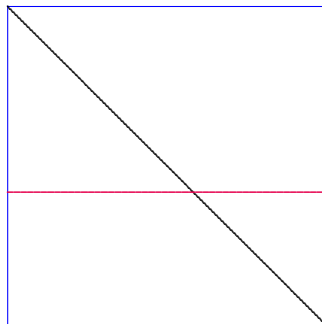
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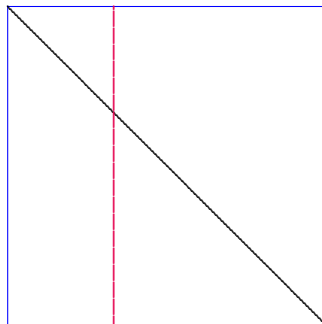
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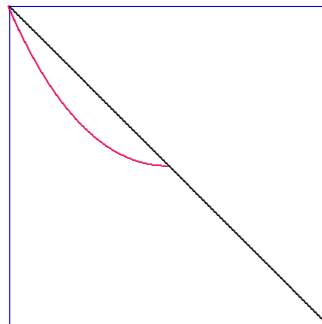
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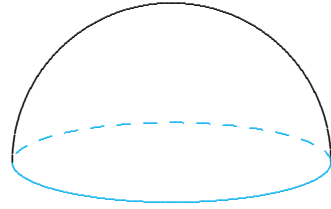
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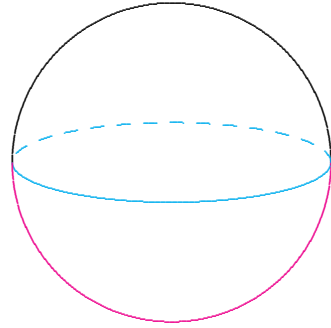
If $(\Sigma, \partial\Sigma) \rightarrow (Z, P)$ is any holomorphic curve with boundary representing \mathbf{a} , then Σ is either a holomorphic disk as above, or is a factor $\mathbb{C}\mathbb{P}_1$ of $Z = \mathbb{C}\mathbb{P}_1 \times \mathbb{C}\mathbb{P}_1$.

Proof. Consider abstract double $D \cup_{\partial D} \overline{D}$.

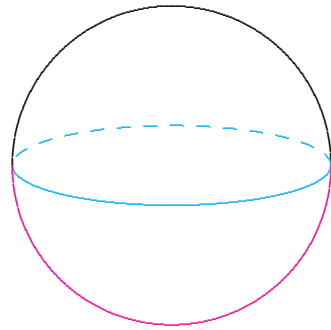
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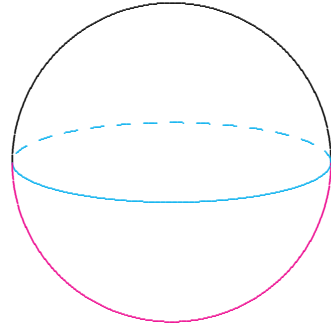
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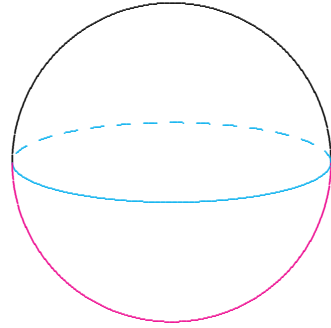
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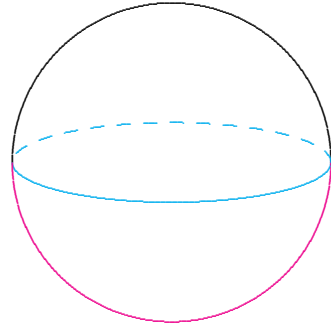
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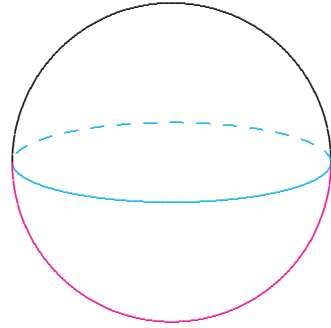
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So $\deg \Phi = 1$.

Proof. Consider abstract double $D \cup_{\partial D} \overline{D}$.



and continuous map

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Φ orientation-preserving; \implies homeomorphism.

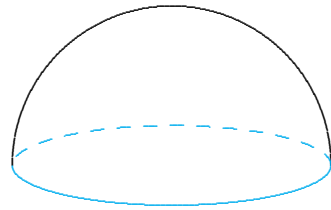
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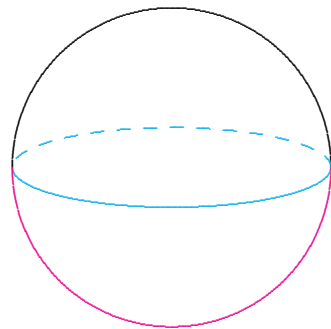
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 E & = & N & \cup_{\nu} & \overline{N} \\
 \downarrow & & \downarrow & & \downarrow \\
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Equals 2 in our case:

$$E \cong \mathcal{O}(2).$$

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$$\begin{aligned}h^1(\mathbb{C}\mathbb{P}_1, \mathcal{O}(2)) &= 0 \\h^0(\mathbb{C}\mathbb{P}_1, \mathcal{O}(2)) &= 3\end{aligned}$$

cf. Kodaira's Theorem
on deformation of complex submanifolds

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Disks all have same ω -area. \implies
any sequence has convergent subsequence...

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Tricky point: disks can degenerate to factor $\mathbb{C}P_1$.

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Since ψ is continuous deformation of antipodal,

Continuity method \Rightarrow each level set **non-empty!**

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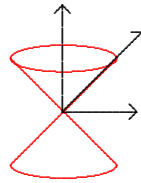
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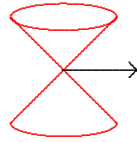
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\Rightarrow up to homothety $T_D M$ carries Lorentz metric, modelled on Killing form of $\mathfrak{sl}(2, \mathbb{R})$.

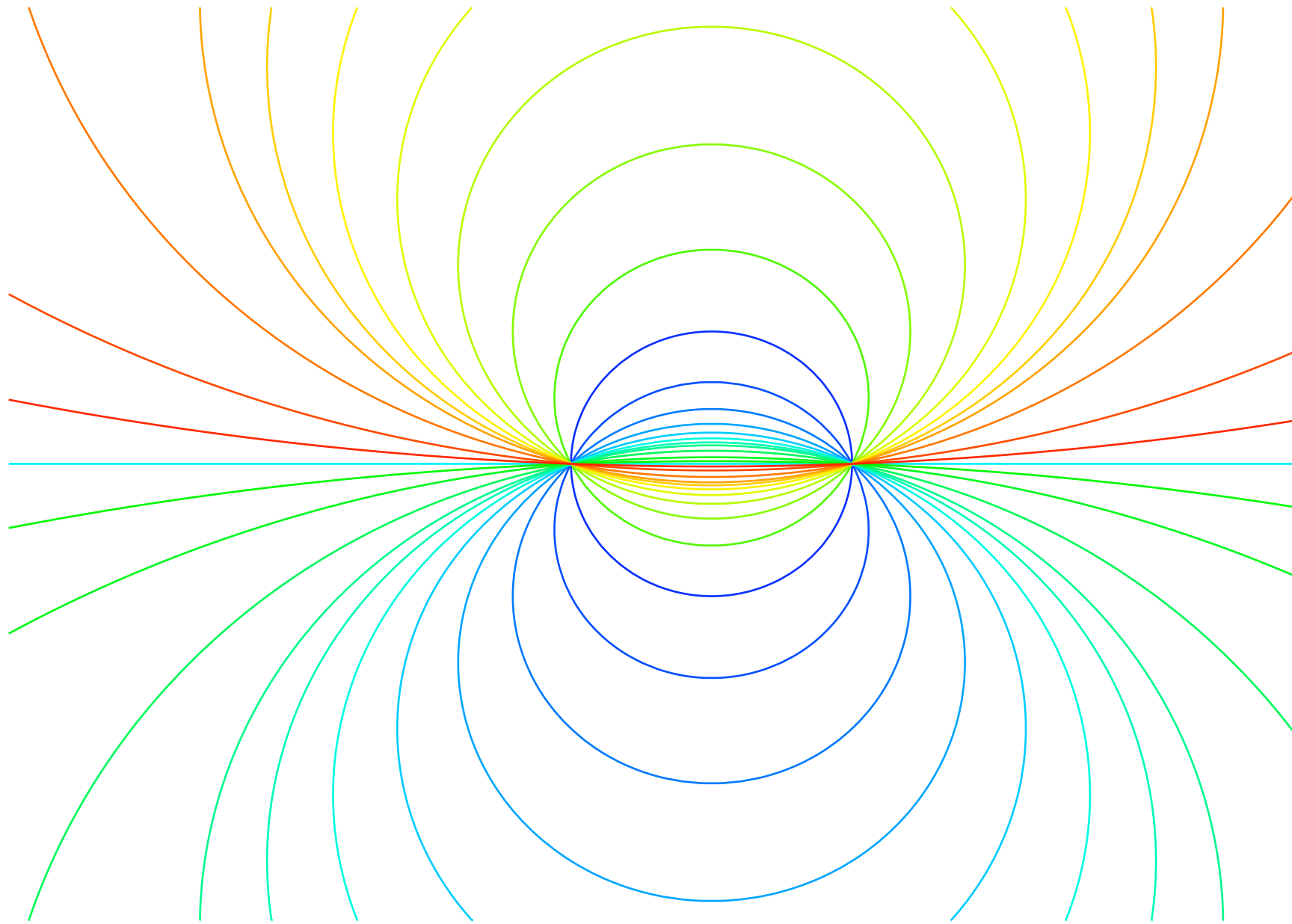
Trichotomy:

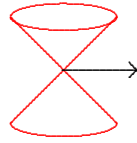
TM	$\mathfrak{sl}(2, \mathbb{R})$
space-like	hyperbolic
null	parabolic
time-like	elliptic



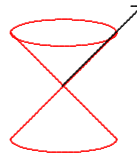


Space-like vector = infinitesimal variation with
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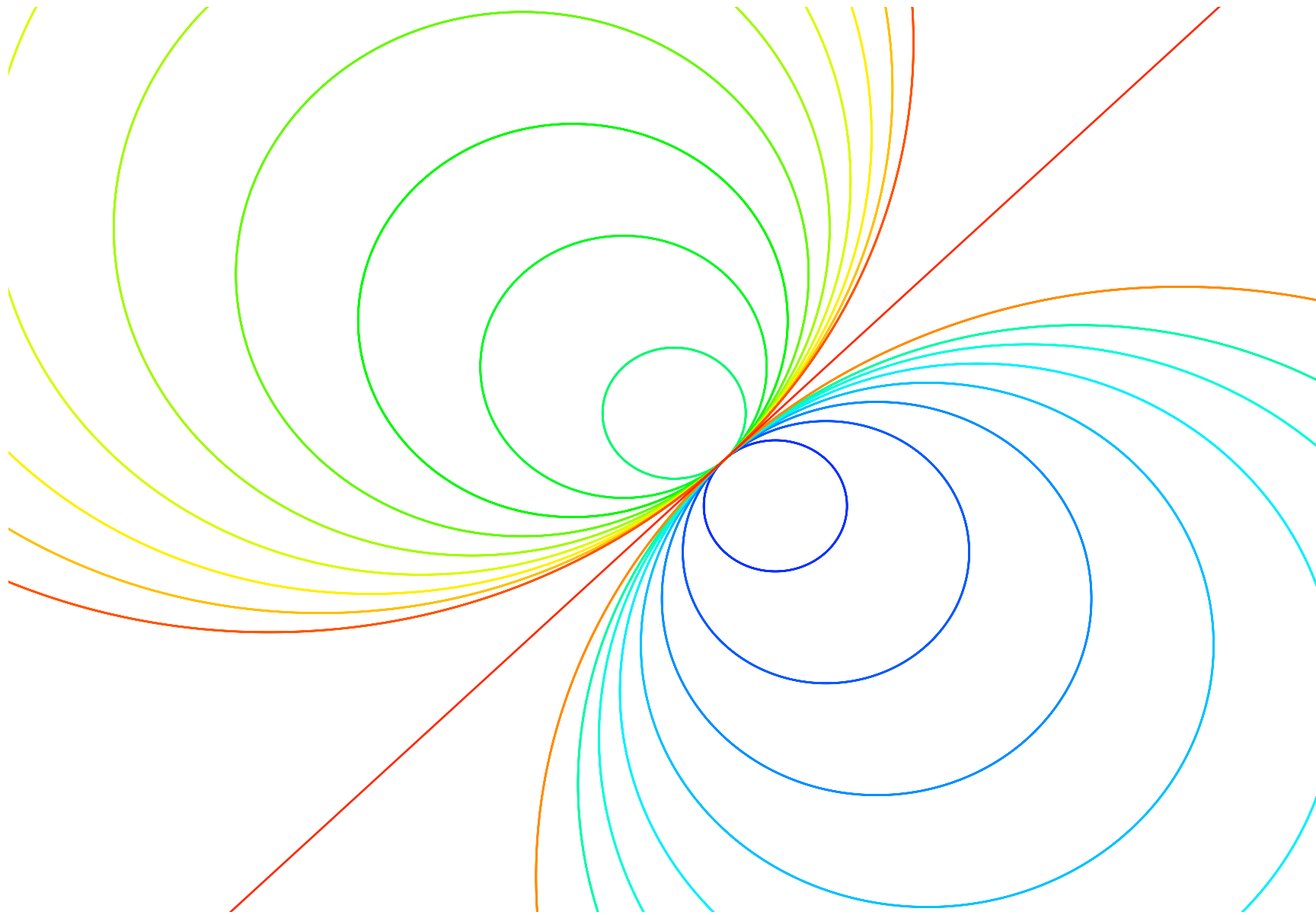


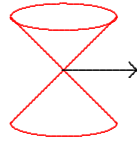


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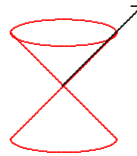


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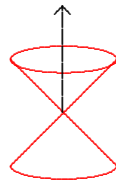




Space-like vector = infinitesimal variation with two distinct zeroes on ∂D .



Null vector = infinitesimal variation with a repeated zero on ∂D .



Time-like vector = infinitesimal variation with a single zero in interior of D : none along ∂D .

Hence area function \mathcal{A} has derivative $\neq 0$
in any time-like direction.

\mathcal{A} is time function on $(M_\psi, [g]_\psi)$!

\therefore No critical points.

\therefore Gromov-compactness \Rightarrow for all $c \in (0, 4\pi)$,
level set $\mathcal{A}^{-1}(c)$ are compact surfaces

Every endless time-like curve goes from
 $\mathcal{A} = 0$ to $\mathcal{A} = 4\pi$.

So $\mathcal{A}^{-1}(c)$ is Cauchy surface.

M globally hyperbolic!

By deformation: Cauchy surface topologically S^2 .

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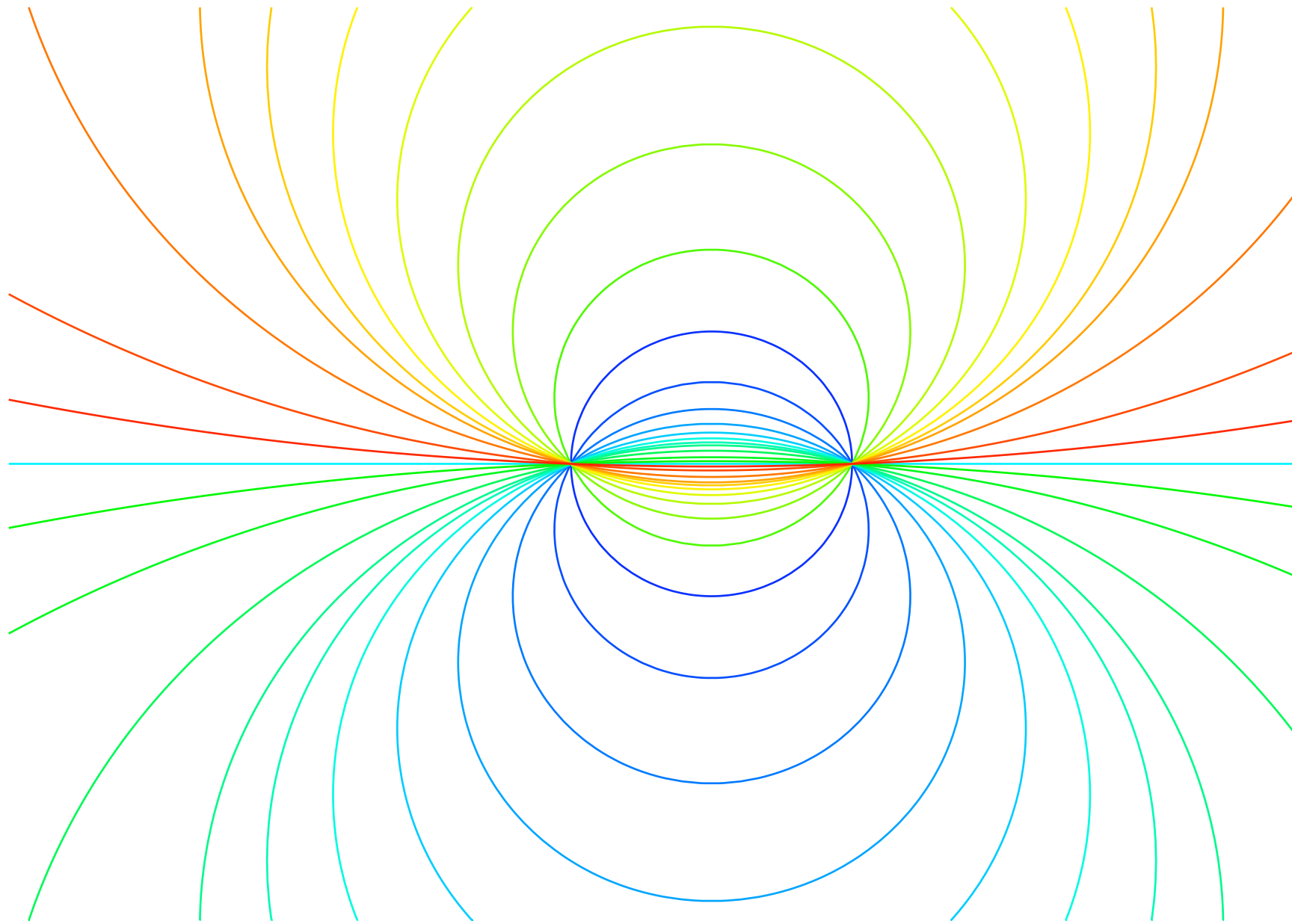
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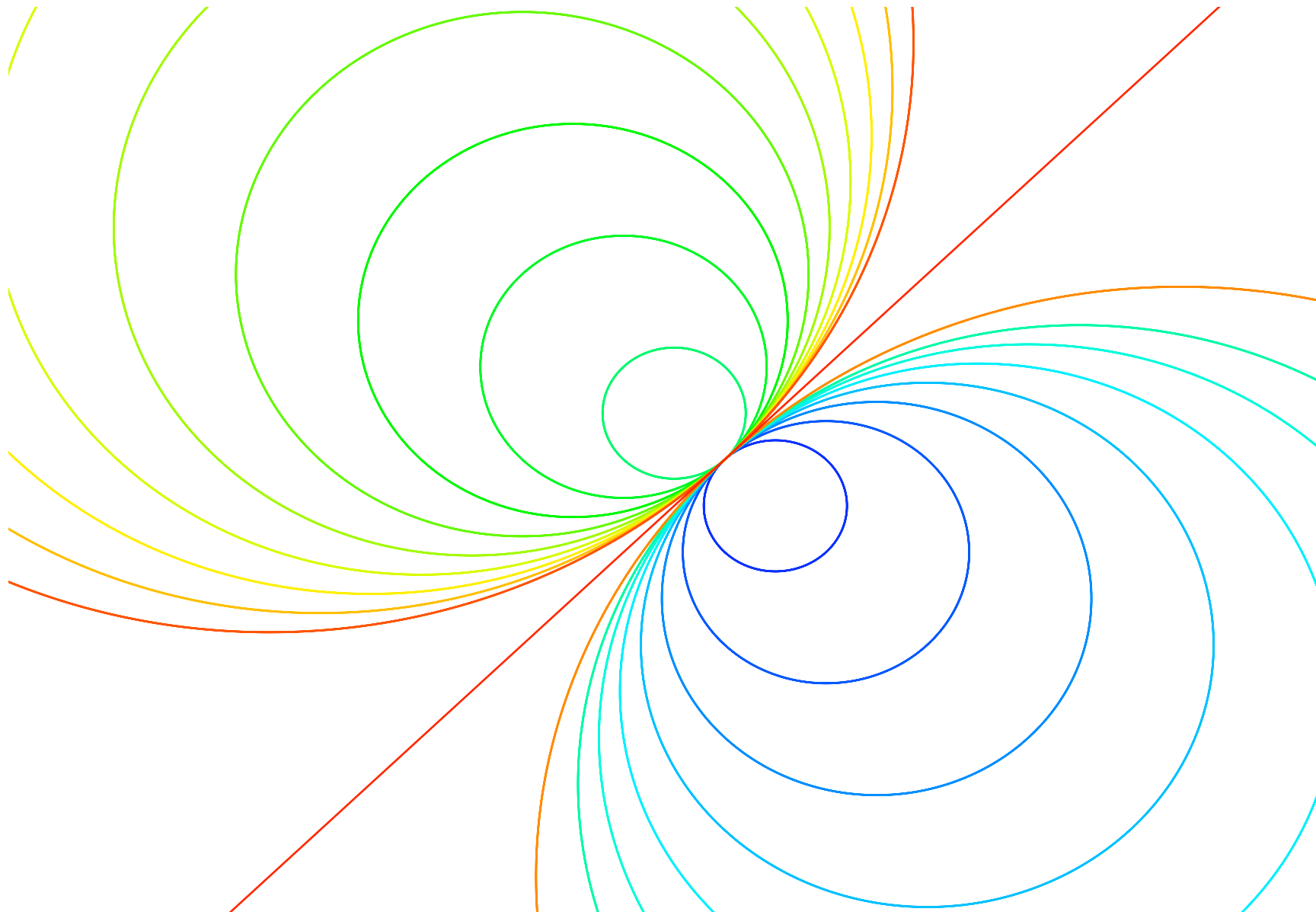
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Also gives direct proof of conformal compactness.